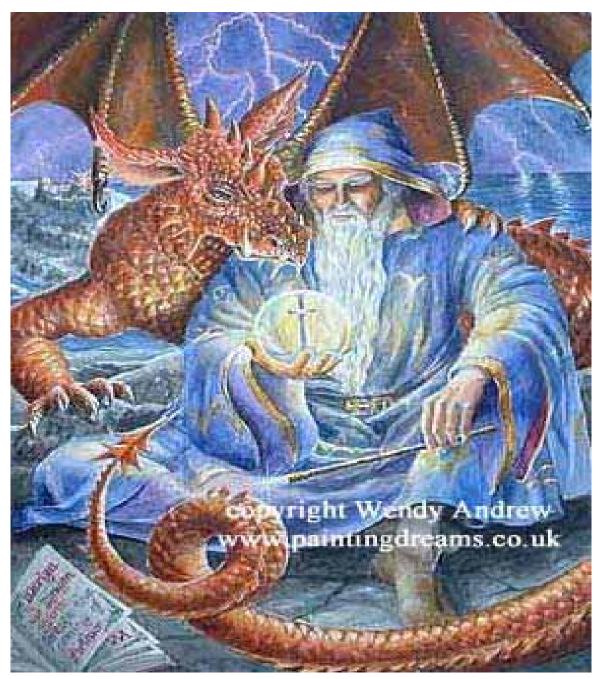
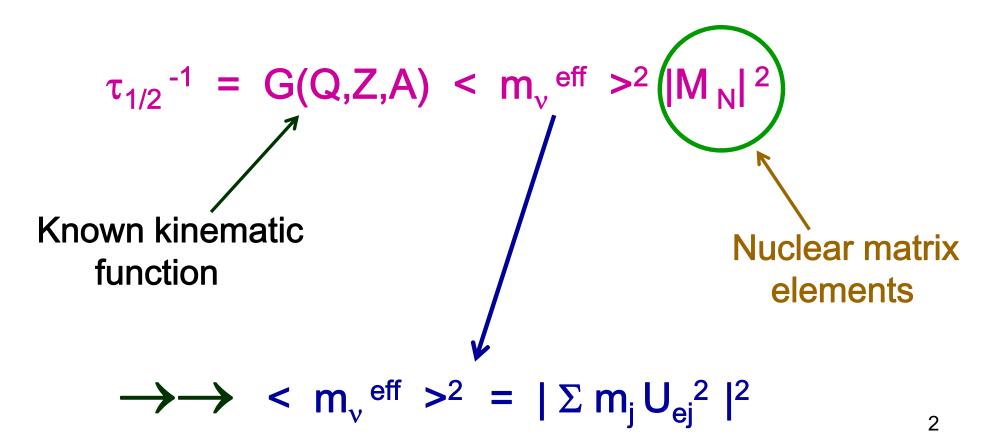
Matrix Elements for $\beta\beta_{0\nu}$



T. Rizzo

10/06/09

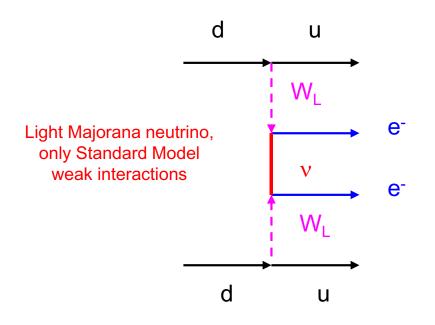
The size of the nuclear matrix element is the largest theoretical uncertainty in the extraction of effective Majorana neutrino masses from the (??) observation of neutrinoless double beta decay assuming light neutrinos. However, it appears that the size of the nuclear community working on this problem is rather small.

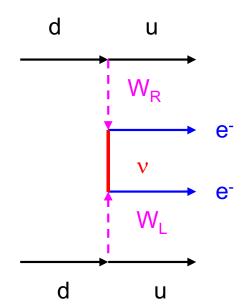


We Must Remember

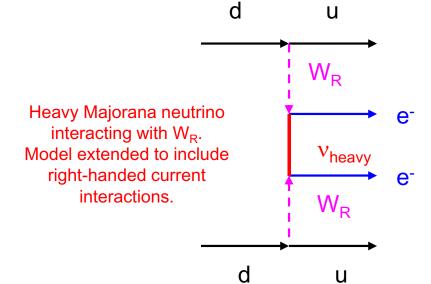
that although the observation of $\beta\beta_{0\nu}$ is certainly a clear signal for $\Delta L=2$ interactions (this is a theorem by Schechter & Valle, '82), it is not always true that the rate for the $\beta\beta_{0\nu}$ process is directly related to the Majorana neutrino mass values in any simple way as many other virtual processes can also contribute to these reactions:

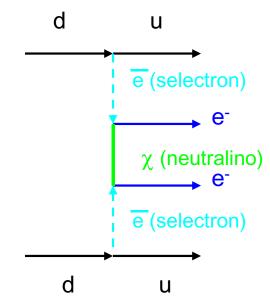
All these diagrams can in principle contribute to the decay amplitude





Light or heavy Majorana neutrino. Model extended to include right-handed W_R. Mixing extended between the left and right-handed neutrinos.



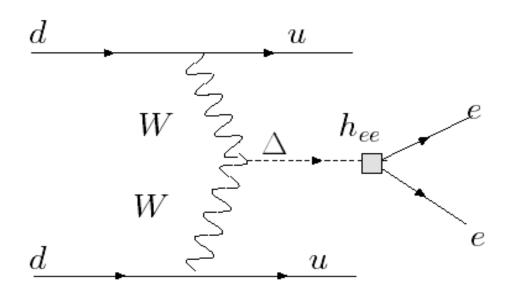


Supersymmetry with R-parity violation. Many new particles invoked. Light Majorana neutrinos exist also.

Vogel

Another Possibility: Left-Right Symmetric Model (LRSM)

$$\mathcal{L}_{\delta_{L,R}^{\pm\pm}} = \frac{g}{2} \left[\delta_{L,R}^{++} \, \overline{l^c} (h_{L,R} \, P_{L,R}) \, l + \delta_{L,R}^{--} \, \overline{l} \left(h_{L,R}^{\dagger} \, P_{R,L} \right) l^c \right]$$



The model includes a doubly charged Higgs that couples to leptons as shown

This is an example of $0\nu\beta\beta$ decay mediated by this coupling. The amplitude scales like

$$\frac{g_2^3 \ h_{ee}}{M_{W_R}^3 M_{\Delta}^2}$$

Another example is the exchange of heavy right-handed v_R and two W_R that scales like a_0^4

$$rac{g_2^4}{M_{W_R}^4 M_{
u_R}}$$

In both cases the amplitude scales like $1/\Lambda^5$ with $\Lambda \sim M_{W(R)} \sim M_{\Delta} \sim M_{v(R)}$

In the amplitudes for these other processes the direct relationship of the observed rate to the v Majorana masses is lost. The nuclear matrix elements in these 'unusual' cases are even less well studied & well known than in the more familiar case of light neutrino exchange.

Measurements of $\beta\beta_{0\nu}$ in several nuclei (as well as of the electron angular correlations*) may be necessary to pin down the 'correct 'mechanism & help with the nuclear uncertainties.

I will, however, assume that the direct 'light Majorana neutrino' assumption holds in the analysis that follows.

An Aside:

The angular correlation, i.e., the distribution of the angle between the two outgoing leptons, is given by

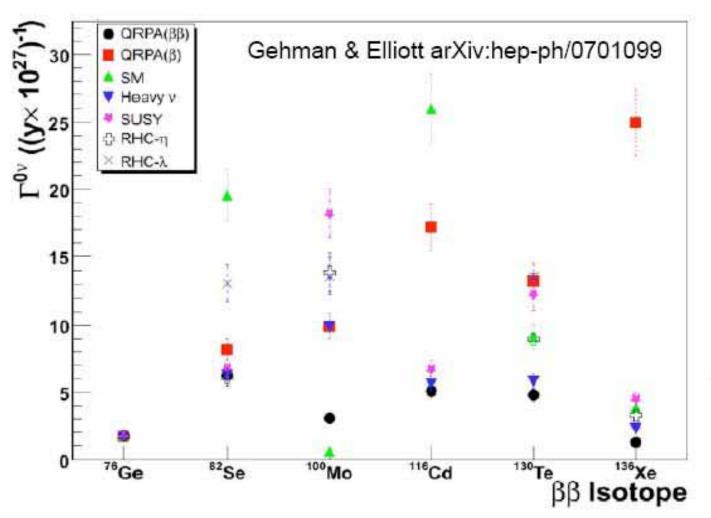
 $d\Gamma \sim 1 - K \cos \phi$ with $K \sim <\beta_1 \beta_2>$ the product of their speeds in the 'light neutrino' scenario

One finds that K~ 0.84 ±0.03 for the various nuclei depending upon A, Z & the Q values...a fairly narrow range.

For more exotic physics models, K can lie ANYWHERE in the range between +1 and -1 so measuring the K value above is reasonably restrictive on model building.

$0\nu\beta\beta$ -decay as a Probe of LNV Interactions

If $0\nu\beta\beta$ is observed, then measurements on 3-4 multiple isotopes might be able to distinguish potential physics mechanisms

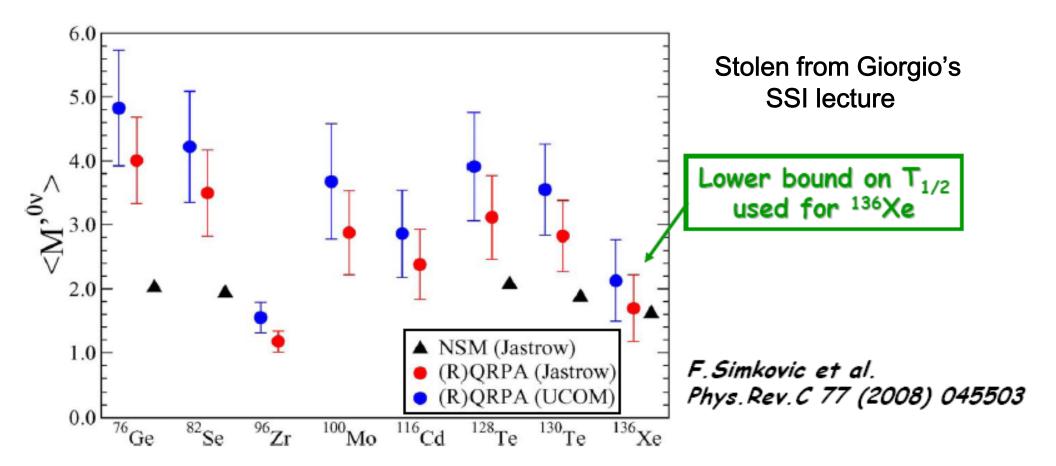


Comparison assumes a single dominant mechanism.

Requires results from 3-4 isotopes & calculation of NME to ~20%

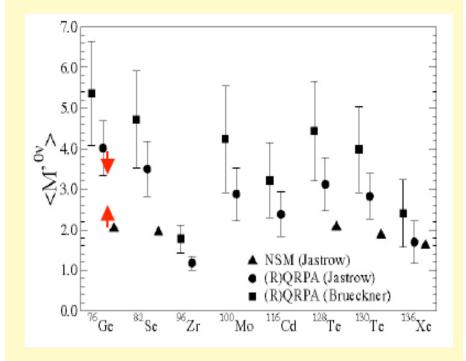
Also see Deppisc & Päs arXi:hep-ph/0612165

Wilkerson



What are these matrix elements? What techniques are being used to obtain them? What are the origins of these sizeable uncertainties? I will start from essentially complete ignorance (which accurately reflects my knowledge!) & then outline what is going on here from the point of view of a HE theorist .. It's not possible to do any of the calculations in real time.

E.g., Barea and Iachello, PRC 79, 044301 (2009)



E.g., Simkovic et al, arXiv:0902.0331

Update from Lisi @MEDEX09

The spread among NME from different nuclear models (QRPA, SM, IBM) is an indicator of **unknown** theoretical systematics. However, within each given model, the two (currently) largest source of correlated uncertainties are actually **known**:

- (1) s.r.c. effects (Jastrow? UCOM? Others?)
- (2) g_A variations (Quenched? Unquenched?)

Quenching issues are particularly embarassing... Let me be provocative:

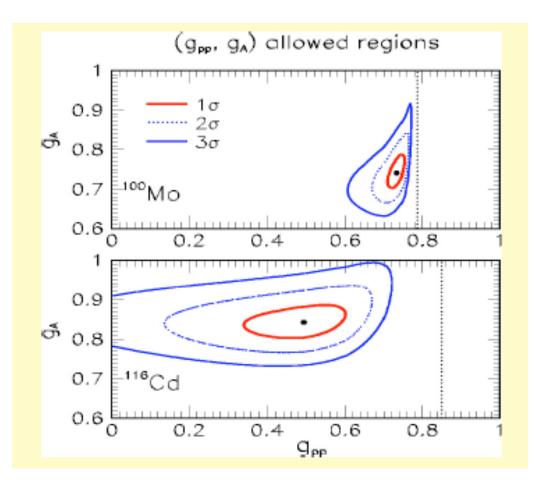
After decades of research on weak interactions in nuclei, it should be time to abandon crude recipes like "take either $g_A \sim 1$ or ~ 1.25 , and evaluate the difference." More properly, one should get a set of g_A expt/theo estimates (with errors), which might well be somewhat different in different nuclei (since g_A is an effective coupling in nuclear medium).

CONCLUSIONS:

In general, progress in this area will require <u>keeping</u> <u>track</u> of all known sources of uncertainties in each nuclear physics model, <u>computing their correlated</u> <u>effects</u> in both standard and nonstandard particle physics scenarios, & <u>constrain them</u> by using as many weak interact. and nuclear structure data as possible

Lisi advocates that, within each NME model, one should do a 'simultaneous fit' to over-constrain the free parameters using all of the existing data & extract their values...for some reason this has not been an accepted approach within the nuclear community. Doing this within each model framework has improved, & will continue to improve, the agreement among the

NME predictions for $\beta\beta_{0\nu}$.



A toy analysis for QRPA picks out parameter values not commonly considered by most authors !!!

CLEAR PROGRESS!!!

No Comment...

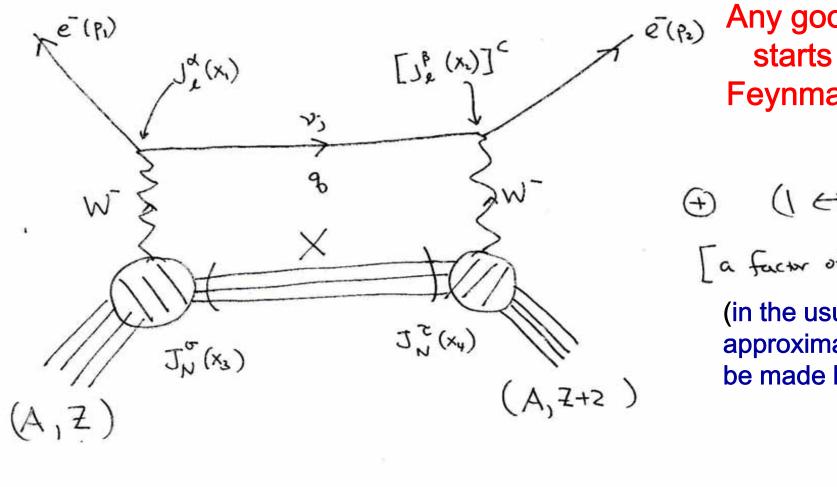
The Matrix Element of What??

The calculations of, and those leading to, the nuclear matrix elements are the result of a long series of approximations. ..

Some of these are clearly valid while others are 'reasonable' & allow for the further explicit calculation. Many approximations are associated with the ratios of the various energy scales in the problem.

Others either seem questionable or, at least to me, unclear without further study.

In all theses calculations the nucleus is considered to be a system of nucleons existing in some mean 'background' field. Pairs of nucleons may also experience a residual interaction.



Any good theorist starts with a Feynman diagram

This is an excellent approximation!

· Treat the external e as a plane wave + apply the Coulomb corrections

Standard approach

later as is done in ordinary β-decay process

Majorana

.
$$\int_{R}^{\infty} (x_{i}) = \overline{e}(x_{i}) \delta \alpha \left\{ \bigcup_{i=1}^{L} P_{L} + \overline{\xi} \bigcup_{i=1}^{R} P_{R} \right\} y_{j}(x_{i})$$
 $= 0 \text{ in the Sm}$

=> In sert into the 'leptonic' part of the diagram - it looks like

$$\sim \int \frac{d^4q}{(2\pi)^4} e^{-ig(x_1-x_1)} = \frac{1}{2} e^{-ig(x_1-x_1)} = \frac{1}{2}$$

Note the extra terms here..

Let's look a bit closer:

[m; {(Ve))2PL+32(Ve)2PR}+38 Ve) Ve) for majorana v's Ue = Ue, RH current SM piece piece GIM (ignore) Suppressed Majorana (Ignore) v masses

· What about the XXXp structure?

hat about the
$$8 \times 8 p$$
 shocking hat about the $8 \times 8 p$ shocking $\left[8 \times 8 p + 8 p$

13/17/050

do integral after we know tensor structure of nuclear operator"

What about the hadronia side ?

$$\langle 5|J_{N}^{h}(x_{1})J_{N}^{v}(x_{2})|i\rangle = \sum_{n} \langle 5|J^{h}(\vec{x}_{1})|n\rangle\langle n|J^{v}(\vec{x}_{2})|i\rangle e \qquad e$$

. Take the time components of the elections wave fuctions & the result of the go-integral above - put it all together + integrate our dt, dt2 adding the (1602) electron interchansed graph

$$\Rightarrow (2\pi) \delta (E_{5} + E_{e_{1}} + E_{e_{2}} - E_{i}) \sum_{n} \left[\frac{\langle 5|J_{\mu}^{h}(\vec{x}_{i})|n\rangle\langle n|J_{\mu}(\vec{x}_{i})|i\rangle}{g_{0}^{*}(E_{n} + g_{0} + E_{e_{2}} - E_{i})} + (16)^{2} \right]$$

· now go = (171 + m)) 1/2 = 171 = 8 ~ / Ave nucleon spacing = 100 MeV >> En En as some aurage value E and replace So we approximate I Injent by [1] This is called the closure approximation + is.

10-15% ??

good to ~15%. to go further we need to Know [J].

$$\int N \sim \overline{\Psi} \quad \tau^{+} \left\{ g_{V}(g_{S}^{2}) 8^{h} - g_{A}(g_{S}^{2}) 8^{h} 8^{s} - i g_{A}(g_{S}^{2}) \frac{\sigma^{h v} g_{V}}{2Mp} + g_{P}(g_{S}^{2}) g_{S}^{h} 8^{s} \right\} \overline{\Psi}$$

$$(g_{O} \text{ to Ch. 10} \quad \int \quad \int \quad \text{Nucleon} \quad \text{Nucleon} \quad \text{Secondo scalar} \quad \text{Wave function} \quad \text{of } B_{J} + D) \quad \text{Isospin} \quad \text{Vector} \quad \text{axial-vector} \quad \text{magnetic} \quad \text{FF} \quad$$

is the most general nucleon current

Non - Relativistic

Approximation.

Ekm << M

$$\vec{J}_{n}(g^{2}) = g_{n}(g^{2}) \vec{\sigma}_{n} + i g_{n}(g^{2}) \frac{\vec{\sigma}_{n} \times \vec{g}}{2m_{p}} - g_{p}(g^{2}) \frac{\vec{g}(\vec{\sigma}_{n} \cdot \vec{g})}{2m_{p}}$$

(1 hucleun magnetic moments

Form Factor
$$g_{V}(g_{0}^{2}) = g_{V} \left\{ 1 + g_{V/2}^{2} \right\}^{-2}$$
, $g_{A}(g_{0}^{2}) = g_{A} \left\{ 1 + g_{V/2}^{2} \right\}^{-2}$

Scales

$$= 1 \quad 20.71 \text{ GeV}^{2}$$

$$= 1.25.7?$$
(may not be true in ~20-40% a large nucleus) quenching?

$$J_{V}^{\mu} \otimes J_{V}^{\mu} \sim \Omega = \tau_{+} \tau_{+} \left\{ \left(h_{F} + h_{GT} \vec{\sigma}_{1} \cdot \vec{\sigma}_{2} \right) - h_{+} S_{12} \right\} \left(S_{12} = 3 \vec{\sigma}_{1} \cdot \hat{g} \vec{\sigma}_{2} \cdot \hat{g} - \vec{\sigma}_{1} \cdot \vec{\sigma}_{2} \right)$$

Fermi Gamow Tensor ~ $3 co^{2} - 1$

$$h_{F} = g_{V}(g_{0}^{2})$$

$$h_{GT} = g_{A}^{2}(g_{0}^{2}) \left[1 - \frac{2}{3} \frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} + \frac{1}{3} \left(\frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} \right)^{2} \right] + \frac{2}{3} \frac{g_{A}^{2}(g_{0}^{2})}{g_{A}^{2}}$$

$$h_{T} = g_{A}^{2}(g_{0}^{2}) \left[\frac{2}{3} \frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} - \frac{1}{3} \left(\frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} \right)^{2} \right] + \frac{1}{3} \frac{g_{A}^{2}(g_{0}^{2})}{g_{A}^{2}}$$

$$h_{T} = g_{A}^{2}(g_{0}^{2}) \left[\frac{2}{3} \frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} - \frac{1}{3} \left(\frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} \right)^{2} \right] + \frac{1}{3} \frac{g_{A}^{2}(g_{0}^{2})}{g_{A}^{2}}$$

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$$h_{T} = g_{A}^{2}(g_{0}^{2}) \left[\frac{2}{3} \frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} - \frac{1}{3} \left(\frac{g_{A}^{2}}{g_{A}^{2} + m_{A}^{2}} \right)^{2} \right] + \frac{1}{3} \frac{g_{A}^{2}(g_{0}^{2})}{g_{A}^{2}}$$

$$\int_{0}^{1} e^{igg} g_{A}^{2} g_{A}^{2$$

& put it all together to get...

Assuming only of -> of transitions so that the outgoing e's are in the S-wave state + so that e (Pir, +P2·r2) ~ 1, we get

(Including the p-waves & Coulomb corrections is straightforward)

$$M_{ov} = \frac{2R}{\pi g_{\lambda}^{2}} \left\langle f \middle| \int_{0}^{\infty} dg \sum_{a,b} \frac{J_{o}(g_{rab}) \left[-h_{F} + h_{GF} \vec{\sigma}_{a} \cdot \vec{\sigma}_{b} \right] + J_{z}(g_{rab}) h_{f} \cdot \hat{S}_{ab} \right]}{g_{+} \cdot \bar{E} - \left(E_{i} + E_{f} \right) / 2}$$

where R is the nuclear radius [a source of 'errors' as various values used!]

Some authors drop the tensor term [230% effect!) to simplify further :

Mov = MGT - gv HF

MF = < 51 [H TaTbli>, MGT = <51 [H Ga. Gb ZaTbli>

With $H \simeq \frac{2R}{\pi r_{ab}} \int_{0}^{\infty} dq \frac{\sin(qr_{ab})}{q + \overline{\epsilon} - (\overline{\epsilon}_i + \overline{\epsilon}_e)/2}$ is the mention potential

Comments

- As you can see, now that we have arrived at the form of the relevant operators, a reasonably large number of approximations have already been made. These are over & above any FURTHER approximations that are made in obtaining the values of the matrix elements themselves but most of these are not necessary to make.
- So far this has all been rather straightforward (??) although a bit messy. HOWEVER, this is where High Energy Physics ends & Nuclear Physics takes over --- which causes an immediate drastic increase in the murkiness of the nuclear matrix element calculations...

Comments (cont.)

One apparently obvious & universal uncertainty in these matrix element calculations is the detailed knowledge required of both the initial & final state nuclear wave functions (as well as those for any of the possible intermediate states!). Remember that these are obtained from some assumed forms for the 'collective' nuclear potential and the N-N interaction which are not known a priori.

Matrix Element Evaluation

The matrix elements themselves are evaluated using several *FAMILIES* of techniques which have different strengths and weaknesses & are based on different physics assumptions:

- QRPA = Quasiparticle Random Phase Approximation, which provides a 'straightforward' (!) calculational technique but allows only limited testing opportunities..most commonly used approach
- NSM = Nuclear Shell Model, which is easily tested by nuclear spectroscopy, etc., but is difficult to apply to heavy nuclei
- IBM = Interacting Boson Model, a modified shell-like model
- Within EACH family are further competing approximations & 23 techniques apparently based on the bias of the different authors

Aside II:

Constraints on $\beta\beta_{0\nu}$ nuclear matrix elements come from many sources - but one potentially important one is the corresponding $\beta\beta_{2\nu}$ process which is lepton number conserving & for which data exists. For the analogous $0^+ \rightarrow 0^+$ transition the rate for an even-even nucleus is given by

$$\tau_{1/2}^{-1} = F(Q,Z,A) |M_{GT}|^2$$

where F is a known function and the nuclear matrix element is essentially pure GT due to selection rules:

$$M_{GT}^{2\nu} = \sum_{m} \frac{\langle f||\sigma\tau_{+}||m\rangle\langle m||\sigma\tau_{+}||i\rangle}{E_{m} - (M_{i} + M_{f})/2}$$

Here |m> label the 1⁺ of the intermediate odd-odd nucleus. Note Fermi matrix elements are NOT constrained.

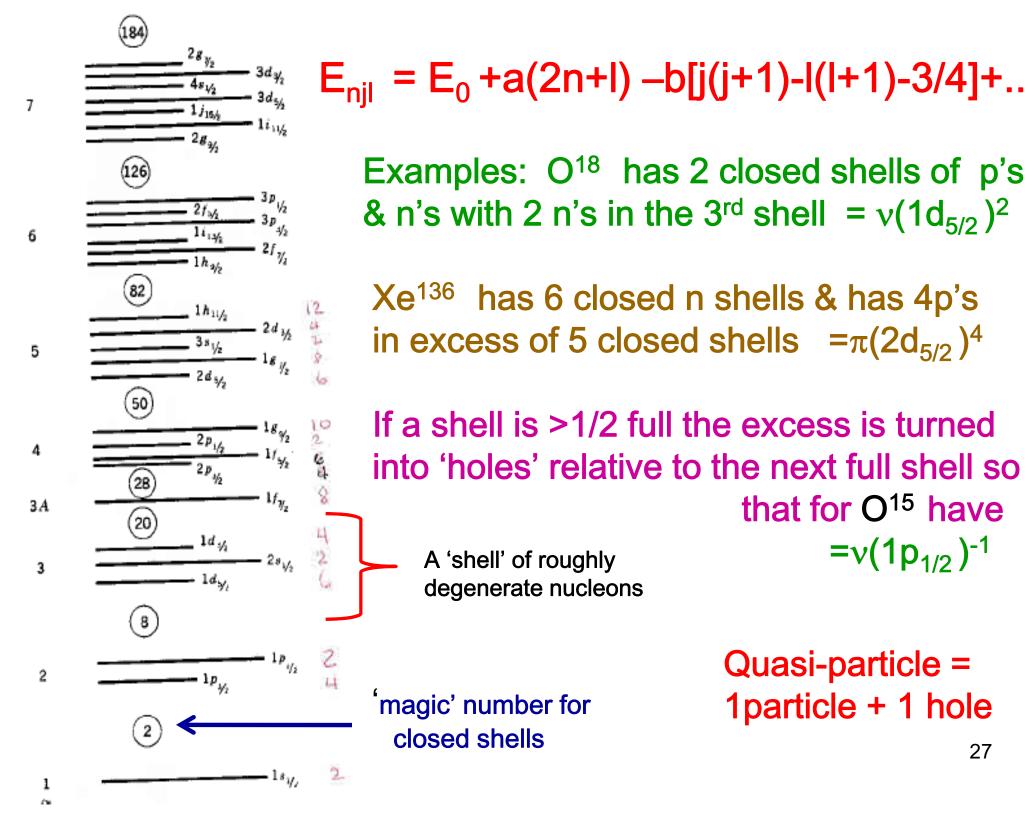
NSM

- At it's most basic level the NSM is the easiest to understand as
 it is very similar to atomic physics. SUPPOSE we knew the
 complete set of wavefunctions and energy levels of the original
 & final nuclei, as well as those for any of the possible intermediate states, in the ββ_{0ν} process.
- We then could diagonalize the effective Hamiltonian & obtain the required matrix elements by numerical integration & then summing over all of (the very large number of) configurations.
- The goal of NSM (as I see it) is to 'explain' nuclear energy levels & to allow for the calculation of various transition rates in a manner following the analogy w/ atomic physics.

To accomplish this they need to (numerically!) obtain the nuclear wavefunctions based on an assumed form of the common central potential (there are several possible choices, e.g., the Woods-Saxon potential or modified oscillator potential) & correcting for, e.g., the 2-body, spin-orbit, and velocity dependent interactions using experimental data from these & related nuclei as input employed in a self-consistent manner. This is clearly an art form as well as science!

The basic concepts are used in all NME calculations

Let's review....



 $(1d_{5/2})^2$ is called a configuration of O^{18} but it is not the only possible one...

• 'Collisions' of 2 nucleons in the same shell can lead to some admixtures of shell model states. For example, in O¹⁸ the 2n's experience a pairing interaction so that the nucleus can be in a superposition of NSM basis states with the same total 2-particle total spin (here=0) and (here even) parity:

$$\Psi = a (d_{5/2})_0^2 + b (s_{1/2})_0^2 + c (d_{3/2})_0^2$$

with the relative weighting depending upon the strength of the interaction. The subscript on these terms tells us that these are all configurations with total J=0. The squares of the coefficients a-c , divided by the maximum number of possible nucleons in a given orbital are called occupation numbers , e.g.,

 $V_{5/2}^2 = |a|^2/3$ for the $d_{5/2}$ configuration above

NSM II

Several issues now arise:

- (i) Given the uncertainties on the 'input' assumptions, the NSM wavefunctions, though leading to testable predictions, are not uniquely determined. E.g., variational techniques with different trial wavefunctions can yield similar results. (The full nuclear wavefunctions are then the symmetrized products of the single nucleon wavefunctions obtained via Slater determinants)
- (ii) They will certainly not be determined for all possible configurations of the initial, final & intermediate nuclei relevant for $\beta\beta_{0\nu}$. Data does not exist for all these states .
- (iii) The wavefunctions will not necessarily be probed by the nuclear spectroscopy data in the same range of 'r₁₂' ~ few (MeV) ⁻¹ as do the 'neutrino potentials' ~(100 MeV) ⁻¹.

NSM III

• (iv) There are obviously very many configurations to sum over & this is usually truncated to something ~ (10^{10-11}) due to CPU issues especially as the number of nucleons grows, i.e., for the heavier nuclei such as those employed in $\beta\beta_{0\nu}$ experiments.

Clearly, these calculations, though constrained by data, can at best be approximate – but, as these calculational techniques are improving (e.g., the use of Monte Carlo techniques) and more & more data is confronted by the NSM calculations, they can be viewed as reasonably trustworthy & the ones of choice in cases where they are applicable similar to what may see for large atoms.

Aside III

- A general correction that is employed in both NSM & QRPA calculations is accounting for the 'overlap' of the two initial nucleons' wavefunctions due to the short ranged correlations from the neutrino potential, e.g., 2 nucleons occupying the 'same place'.
- The Jastrow method rescales both the initial & final state wavefunctions at small separations by a brute-force ad-hoc common correction function which suppresses these overlaps. Unfortunately, this approach, though commonly used, does not conserve the norms of these states. This is ugly...
- The UCOM approach does something similar but through the use of a unitary transformation matrix which is 'softer' as well as norm-preserving & is thus a superior approach theoretically & is more favored by recent analyses & data fits.

QRPA

- The goal of QRPA is to reduce the complexities of the NSM calculations especially for heavier nuclei. To me, while it may reduce apparent calculational complexity, it's theoretical structure seems forbiddingly difficult to understand for an amateur.
- Similar to NSM the individual oscillator-like wavefunctions arise from the mean field core plus residual interactions obtained by summing over all of the N-N pair interactions & fit to the available data from the relevant nuclei. Many different N-N potentials are in use.
- QRPA comes in MANY, MANY different versions: RQRPA,
 Full-QRPA, SCQRPA, HQRPA, and on & on

QRPA II

• But QRPA goes further through the pairing of nuclei using BCS-like techniques & by then performing a double multipole expansion of the relevant nuclear matrix elements in terms of the J_I of the intermediate states which are restricted to be within ~10-20 MeV of either ground state & in terms of the relative J_R of the annihilating neutrons and final state protons. This makes the notation appear very 'cumbersome' to say the least!

 An additional interaction is included in QRPA between particles & holes which in principle differs in strength (g_{ph}) from the usual particle –particle interaction (see below). The value used comes data fitting.

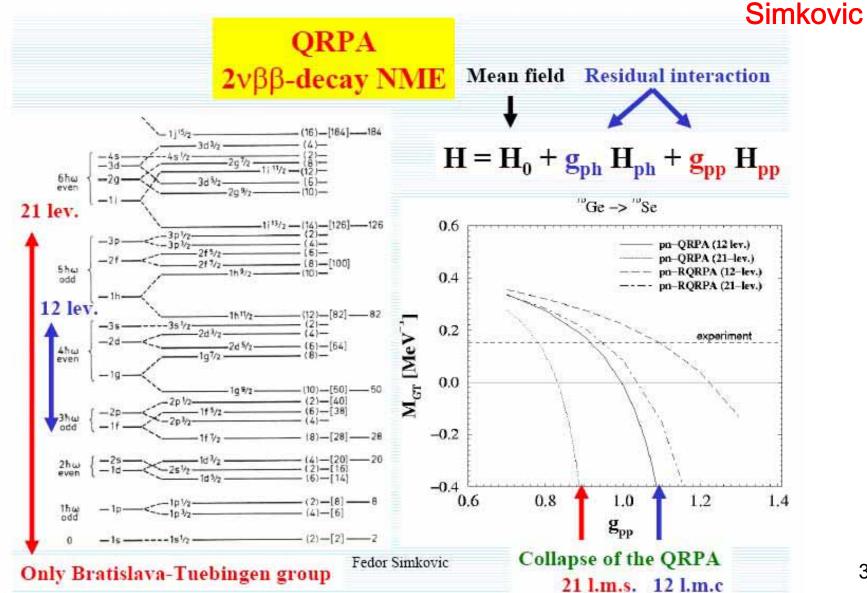
QRPA III

- The nuclei are built up of a core plus bosonic quasiparticles which are point-like bound states (i.e., obeying Bose-Einstein statistics) of nucleon particles & holes..thus, the Exclusion Principle is violated(!) The Renormalized PQRA tries to take this into account & would seem to be superior but violates sum rules. While an arbitrarily large number of possible p,n single particle states may be included the subset of possible quasiparticle states included is computationally restricted.
- The number of states included in various calculations seem to vary significantly w/ up to ~30-40% influence on matrix elements..certainly MORE is better but harder. Some authors include only two shells in their model space while others extend this to 5.

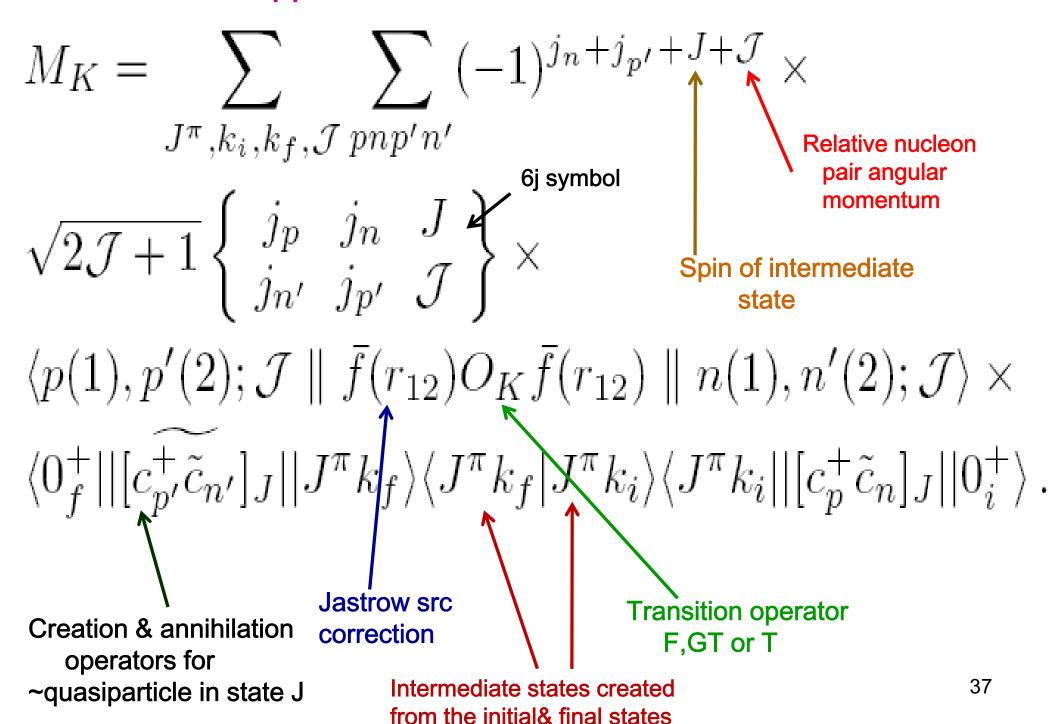
QRPA IV

- QRPA calculations are especially sensitive to deviations in the nuclear shape since multipole expansions are performed.
 Some authors have tried to accommodate deformations as a part of the QRPA formalism but these involve a far greater number of potentially contributing configurations.
- The two other free parameters in PQRA include the overall N-N coupling strength: $g_{pp} \sim 1$. This can generally be fitted by making use of the corresponding 2v decay mode. The last parameter is the axial-vector coupling g_A (~1??) with its associated quenching issue.

• If the N-N coupling becomes too strong then the calculations can 'collapse', i.e., fail in the interesting parameter region due to Exclusion Principle violations.



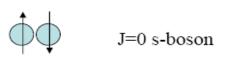
In the Closure Approximation...



IB(F)M*

is an old idea ('74) which has only been recently ('09) applied to the problem of double-beta decay. A problem w/ the NSM is that there are SO many nucleons in very many possible states when A is large. The goal of IBM, like the QRPA scheme, is to simplify this structure so that the multitude of states entering into, e.g., a decay calculation can be significantly reduced while still getting the essential physics correct.

NSM tells us that low-lying states of E-E nuclei are made of nucleon pairs which are either total spin-0 or 2. IBM postulates only valence s- or d-bosons outside of an inert core instead of p's or n's. Extra unpaired nucleons remain as fermions in more general nuclei; IBM-2 treats p's & n's pairings separately.





J=2 d-boson

lachello



Unpaired fermions

All can be calculated at once using the compact expression:

$$V_{s_{1},s_{2}}^{(\lambda)} = \frac{1}{2} \sum_{n,n'} \tau_{n}^{+} \tau_{n'}^{+} \left[\sum_{n}^{(s_{1})} \times \sum_{n'}^{(s_{2})} \right]^{(\lambda)} \cdot V(r_{nn'}) C^{(\lambda)}(\Omega_{nn'})$$

current-current operators of different L

$$\lambda = 0, s_1 = s_2 = 0(F)$$

$$\lambda = 0, s_1 = s_2 = 1(GT)$$

$$\lambda = 2$$
, $s_1 = s_2 = 1(T)$

Neutrino potentials

In second quantized form:

$$\begin{split} &V_{s_{1},s_{2}}^{(\lambda)} = -\frac{1}{4} \sum_{j_{1}j_{2}} \sum_{j'_{1}j'_{2}} \sum_{J} (-1)^{J} \sqrt{1 + (-1)^{J} \delta_{j_{1}j_{2}}} \sqrt{1 + (-1)^{J} \delta_{j'_{1}j'_{2}}} \\ &\times G_{s_{1}s_{2}}^{(\lambda)} (j_{1}j_{2}j'_{1}j'_{2};J) \bigg[\Big(\pi_{j_{1}}^{\dagger} \times \pi_{j_{2}}^{\dagger} \Big)^{(J)} \cdot \Big(\tilde{V}_{j'_{1}} \times \tilde{V}_{j'_{2}} \Big)^{(J)} \bigg] \end{split}$$

with angular momentum J

Creates a pair of protons Annihilates a pair of neutrons with angular momentum J

lachello

The fermion operator V is mapped onto the boson space by using:

$$\left(\pi_{j}^{\dagger} \times \pi_{j}^{\dagger}\right)^{(0)} \mapsto A_{\pi}(j)s_{\pi}^{\dagger} \leftarrow \left(\pi_{j}^{\dagger} \times \pi_{j'}^{\dagger}\right)_{M}^{(2)} \mapsto B_{\pi}(j,j')d_{\pi,M}^{\dagger}$$

LO mapping from nucleon to pairs of bosons.

$$\begin{split} &V_{s_{1}s_{2}}^{(\lambda)} \longmapsto -\frac{1}{2} \sum_{j_{1}} \sum_{j_{1}'} G_{s_{1}s_{2}}^{(\lambda)} \left(j_{1}j_{1}j_{1}' \ j_{1}'; 0 \right) A_{\pi}(j_{1}) A_{\nu}(j_{1}') S_{\pi}^{\dagger} \cdot \tilde{S}_{\nu} \\ &- \frac{1}{4} \sum_{j_{1}j_{2}} \sum_{j_{1}'j_{2}'} \sqrt{1 + \delta_{j_{1}j_{2}}} \sqrt{1 + \delta_{j_{1}j_{2}'}} G_{s_{1}s_{2}}^{(\lambda)}(j_{1}j_{2}j_{1}') J_{2}^{\dagger}; 2) B_{\pi}(j_{1}, j_{2}) B_{\nu}(j_{1}', j_{2}') d_{\pi}^{\dagger} \cdot \tilde{d}_{\nu} \end{split}$$

Neutron pairs annihilated Proton pairs created

Modified matrix elements

The basis

$$\left(S^{\dagger}\right)^{\frac{n-\nu}{2}}\!\left(D^{\dagger}\right)^{\!\!\frac{\nu}{2}}\!\left|0\right\rangle$$

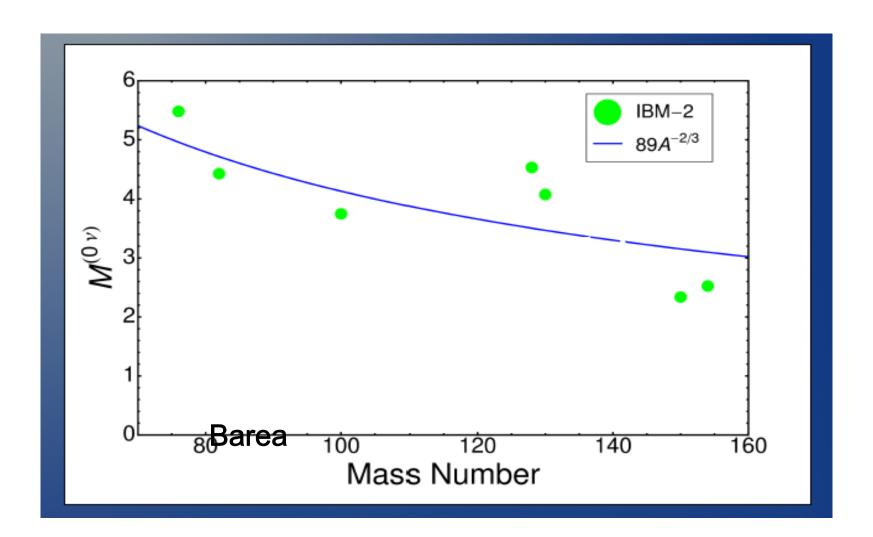
is constructed with operators:

$$\begin{split} S_{\pi}^{\dagger} &= \sum_{j} \alpha_{j} \sqrt{\frac{\Omega_{j}}{2}} \left(\pi_{j}^{\dagger} \times \pi_{j}^{\dagger} \right)^{(0)} \\ D_{\pi}^{\dagger} &= \sum_{j \leq j'} \beta_{jj'} \frac{1}{\sqrt{1 + \delta_{jj'}}} \left(\pi_{j}^{\dagger} \times \pi_{j'}^{\dagger} \right)^{(2)} \end{split}$$

now input the appropriate wavefunctions from fits to spectrum data

with $\Omega_j = j + 1/2$.

• The predictions of IBM are generally slightly larger than, but in general agreement with, the QRPA predictions at the level of the expected ~25% uncertainty.



Conclusions:

My impression is that, certainly for the case of the XENON nucleus, by a comparison of the predictions, the matrix element uncertainties are apparently slowly shrinking through both more sophisticated calculations as well as systematic analyses using better common input data.

You EXOs are Smart and/or Lucky!

My bet is that by the time data is available the uncertainty on the NME for XENON will be at the ~20% level (or less) which seems to be more than adequate for neutrino mass extraction. How good these calculations will be for other nuclei, which are clearly necessary to test the underlying theory, is unclear. Observing this rare process will, of course, produce faster theoretical results.