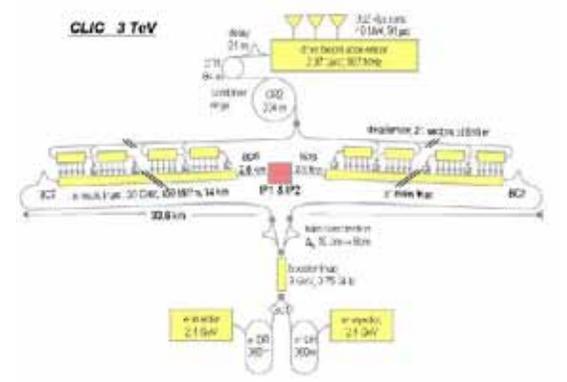
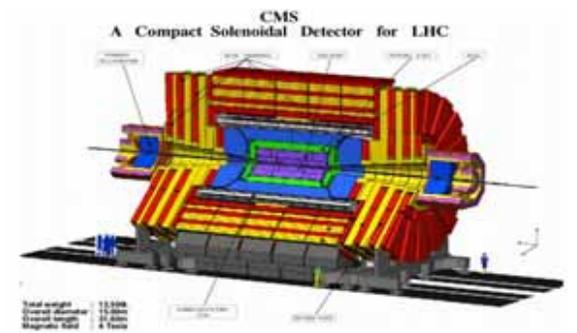
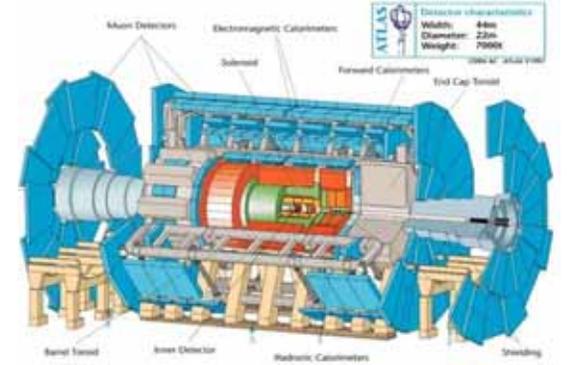
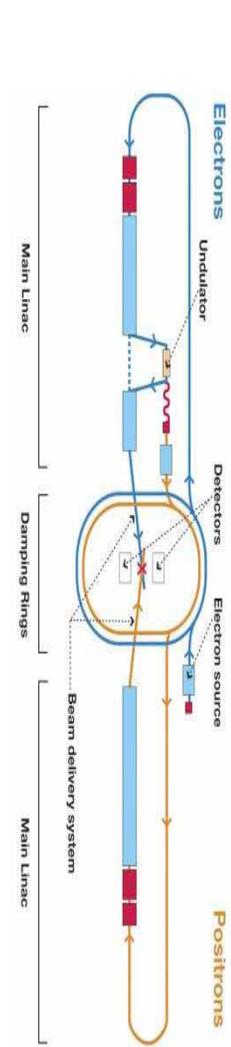
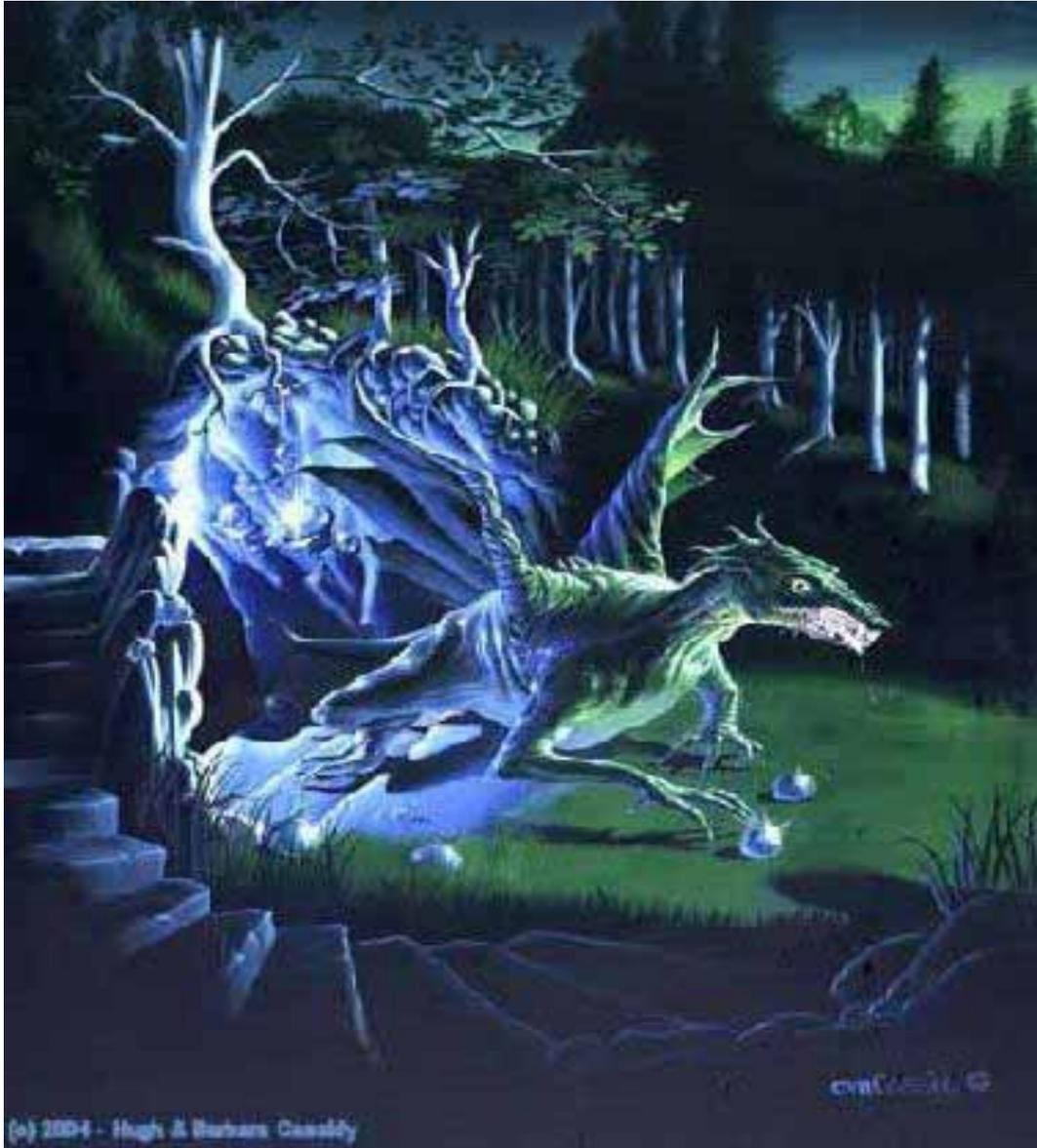
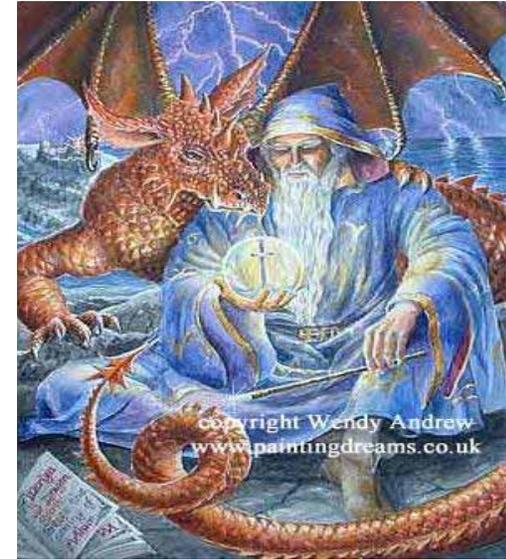


# Emerging BSM @ the 7 TeV LHC & the LC



# THE QUESTIONS:

- How is electroweak symmetry broken?
- How is the hierarchy stabilized?
- What is the origin of flavor ?
- Are there more than 4 dimensions?
- .....



→ We have a lot of questions & we all believe that the LHC will provide at least some partial answers...

BUT WHEN ??

& what will it mean for the  
LC ?



When we consider LHC implications for the LC there are three different time scales:

(i) The 'classic' case : 14 TeV & 100+ fb<sup>-1</sup> LHC....>2016??

(ii) The 'near horizon' : 7 TeV & ~5 fb<sup>-1</sup> LHC.....2012

(iii) → But both ATLAS & CMS are **already** producing important BSM results right now using the first ~35-42 pb<sup>-1</sup> of 7 TeV LHC data

(i) Has been discussed for **many years** & will not be considered here. **Instead** we will discuss (ii) & perhaps more to the point (iii)

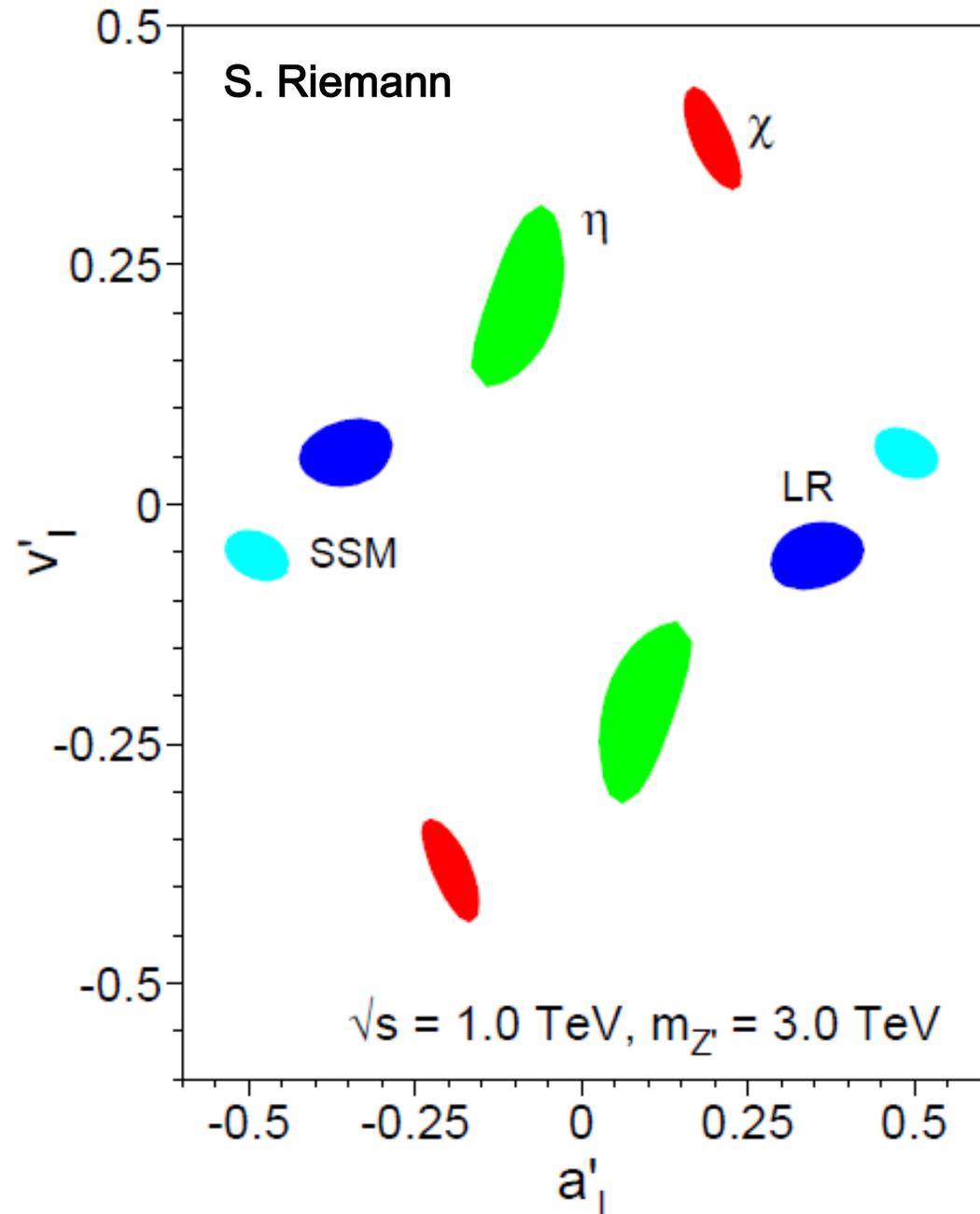
- The role of the LC is to make detailed studies of any objects related to the new physics discovered at the LHC... either through direct production (e.g., Higgs, sleptons, gauginos) or via 'indirect' precision measurements (e.g., Z' coupling determinations , etc)
- The 'canonical' version is LC500.. but LC1000 (or above) may be necessary to explore the NP found by the LHC .
  - The mass scale of NP is the most critical piece of info that the LHC can provide so that the necessary LC  $\sqrt{s}$  can be 'determined' (i.e., physics 'determined' & not by \$\$)
- There is already some evidence along these lines
- Warning: Everything I say has some caveats !

## We will consider two traditional examples\* :

- Indirect Z' coupling determination via 'contact' interactions
  - Gaugino & slepton accessibility for detailed studies
- These results seem to provide us with some evidence already that going to  $\sqrt{s} = 1$  TeV for the LC may be necessary

\* One more will be left to some backup slides

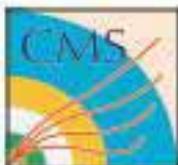
# ILC Indirect Z' Coupling Determinations



This has been 'bread & butter' LC physics for >2 decades & is quite mature.

These studies tell us that if  $M_{Z'} < 3-4\sqrt{s}$  & known from the LHC then interesting **coupling info** can be extracted aiding in Z' identification by detailed **cross section and asymmetry measurements using polarized beams**

Where are we now?



# Z' searches



## Search for resonance in di-lepton mass distribution

- Identify 2 muons with  $p_T > 20$  GeV or 2 electrons with  $p_T > 25$  GeV
- Dilepton invariant mass spectra consistent with SM expectations
- No sign of new resonance
- $Z'_{SSM}$  with Standard-Model-like couplings can be excluded below 1140 GeV
- Superstring-inspired  $Z'$  excluded below 887 GeV
- RS Kaluza-Klein gravitons below 855–1079 GeV for couplings of 0.05–0.1

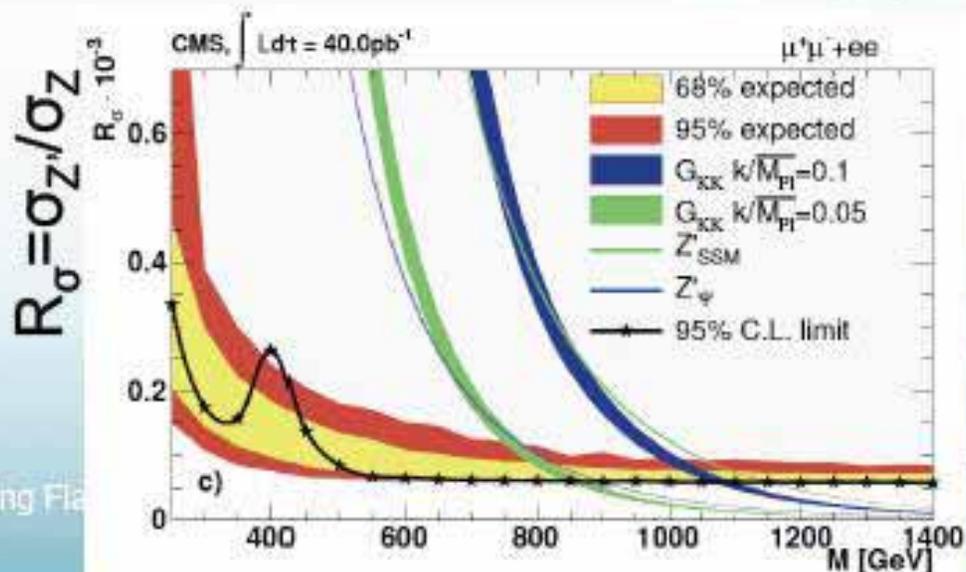
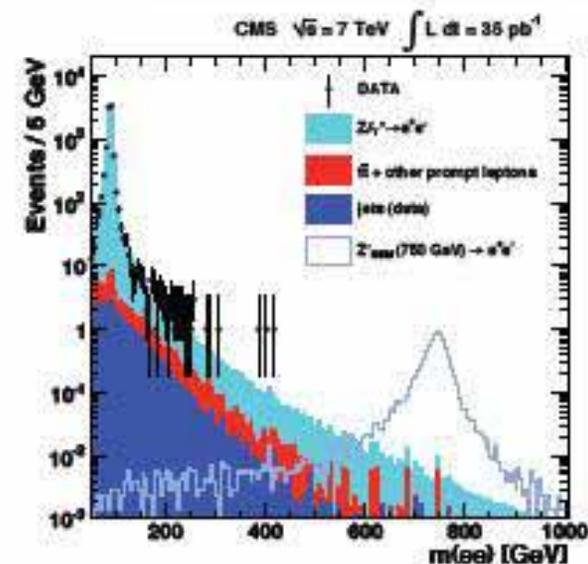
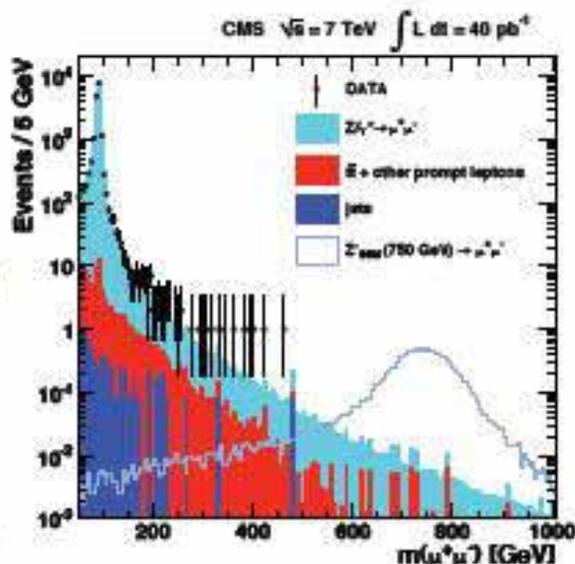
(all at 95% C.L.)

La Thuile 2011

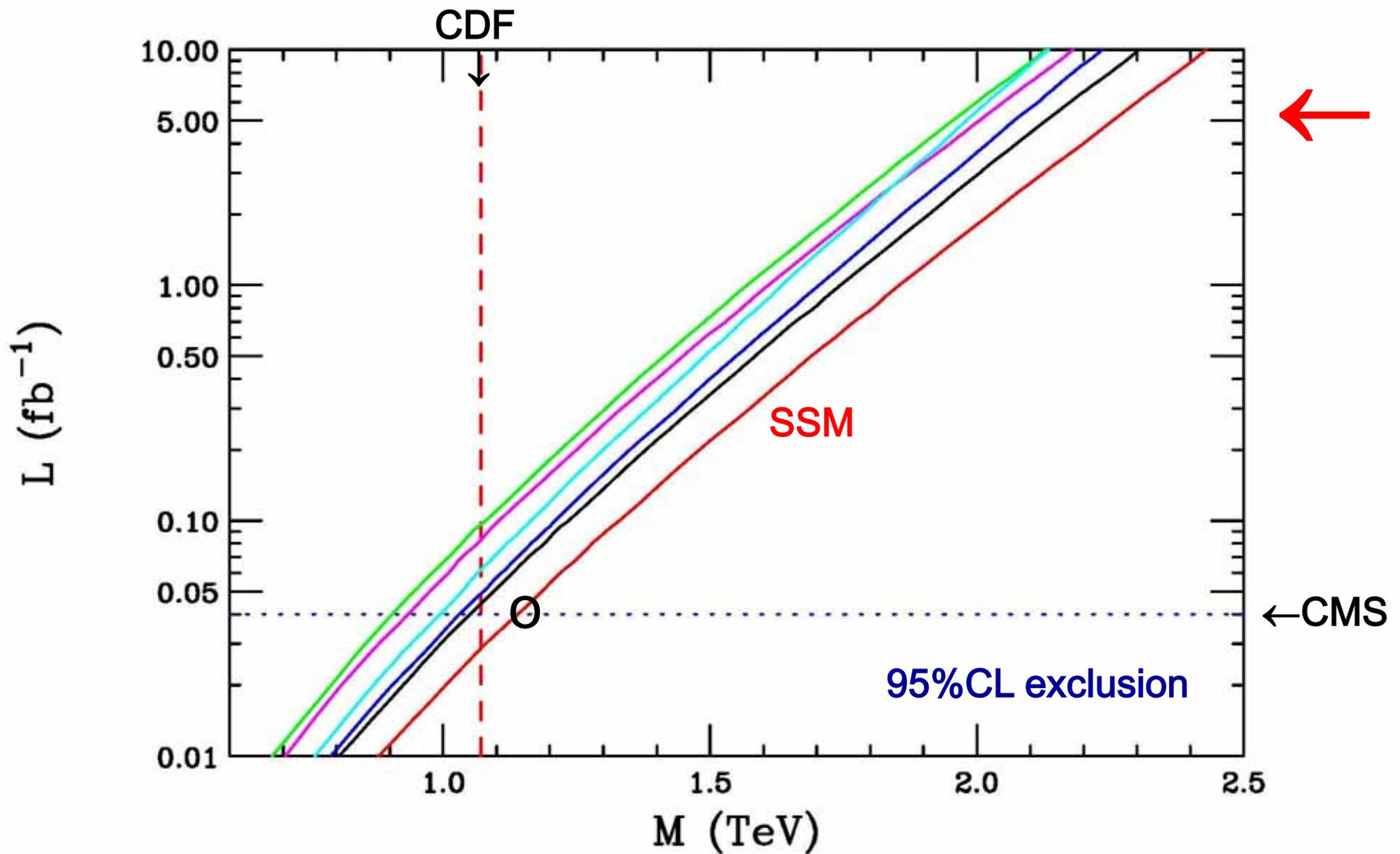
Hemming Fla

EXO-10-013

3/4/2011



# CMS & ATLAS already exceed the Tevatron searches



If no  $Z'$  is seen with  $\sim 5 \text{ fb}^{-1}$  then if one does exist it's too heavy/weakly coupled for LC500. In some models things may already be iffy due to internal relations.

# Example



## W' searches

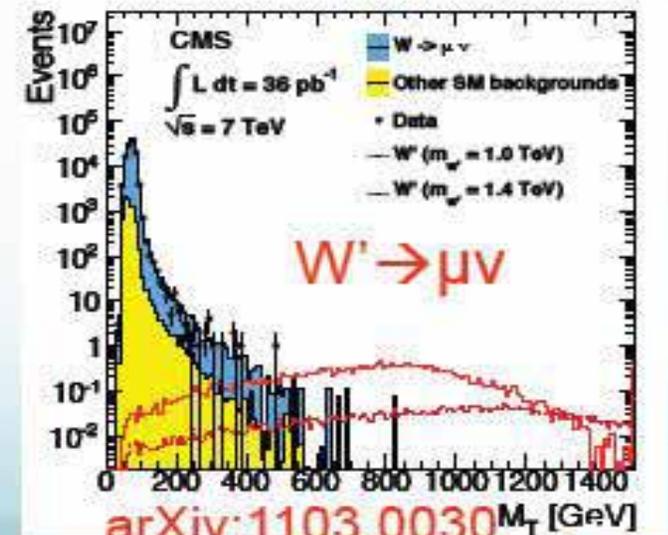
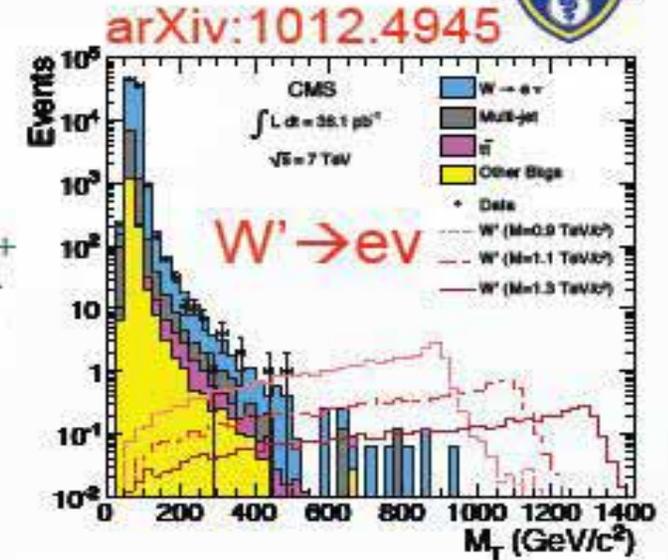
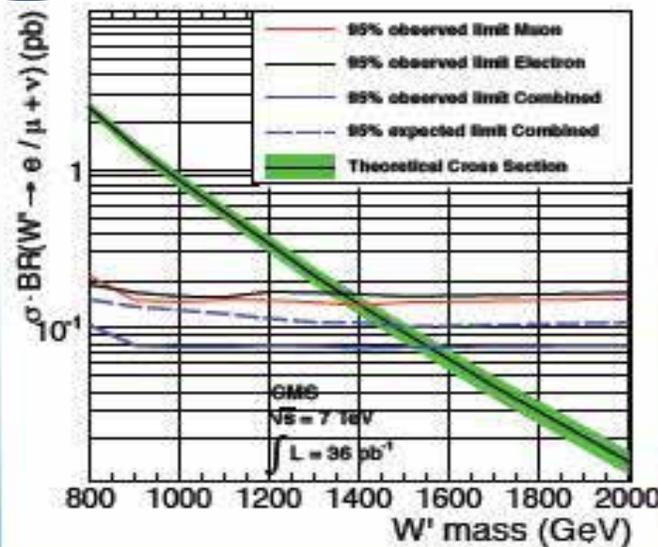


- W' a heavy analogue of SM W with same couplings
- identify high  $p_T$  lepton (e or  $\mu$ ),  $p_T > 30$  GeV
- Search for peak/enhancement in transverse mass spectrum (e/ $\mu$  + missing transverse energy)

$$M_T = \sqrt{2 \cdot p_T \cdot E_T^{\text{miss}} \cdot (1 - \cos \Delta\phi_{\mu,\nu})}$$

- Data agree with SM expectation
  - from  $W' \rightarrow e\nu$  channel exclude  $W'$  masses below 1.36 TeV
  - from  $W' \rightarrow \mu\nu$  channel exclude  $W'$  masses below 1.40 TeV
- Combination of e and  $\mu$  channel results in 95% CL exclusion of  $W'$  masses below 1.58 TeV

more stringent limit compared D0 & CDF (1.1 TeV)

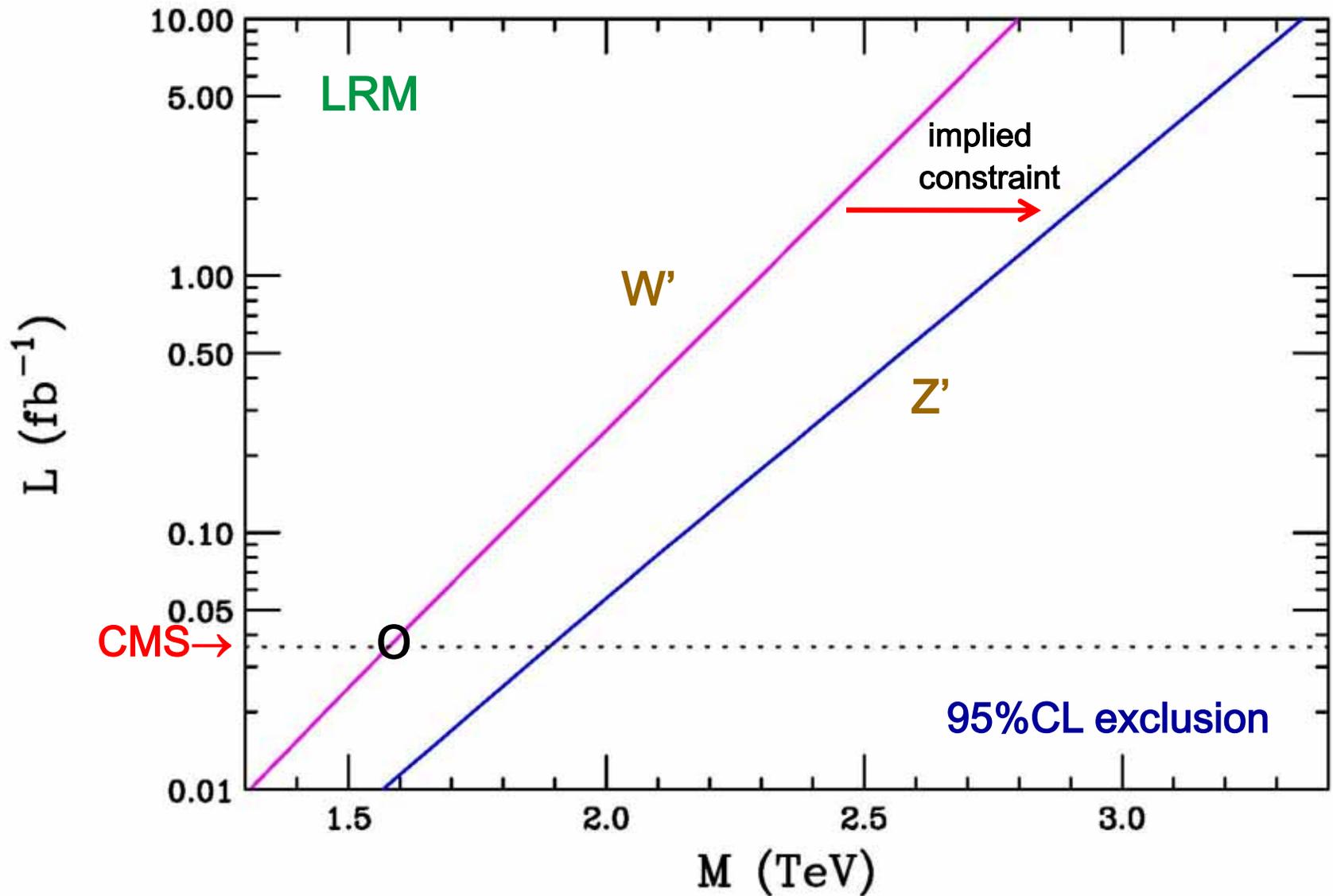


March 4th, 2011

3/4/2011  
La Thuile 2011

This is the vanilla LRM W'.... The LRM Z' is NOT independent!

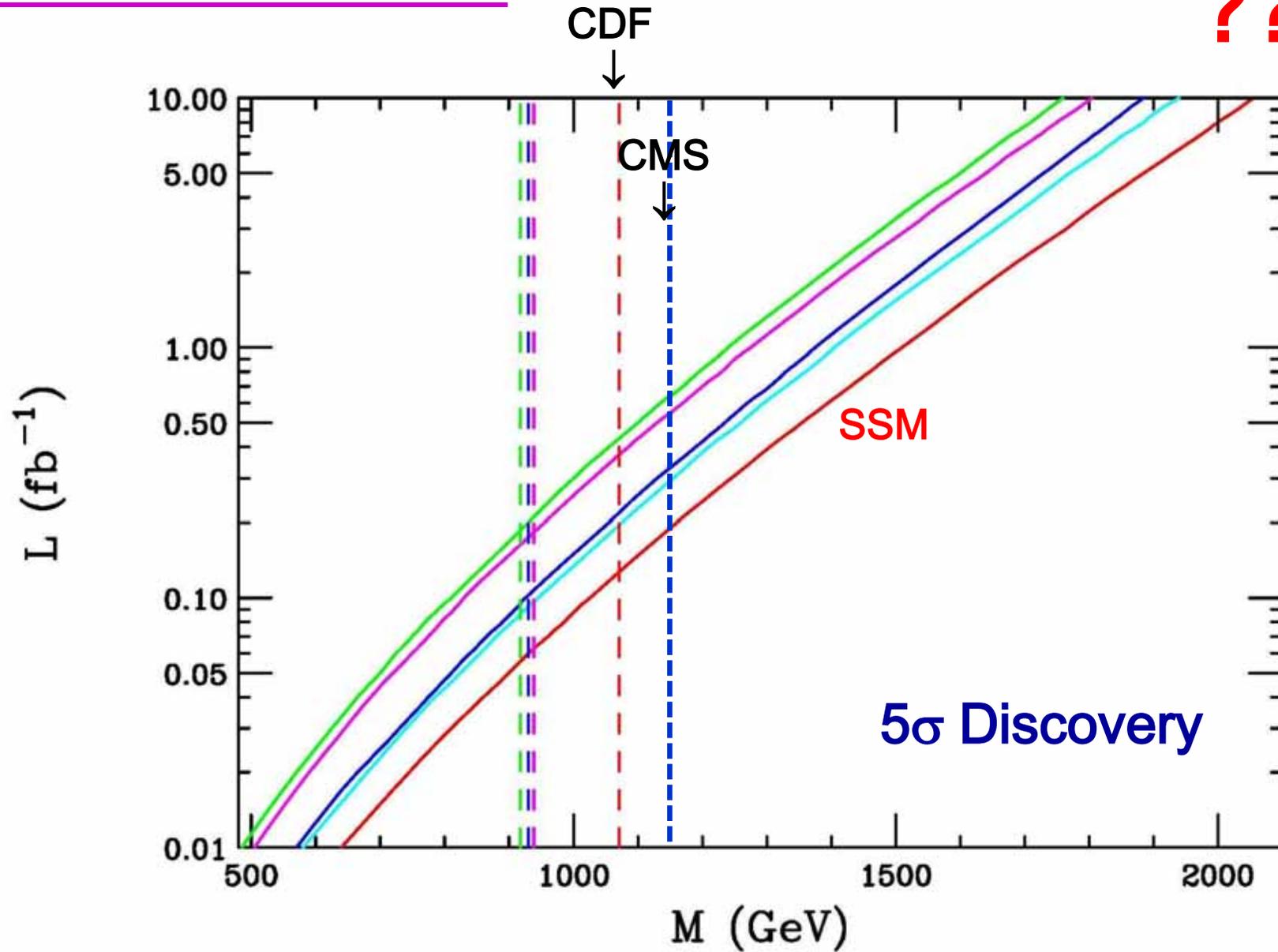
The  $W'$  and  $Z'$  masses are linked in the LRM by SSB so constraining the  $W'$  also leads to a bound on  $Z'$  :



The vanilla LRM  $Z'$  is already  $>1.89$  TeV  $\sim 3.8 \sqrt{s}$  of the LC500 & likely beyond the mass range to be probed there..

# The Near Future

???

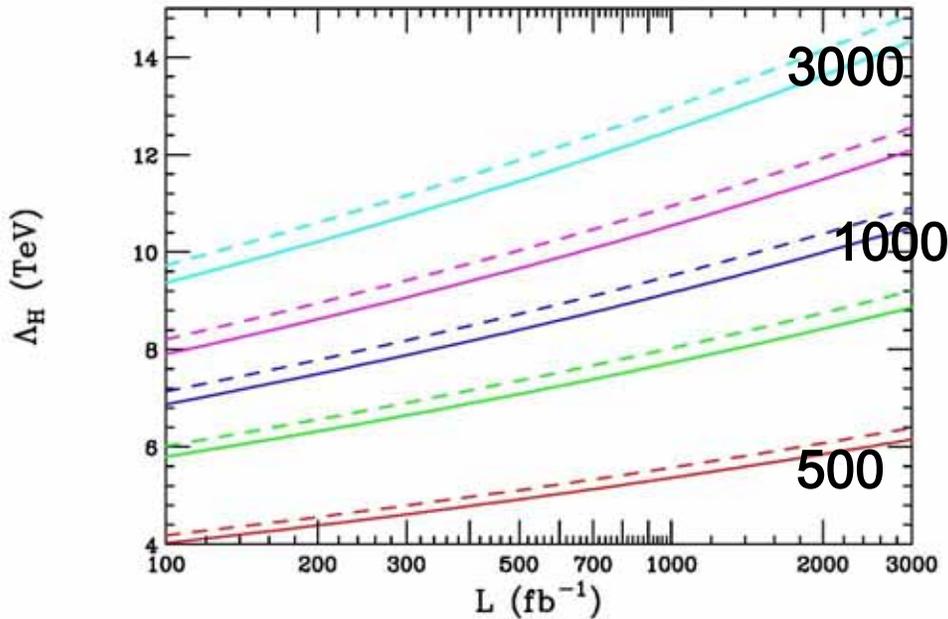


Discovery in the  $\sim 1.5$ - $2 \text{ TeV}$  range will be possible by the end of 2012  
No likely discovery possible above CMS limit until  $\sim 200 \text{ pb}^{-1}$  is achieved

# BTW...ADD visibility at LC500 is almost excluded

Gravity is Universal !

## LC Search Reach



LHC 95%CL exclusion in dijets is already approaching the LC500 discovery reach !

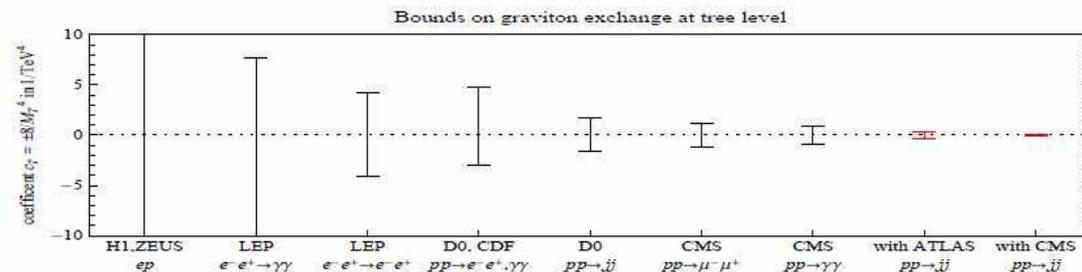
A.Strumia, Moriond11

KK graviton exchange is a dim-8 contact int. that can be seen at both LHC & LC.

But the LC500 window is closing...

Bounds on  $M_T$

| Experiment                 | Process                                     | +        | -        |
|----------------------------|---|----------|----------|
| LEP                        | $e^+e^- \rightarrow \gamma\gamma$           | 0.93 TeV | 1.01 TeV |
| LEP                        | $e^+e^- \rightarrow e^+e^-$                 | 1.18 TeV | 1.17 TeV |
| CDF                        | $p\bar{p} \rightarrow e^+e^-, \gamma\gamma$ | 0.99 TeV | 0.96 TeV |
| DØ                         | $p\bar{p} \rightarrow e^+e^-, \gamma\gamma$ | 1.28 TeV | 1.14 TeV |
| DØ                         | $p\bar{p} \rightarrow jj$                   | 1.48 TeV | 1.48 TeV |
| CMS at 7 TeV with 34/pb    | $pp \rightarrow \gamma\gamma$               | 1.72 TeV | 1.70 TeV |
| CMS at 7 TeV with 40/pb    | $pp \rightarrow \mu^-\mu^+$                 | 1.6 TeV  | 1.6 TeV  |
| ATLAS at 7 TeV with 3.1/pb | $pp \rightarrow jj$                         | 2.2 TeV  | 2.1 TeV  |
| CMS at 7 TeV with 36/pb    | $pp \rightarrow jj$                         | 4.2 TeV  | 3.4 TeV  |



→ Typical SUSY studies at the LC usually involve the (precision) determination of slepton and electroweak gaugino masses & couplings

To perform these studies both these types of sparticles must be, at the very least, kinematically accessible at a given  $\sqrt{s}$ . Let's assume this is a requirement..

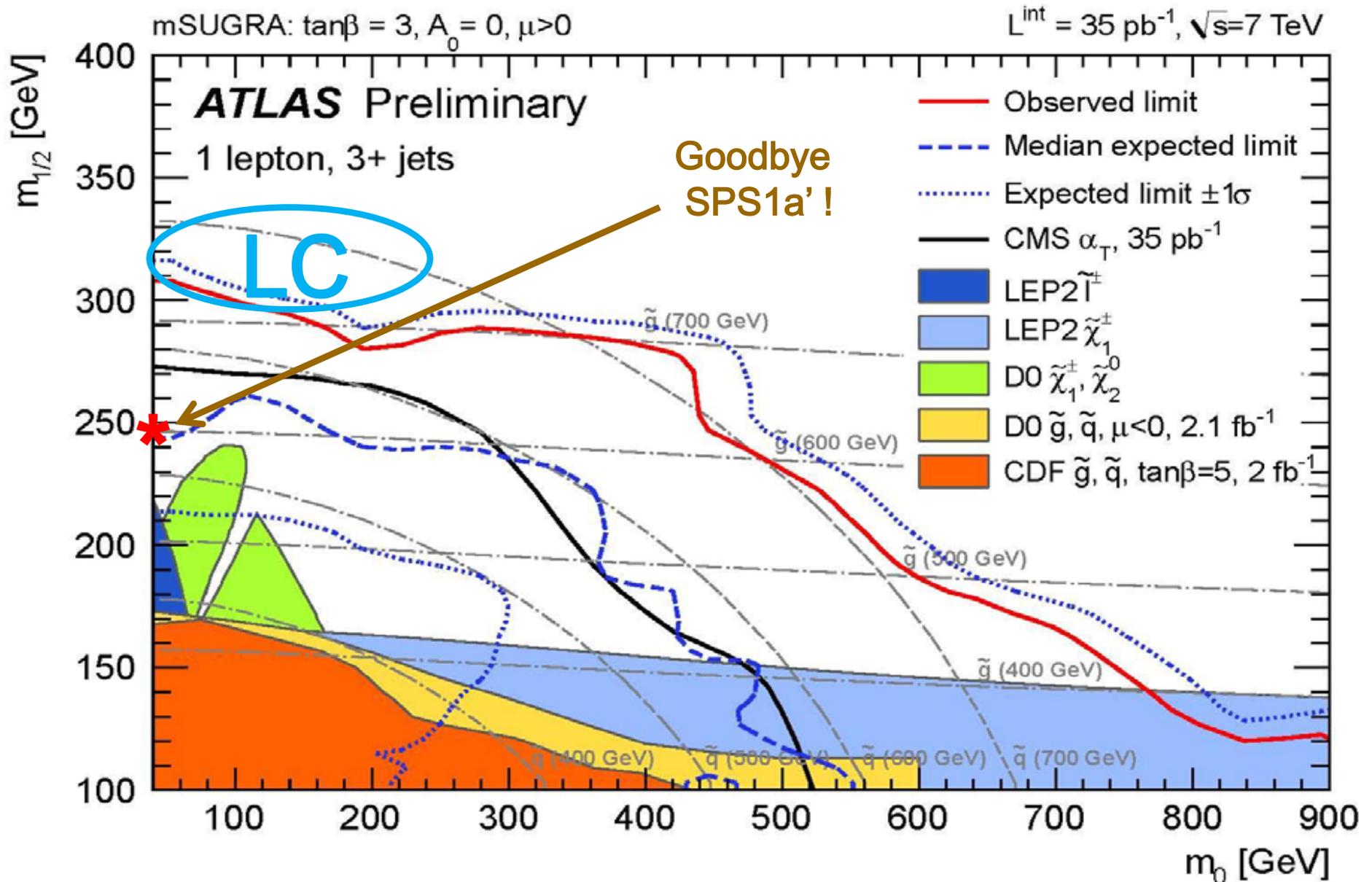
As the LHC lumi is accumulated & if SUSY is not seen life becomes quite tough for LC500 (as well as for some theorists!)

Already life has gotten tough in some scenarios like mSUGRA...

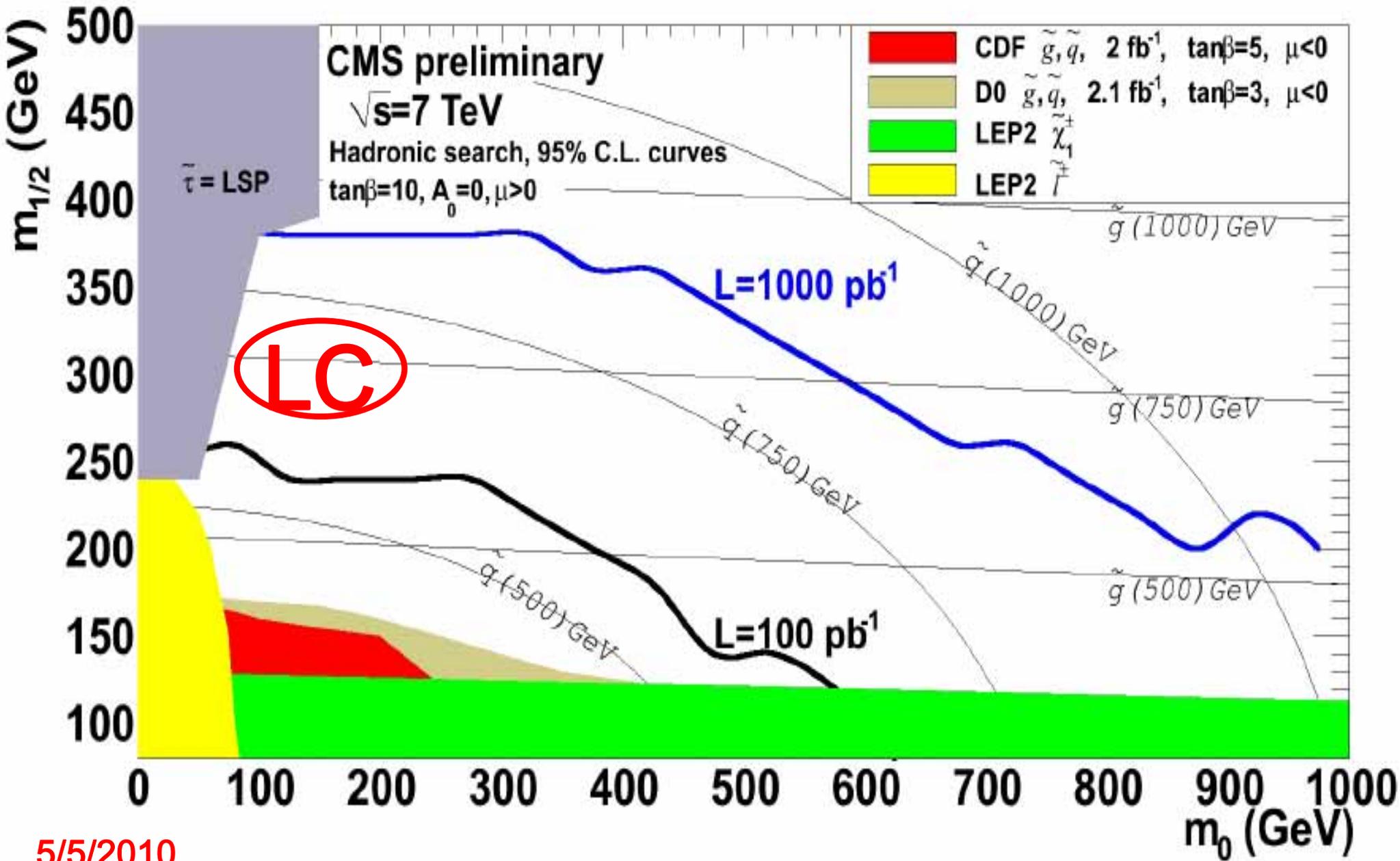
# SUSY searches at the LHC have gotten real:

## The Amazing Power of $\sqrt{s}$ !

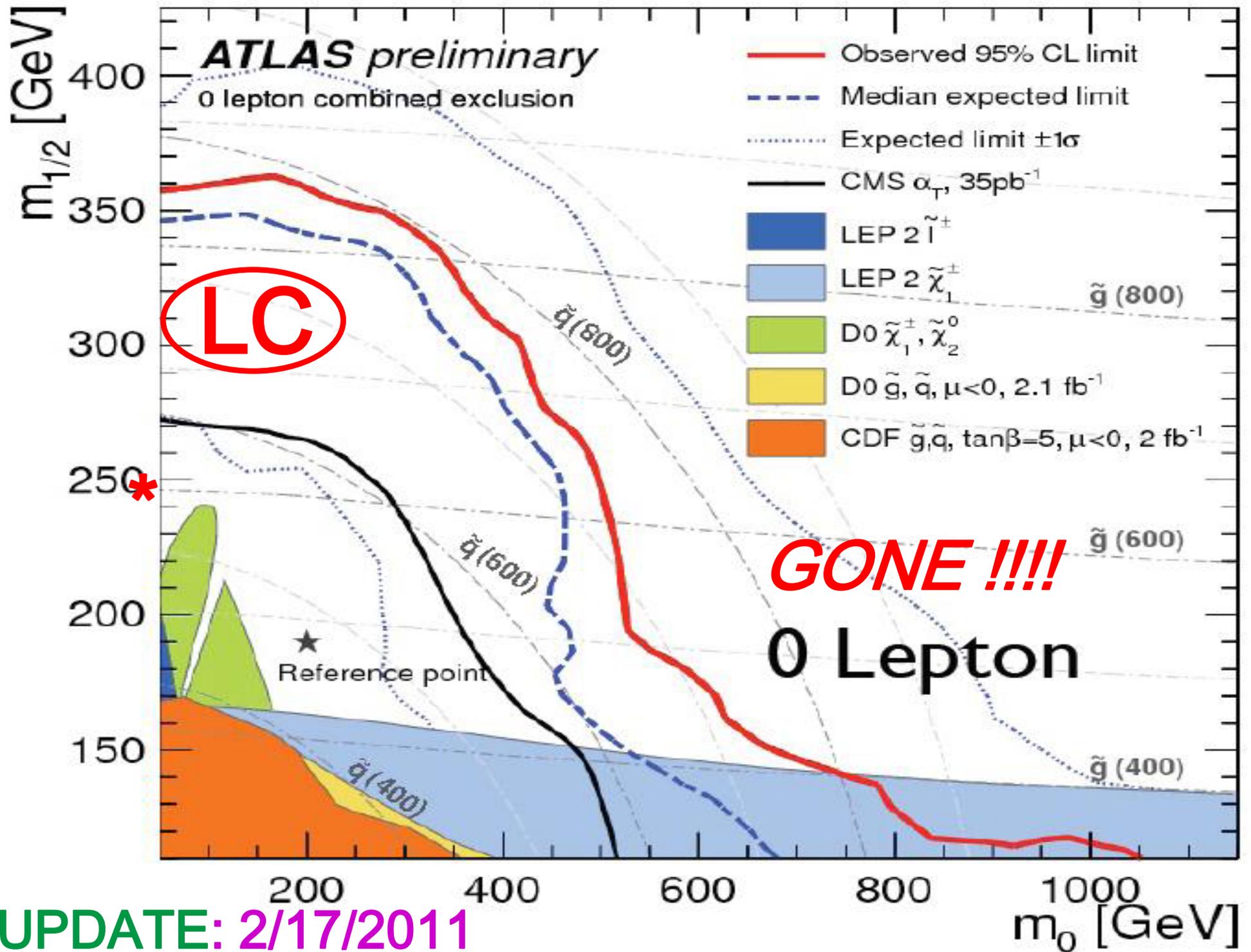
1/27/2011

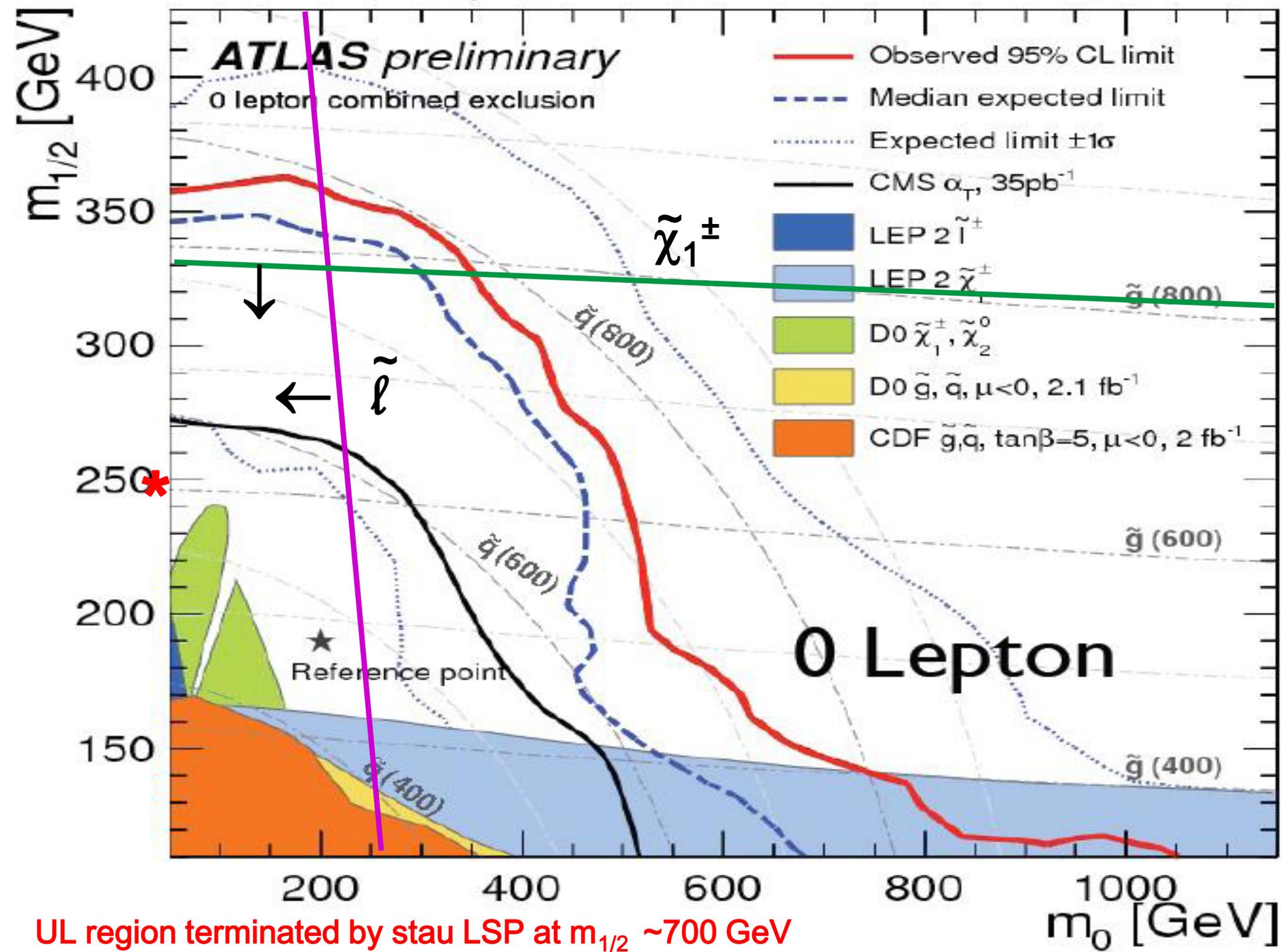


**Expectation: If no signal @ 7 TeV w/ 1 fb<sup>-1</sup>, then a 500 GeV LC is not a good place to study mSUGRA/CMSSM**

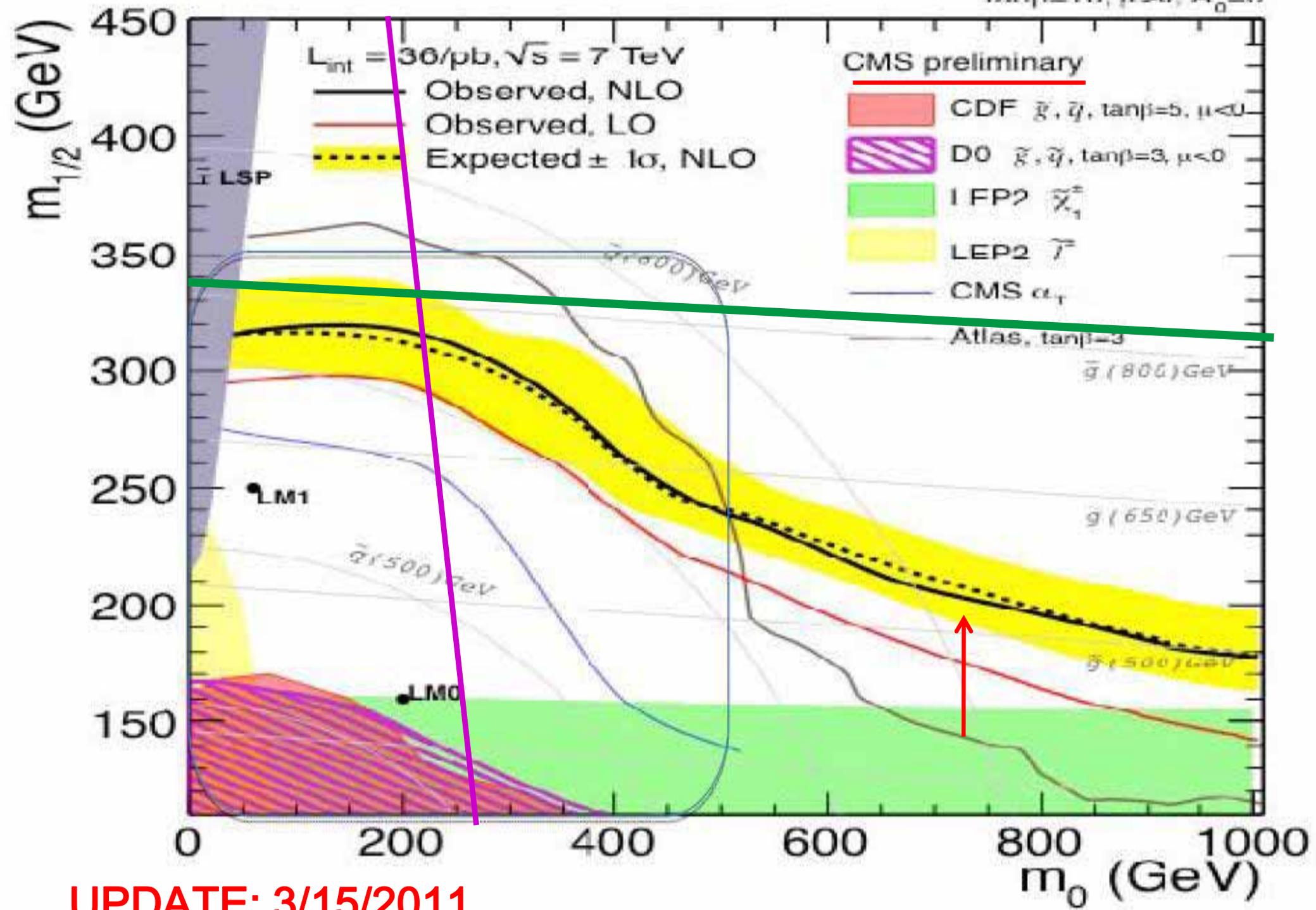


5/5/2010

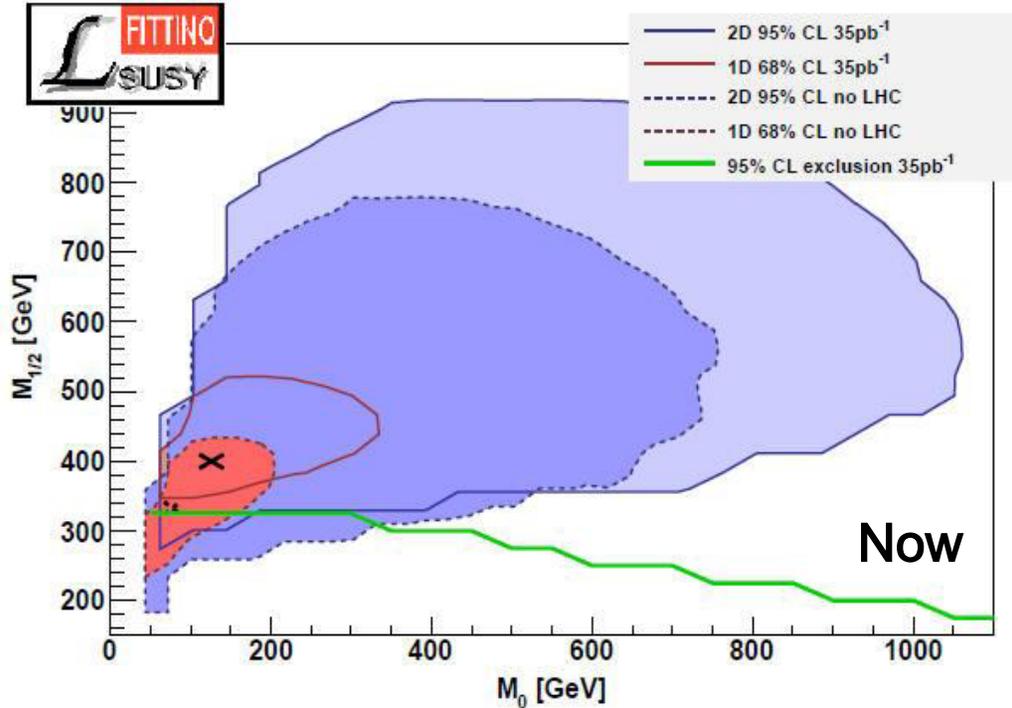




$\tan\beta=10, \mu>0, A_0=0$

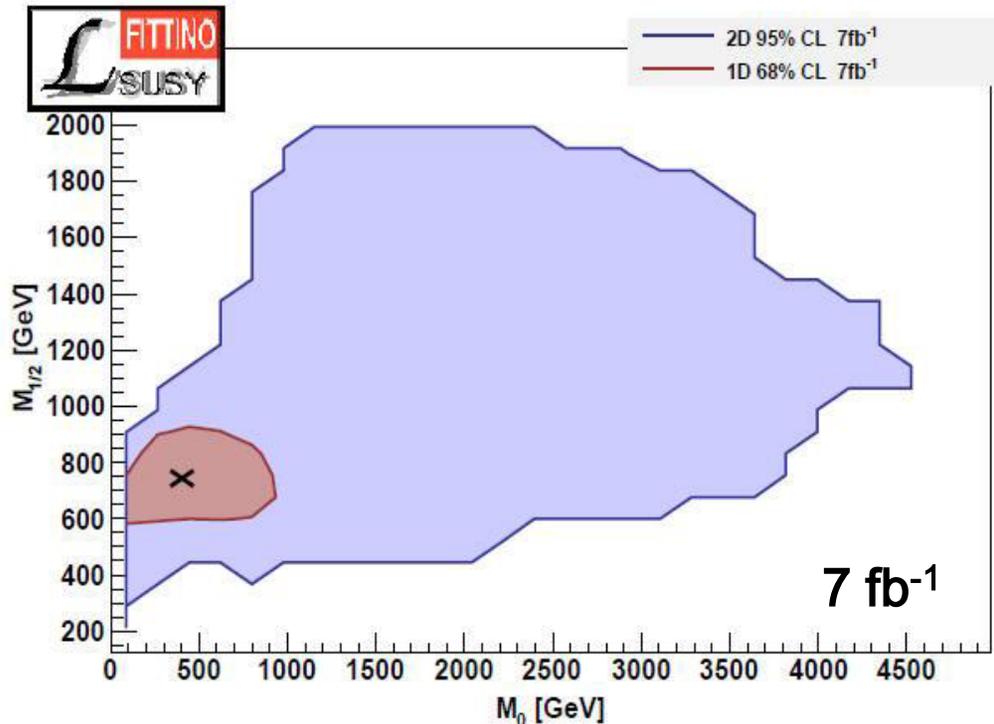


UPDATE: 3/15/2011



# mSUGRA FITS

If no signal is found then the mSUGRA best fits to data will be pushed upwards to higher & higher masses...



They tend to like smaller  $m_0$  values & slide upwards as nothing is seen

So for mSUGRA the answer is *in* :

LC500 is NOT a great place to study mSUGRA !!

→ But remember **SUSY**  $\neq$  **mSUGRA** !

- So the more important question is: ‘Does a more **general version** of the MSSM allow for SUSY to remain ‘accessible’ at LC500 if the LHC sees **nothing** during 2011-12 @ 7 TeV ?’
- We can get a **partial** answer to this question based on a set of more general pMSSM models that some of us studied in detail for the **LHC**..but there’s still a lot we don’t know...

# Issues:

- The general MSSM is too difficult to study due to the large number of soft SUSY breaking parameters ( $\sim 100$ ).
- Analyses are limited to specific SUSY breaking scenarios having only a few parameters...can we be more general ?

## Model Generation Assumptions :

- The most general, CP-conserving MSSM with R-parity
- Minimal Flavor Violation at the TeV scale
- The lightest neutralino is the LSP & a thermal relic.
- The first two sfermion generations are degenerate & have negligible Yukawa's.
- No assumptions about SUSY-breaking or unification

This leaves us with the pMSSM:

→ the MSSM with 19 real, TeV-scale parameters...

# 19 pMSSM Parameters

10 sfermion masses:  $m_{Q_1}, m_{Q_3}, m_{u_1}, m_{d_1}, m_{u_3}, m_{d_3}, m_{L_1},$   
 $m_{L_3}, m_{e_1}, m_{e_3}$

3 gaugino masses:  $M_1, M_2, M_3$

3 tri-linear couplings:  $A_b, A_t, A_\tau$

3 Higgs/Higgsino:  $\mu, M_A, \tan\beta$

# How? Perform 2 Random Scans

## Flat Priors

emphasizes moderate masses

$$100 \text{ GeV} \leq m_{\text{sfermions}} \leq 1 \text{ TeV}$$

$$50 \text{ GeV} \leq |M_1, M_2, \mu| \leq 1 \text{ TeV}$$

$$100 \text{ GeV} \leq M_3 \leq 1 \text{ TeV}$$

$$\sim 0.5 M_Z \leq M_A \leq 1 \text{ TeV}$$

$$1 \leq \tan\beta \leq 50$$

$$|A_{t,b,\tau}| \leq 1 \text{ TeV}$$

## Log Priors

emphasizes lower masses but also extends to higher masses

$$100 \text{ GeV} \leq m_{\text{sfermions}} \leq 3 \text{ TeV}$$

$$10 \text{ GeV} \leq |M_1, M_2, \mu| \leq 3 \text{ TeV}$$

$$100 \text{ GeV} \leq M_3 \leq 3 \text{ TeV}$$

$$\sim 0.5 M_Z \leq M_A \leq 3 \text{ TeV}$$

$$1 \leq \tan\beta \leq 60$$

$$10 \text{ GeV} \leq |A_{t,b,\tau}| \leq 3 \text{ TeV}$$

- Flat Priors :  $10^7$  models scanned , 68422 survive
- Log Priors :  $2 \times 10^6$  models scanned , 2908 survive

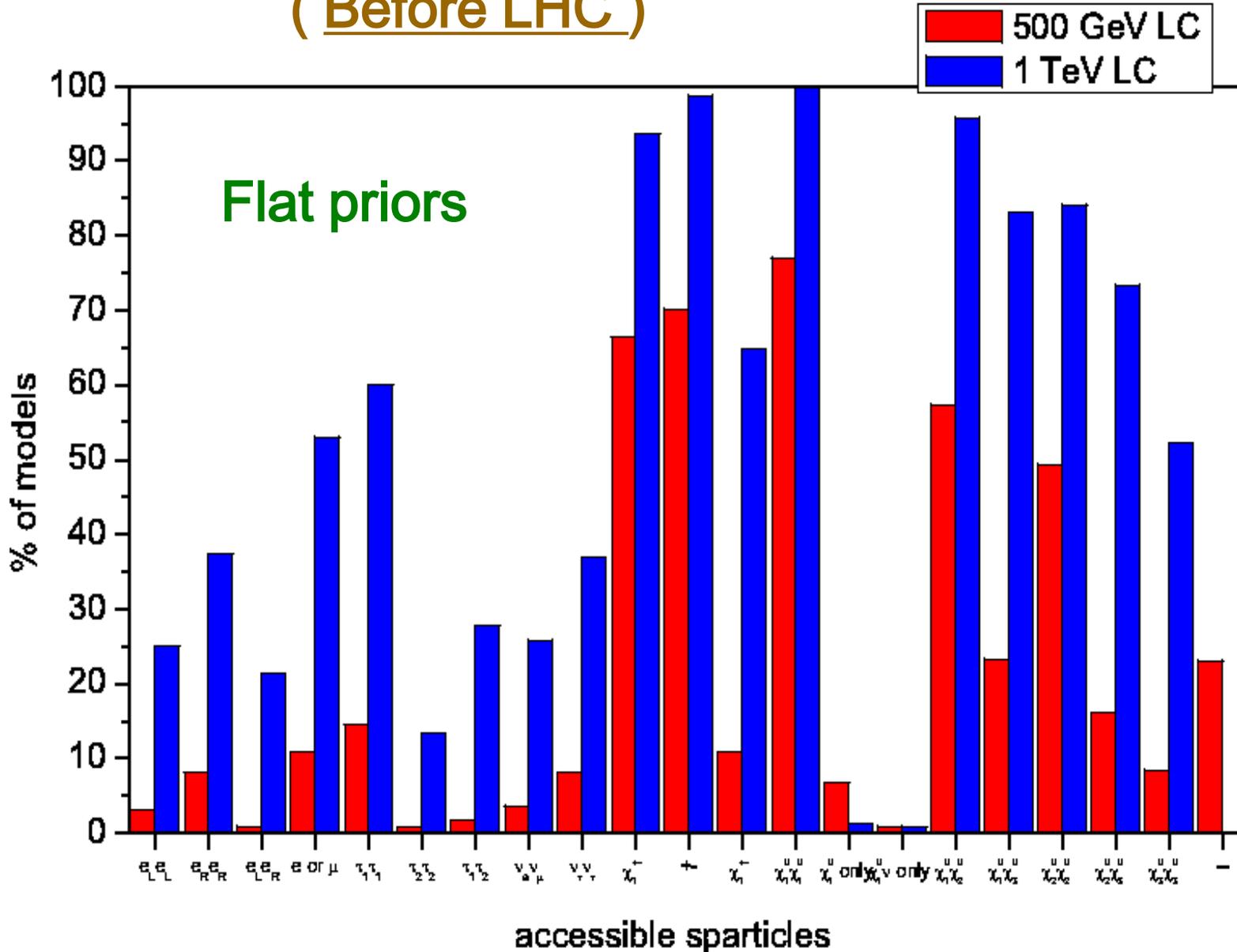
→ Comparison of these two scans will show the prior sensitivity.

## Some Constraints

- **W/Z ratio**                       **$b \rightarrow s \gamma$**
  - **$\Delta(g-2)_\mu$**                        **$\Gamma(Z \rightarrow \text{invisible})$**
  - **Meson-Antimeson Mixing**
  - **$B_s \rightarrow \mu\mu$**                        **$B \rightarrow \tau\nu$**
- 
- **DM density:  $\Omega h^2 < 0.121$ . We treat this only as an *upper bound* on the neutralino thermal relic contribution**
  - **Direct Detection Searches for DM (CDMS, XENON...)**
  - **LEP and Tevatron Direct Higgs & SUSY searches : there are *many* of these searches & they are quite complicated with many caveats.... **These needed to be 'revisited' for the more general case considered here  $\rightarrow$  simulations limit model set size  $\sim 1$  core-century for set generation****

# Model Set Kinematic Accessibility at the ILC : I

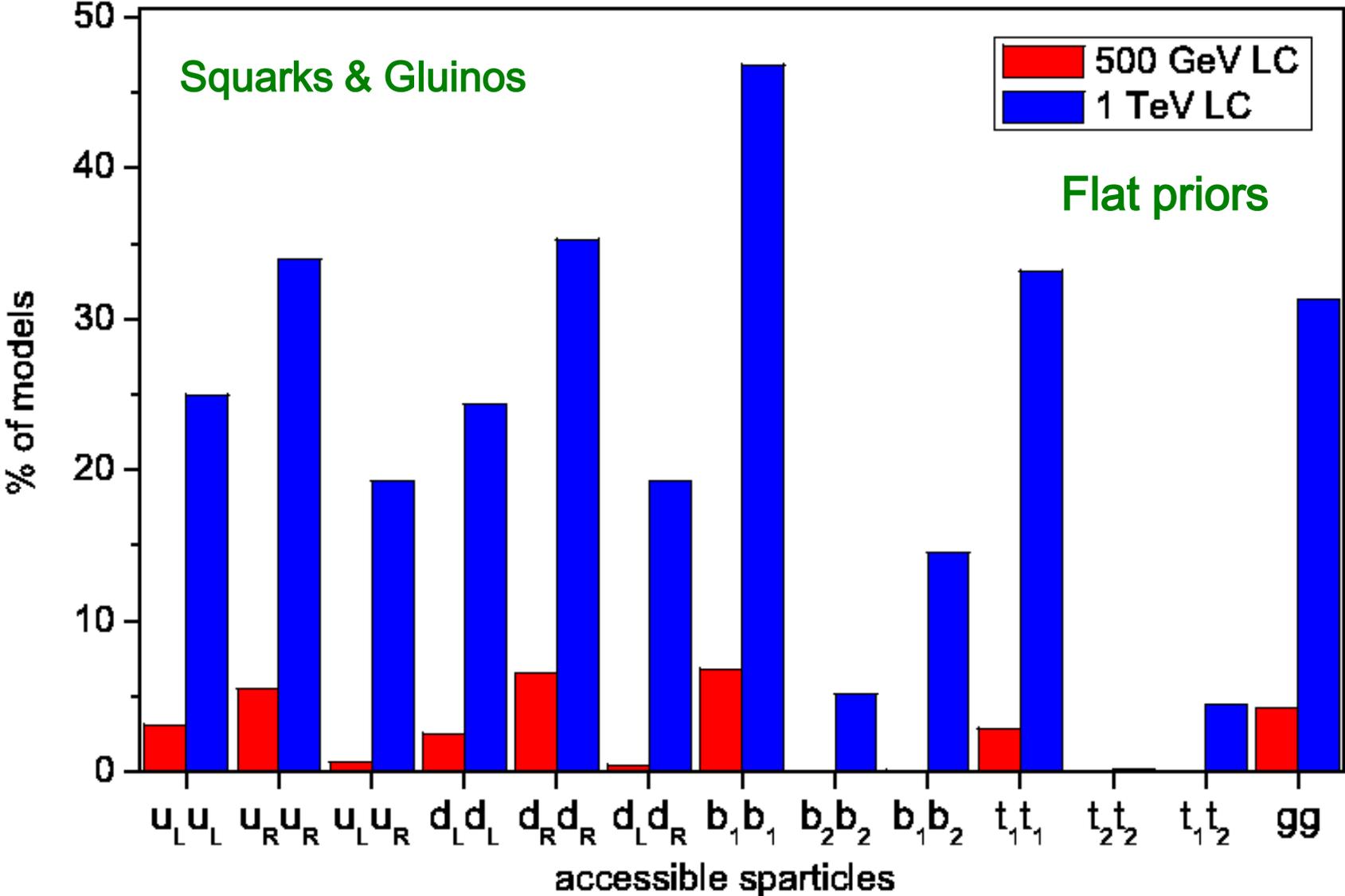
( Before LHC )



| Final State                               |
|---|
| $\tilde{e}_L^+ \tilde{e}_L^-$             |
| $\tilde{e}_R^+ \tilde{e}_R^-$             |
| $\tilde{e}_L^\pm \tilde{e}_R^\mp$         |
| $\tilde{\mu}_L^+ \tilde{\mu}_L^-$         |
| $\tilde{\mu}_R^+ \tilde{\mu}_R^-$         |
| Any selectron or smuon                    |
| $\tilde{\tau}_1^+ \tilde{\tau}_1^-$       |
| $\tilde{\tau}_2^+ \tilde{\tau}_2^-$       |
| $\tilde{\tau}_1^\pm \tilde{\tau}_2^\mp$   |
| $\tilde{\nu}_{e\mu} \tilde{\nu}_{e\mu}^*$ |
| $\tilde{\nu}_\tau \tilde{\nu}_\tau^*$     |
| $\tilde{\chi}_1^+ \tilde{\chi}_1^-$       |
| Any charged sparticle                     |
| $\tilde{\chi}_1^\pm \tilde{\chi}_2^\mp$   |
| $\tilde{\chi}_1^0 \tilde{\chi}_1^0$       |
| $\tilde{\chi}_1^0 \tilde{\chi}_1^0$ only  |
| $\tilde{\chi}_1^0 + \tilde{\nu}$ only     |
| $\tilde{\chi}_1^0 \tilde{\chi}_2^0$       |
| $\tilde{\chi}_1^0 \tilde{\chi}_3^0$       |
| $\tilde{\chi}_2^0 \tilde{\chi}_2^0$       |
| $\tilde{\chi}_2^0 \tilde{\chi}_3^0$       |
| $\tilde{\chi}_3^0 \tilde{\chi}_3^0$       |
| Nothing                                   |

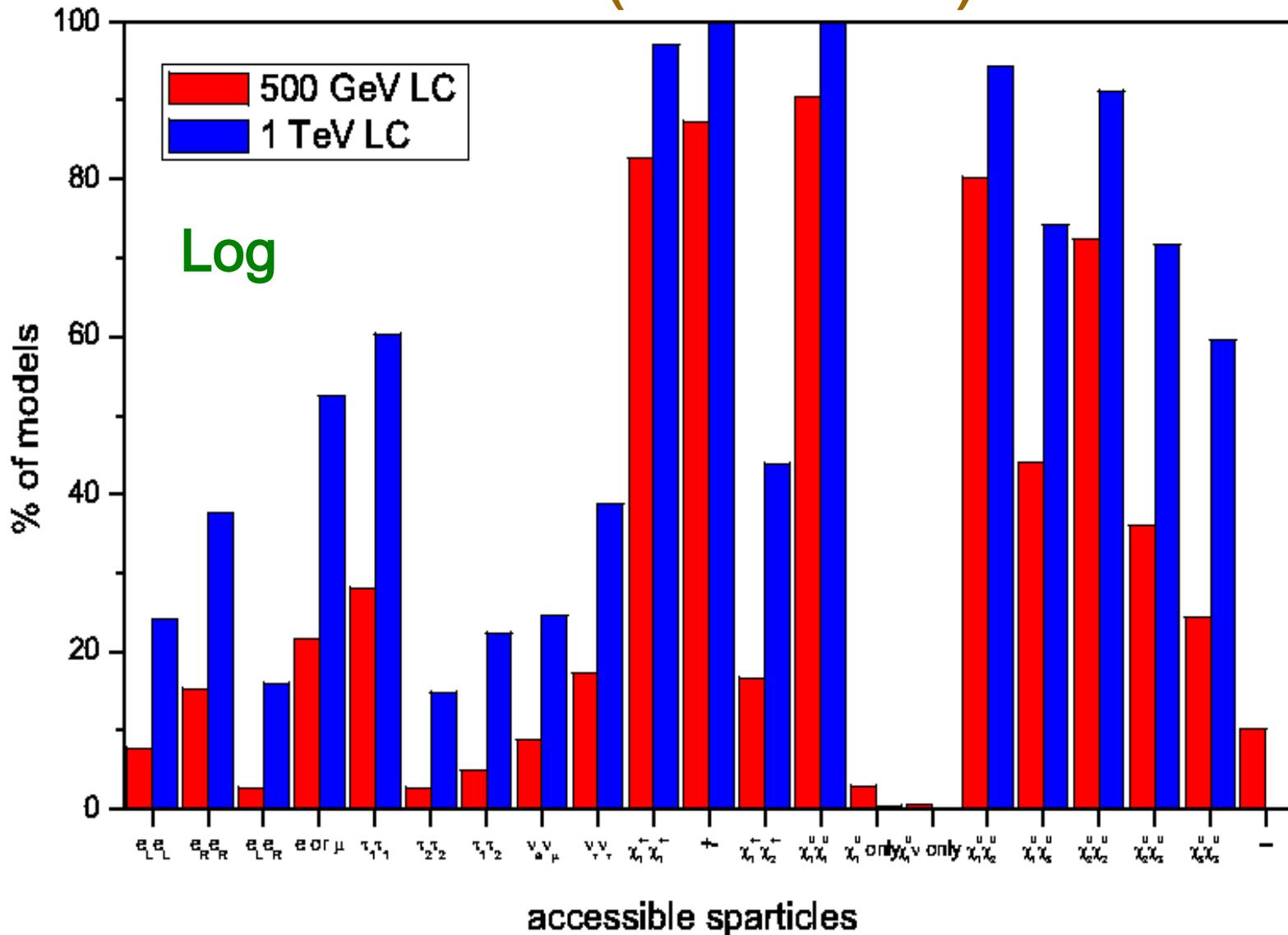
# Model Set Kinematic Accessibility at the ILC : II

( Before LHC )



# Model Set Kinematic Accessibility at the ILC : III

( Before LHC )

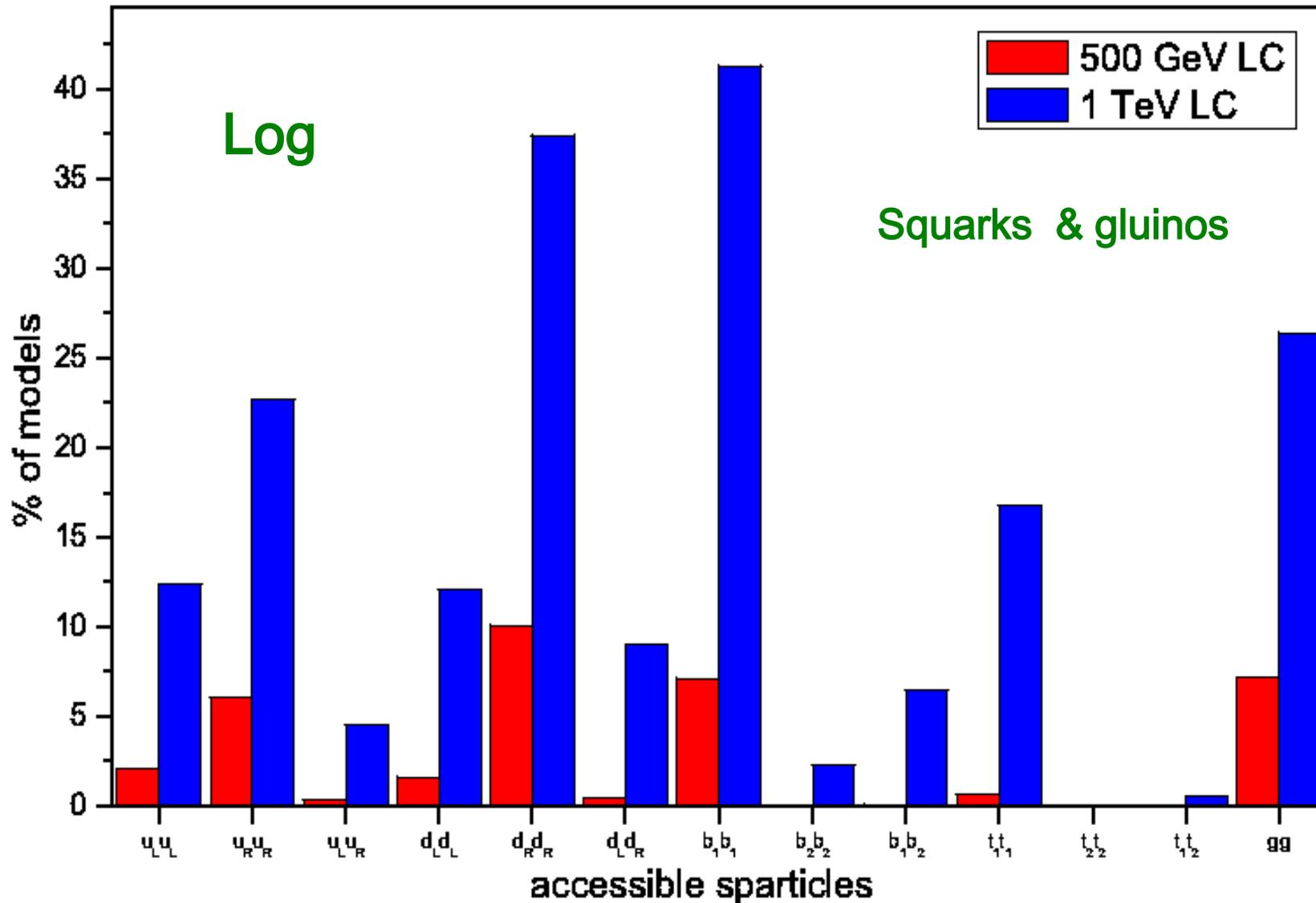


| Final State                               |
|---|
| $\tilde{e}_L^+ \tilde{e}_L^-$             |
| $\tilde{e}_R^+ \tilde{e}_R^-$             |
| $\tilde{e}_L^\pm \tilde{e}_R^\mp$         |
| $\tilde{\mu}_L^+ \tilde{\mu}_L^-$         |
| $\tilde{\mu}_R^+ \tilde{\mu}_R^-$         |
| Any selectron or smuon                    |
| $\tilde{\tau}_1^+ \tilde{\tau}_1^-$       |
| $\tilde{\tau}_2^+ \tilde{\tau}_2^-$       |
| $\tilde{\tau}_1^\pm \tilde{\tau}_2^\mp$   |
| $\tilde{\nu}_{e\mu} \tilde{\nu}_{e\mu}^*$ |
| $\tilde{\nu}_\tau \tilde{\nu}_\tau^*$     |
| $\tilde{\chi}_1^+ \tilde{\chi}_1^-$       |
| Any charged sparticle                     |
| $\tilde{\chi}_1^\pm \tilde{\chi}_2^\mp$   |
| $\tilde{\chi}_1^0 \tilde{\chi}_1^0$       |
| $\tilde{\chi}_1^0 \tilde{\chi}_1^0$ only  |
| $\tilde{\chi}_1^0 + \tilde{\nu}$ only     |
| $\tilde{\chi}_1^0 \tilde{\chi}_2^0$       |
| $\tilde{\chi}_1^0 \tilde{\chi}_3^0$       |
| $\tilde{\chi}_2^0 \tilde{\chi}_2^0$       |
| $\tilde{\chi}_2^0 \tilde{\chi}_3^0$       |
| $\tilde{\chi}_3^0 \tilde{\chi}_3^0$       |
| Nothing                                   |

# Model Set Kinematic Accessibility at the ILC : IV

T

( Before LHC )



# ATLAS SUSY Analyses w/ a Large Model Set

- We passed these points through the ATLAS inclusive MET analyses (@ both 7 & 14 TeV !), designed for mSUGRA , to explore this broader class of models (~150 core-yrs)
- We used the ATLAS SM backgrounds with their associated systematic errors, search analyses/cuts & criterion for SUSY discovery.
- We verified that we can approximately reproduce the 7 & 14 TeV ATLAS results for their benchmark mSUGRA models with our analysis techniques for each channel. ..BUT beware of some analysis differences:

# ATLAS

ISASUGRA generates spectrum  
& sparticle decays

Partial NLO cross sections using  
PROSPINO & CTEQ6M

Herwig for fragmentation &  
hadronization

GEANT4 for full detector sim

# US

SuSpect generates spectra  
with SUSY-HIT<sup>#</sup> for decays

NLO cross section for all 85  
processes using PROSPINO\*\*  
& CTEQ6.6M

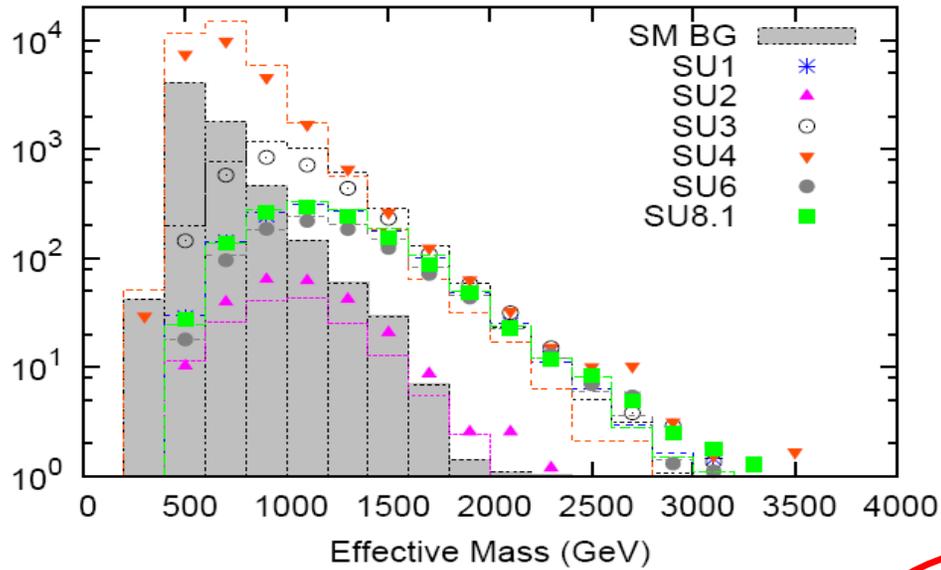
PYTHIA for fragmentation &  
hadronization

PGS4-ATLAS for fast detector  
simulation

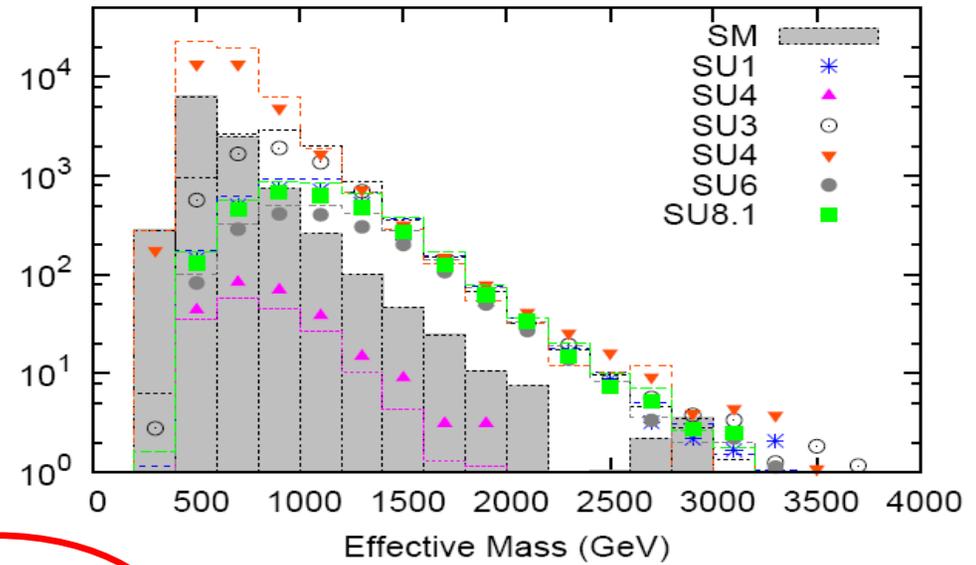
\*\* version w/ negative K-factor errors corrected

# version w/o negative QCD corrections, with 1<sup>st</sup> & 2<sup>nd</sup> generation fermion masses & other very numerous PS fixes included. e.g., explicit small  $\Delta m$  chargino decays, etc.

$M_{\text{eff}}$  distribution for 4-jet, 0 lepton analysis



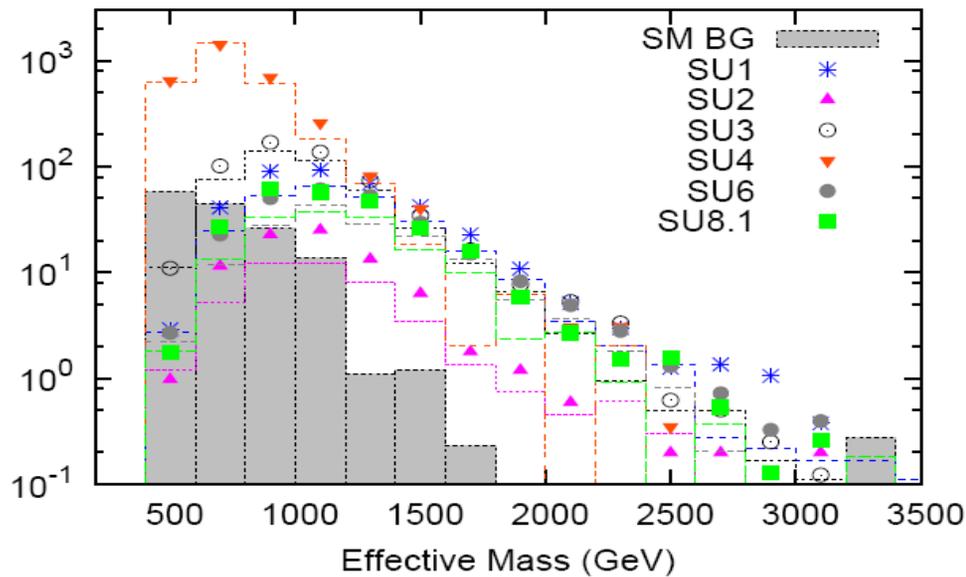
$M_{\text{eff}}$  distribution for 2-jet, 0 lepton analysis



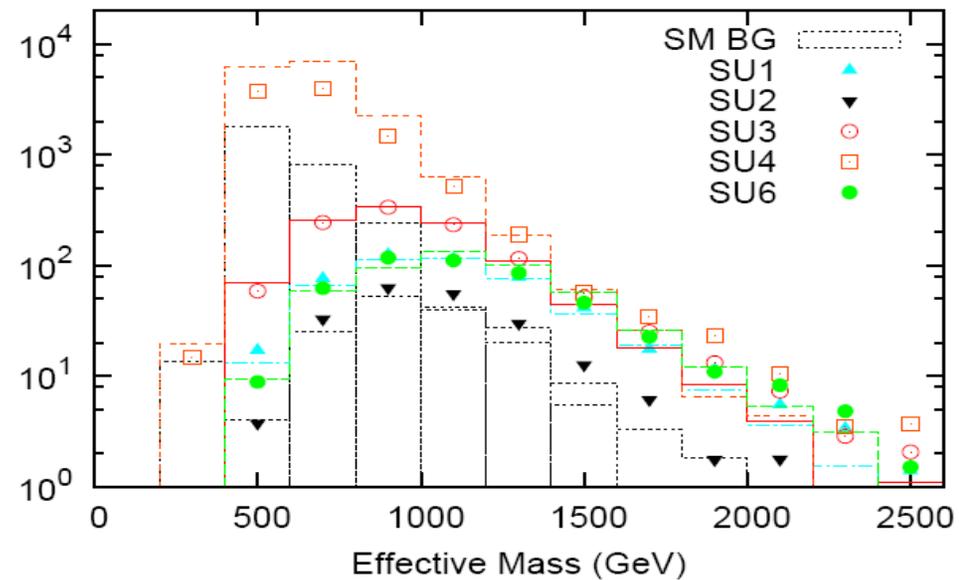
14 TeV

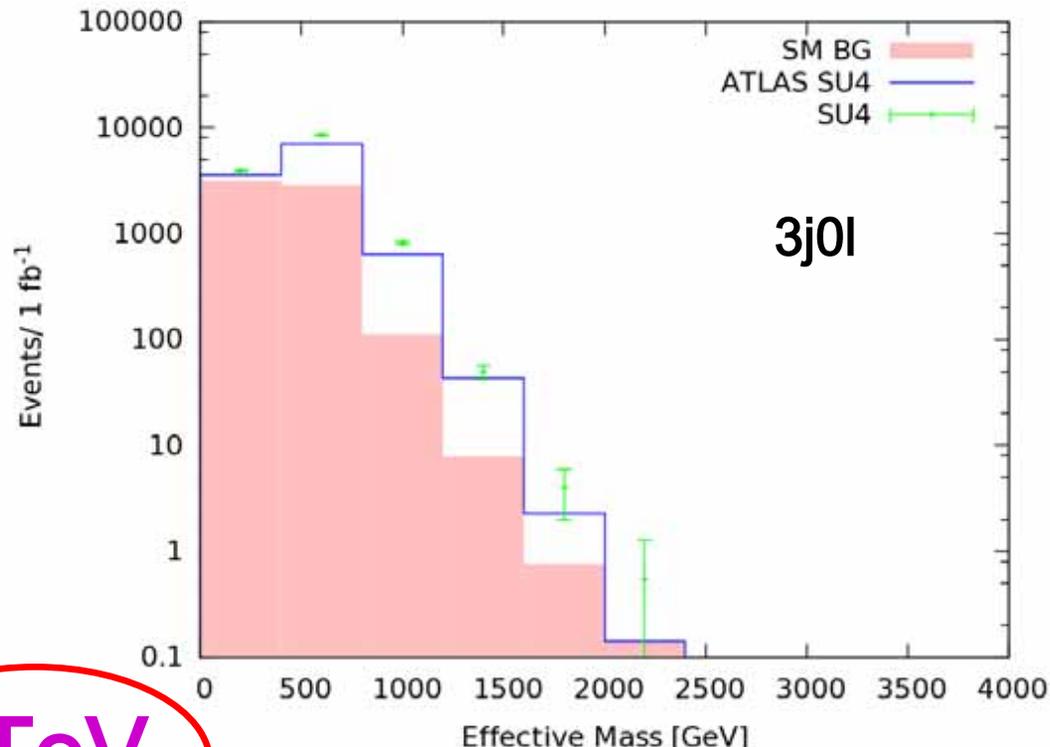
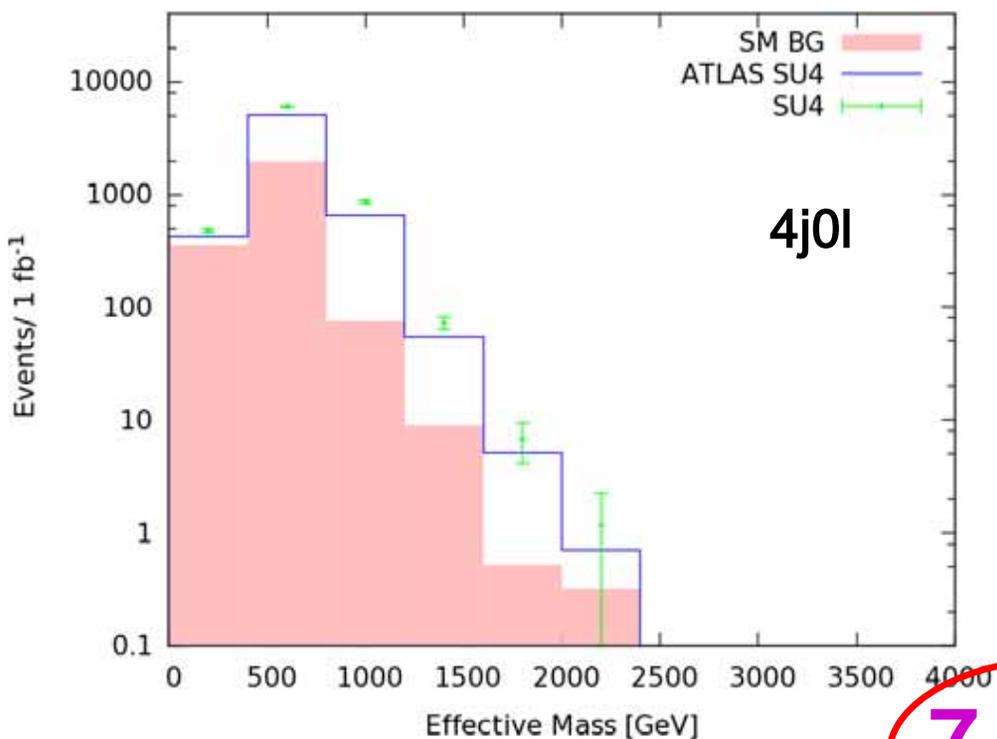
4j

$M_{\text{eff}}$  distribution for  $\tilde{\chi}^0_1$  lepton analysis

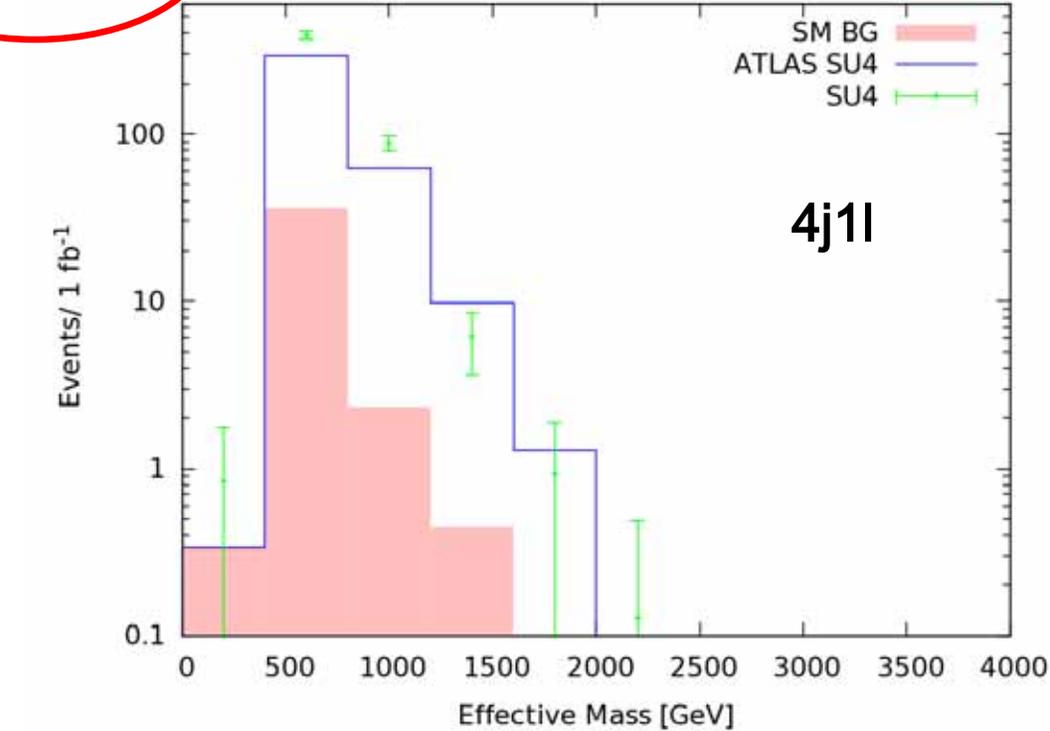
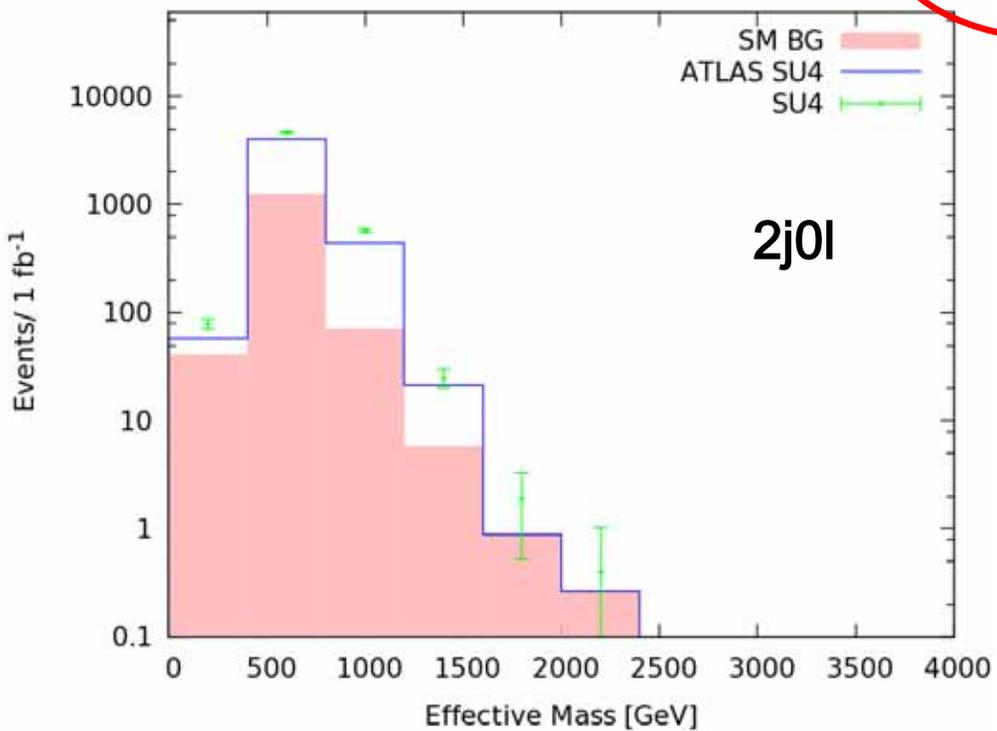


$M_{\text{eff}}$  distribution for b-jet analysis





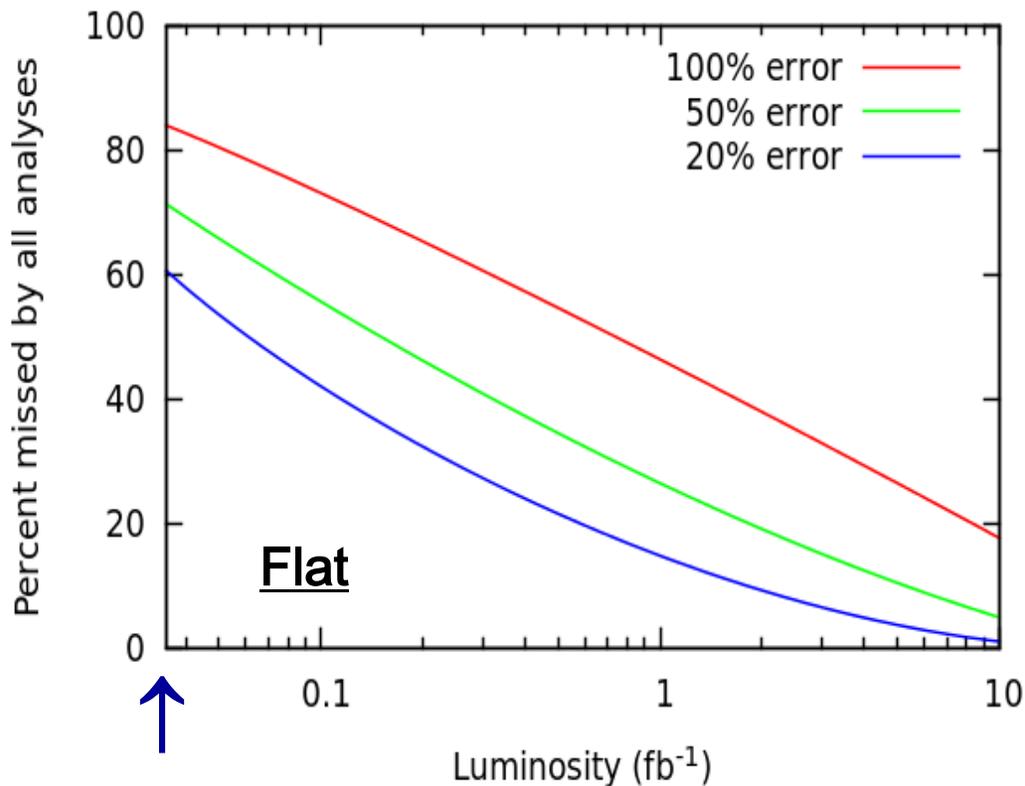
7 TeV



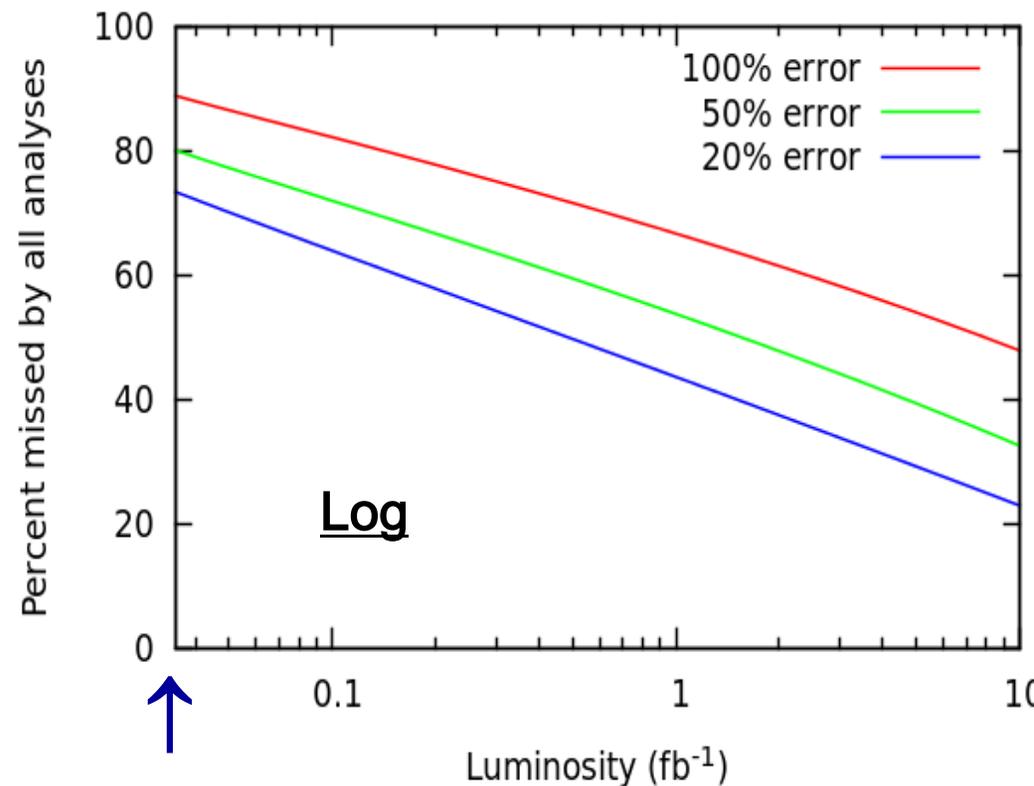
# How good is the pMSSM coverage @ 7 TeV as the luminosity evolves ??

The coverage is quite good for both model sets !

Flat priors



Log priors



# ATLAS pMSSM Model Coverage\*

RIGHT NOW for  $\sim 35 \text{ pb}^{-1}$  @ 7 TeV

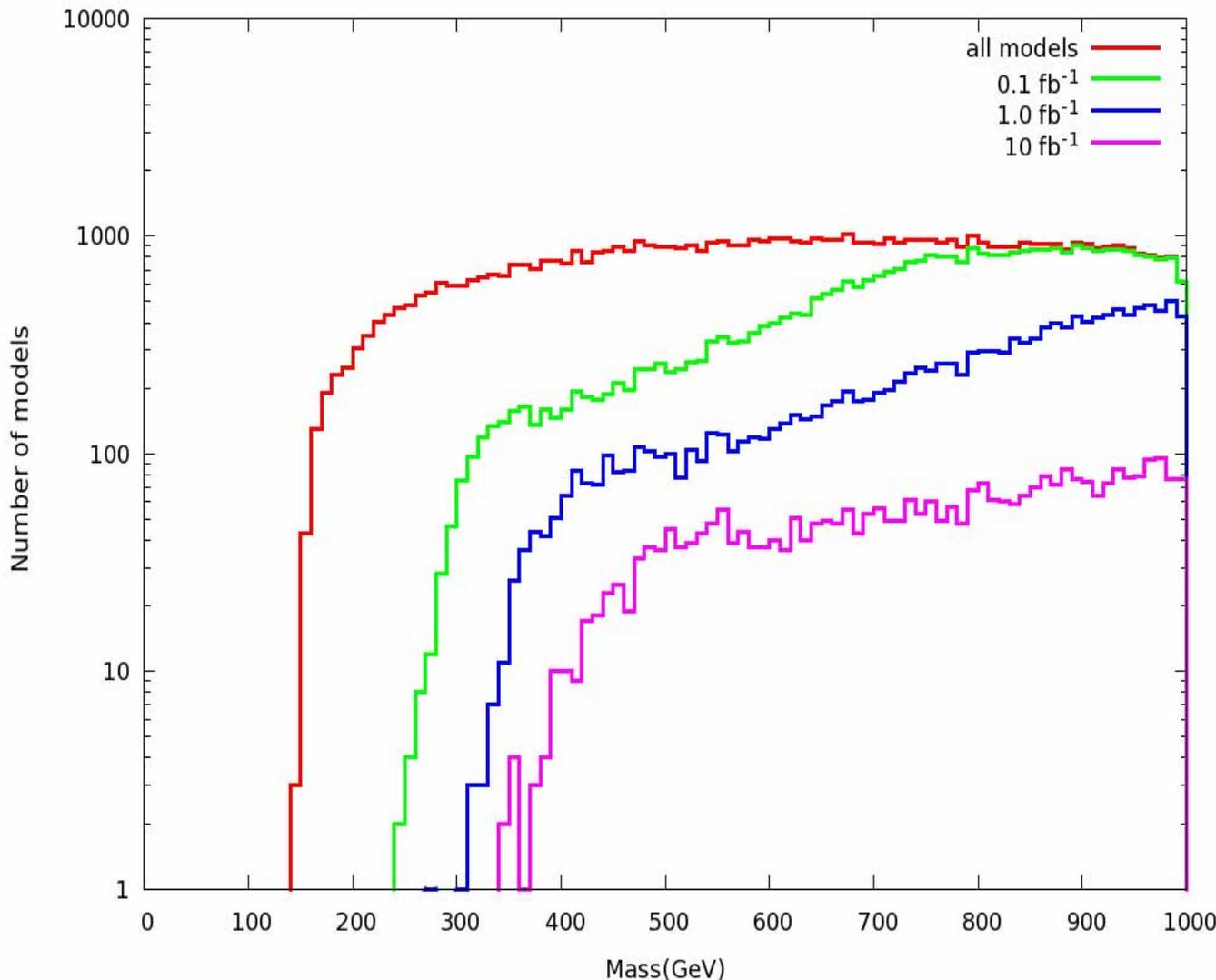
|                                |             |            |            |
|--------------------------------|-------------|------------|------------|
| <u><math>\delta B</math></u> : | <u>100%</u> | <u>50%</u> | <u>20%</u> |
| FLAT:                          | 16%         | 29%        | 39%        |
| LOG:                           | 11%         | 20%        | 27%        |

**Wow!** This is actually quite impressive as these LHC SUSY searches are **just beginning** !!!

\* Fraction of models that **SHOULD** have been found but weren't if all ATLAS analyses were done as stated

- **What do LHC searches do?** If nothing is found they just **'remove'** models from the **'allowed'** set until few remain

g Mass Distribution for FLAT models failed for 50% error



LHC SUSY searches are preferentially sensitive to squarks & gluinos

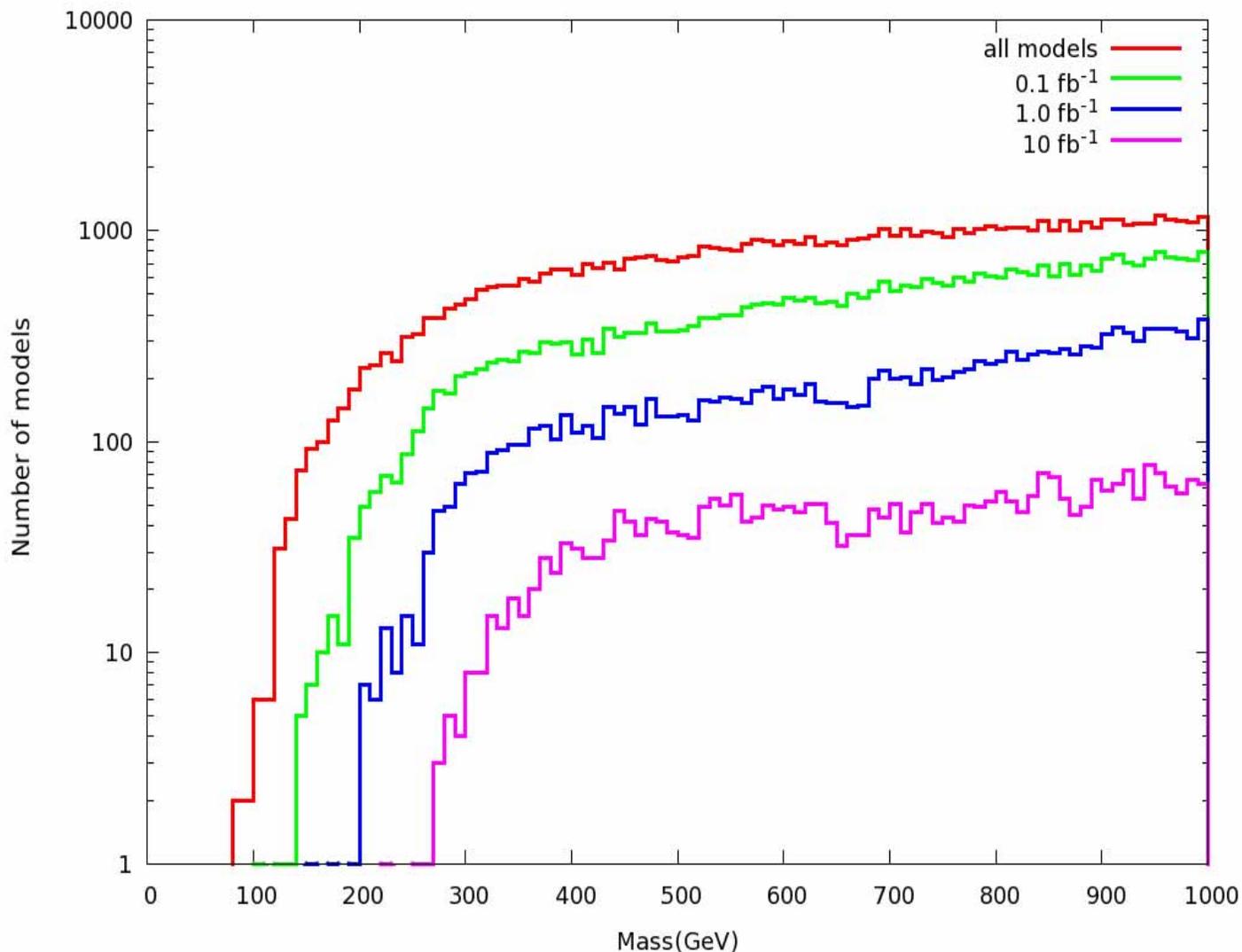
Here we see the gluino population **falling** as models are 'removed' ..

More models w/ lighter gluinos are removed

What about the other sparticles?

For the **other colored sparticles** we would expect to see the same effect as there is nothing special about gluinos.

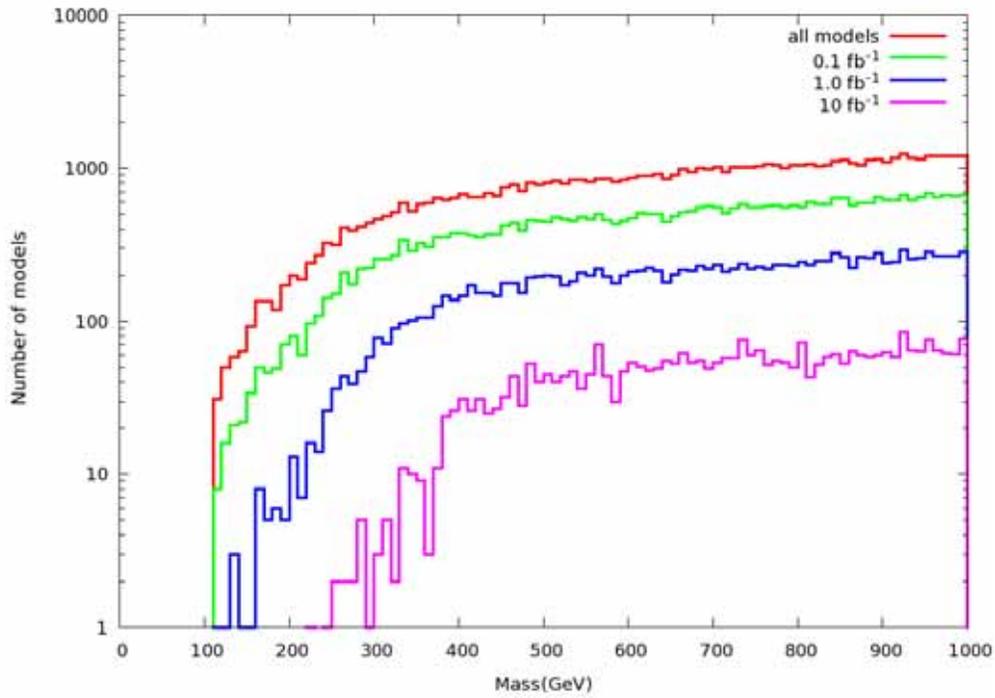
$u_L$  Mass Distribution for FLAT models failed for 50% error



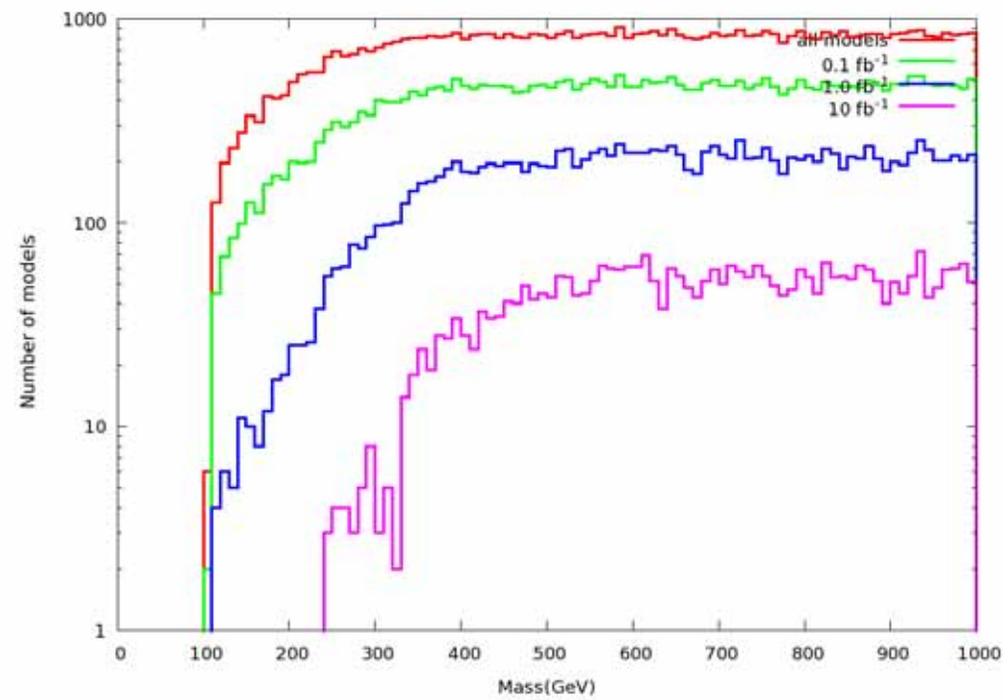
For squarks we also see the depletion of **lighter states** as the LHC lumi goes up.

But what about **color singlets** at they are not (much) directly produced at the LHC & are most relevant for the LC ?

$e_L$  Mass Distribution for FLAT models failed for 50% error

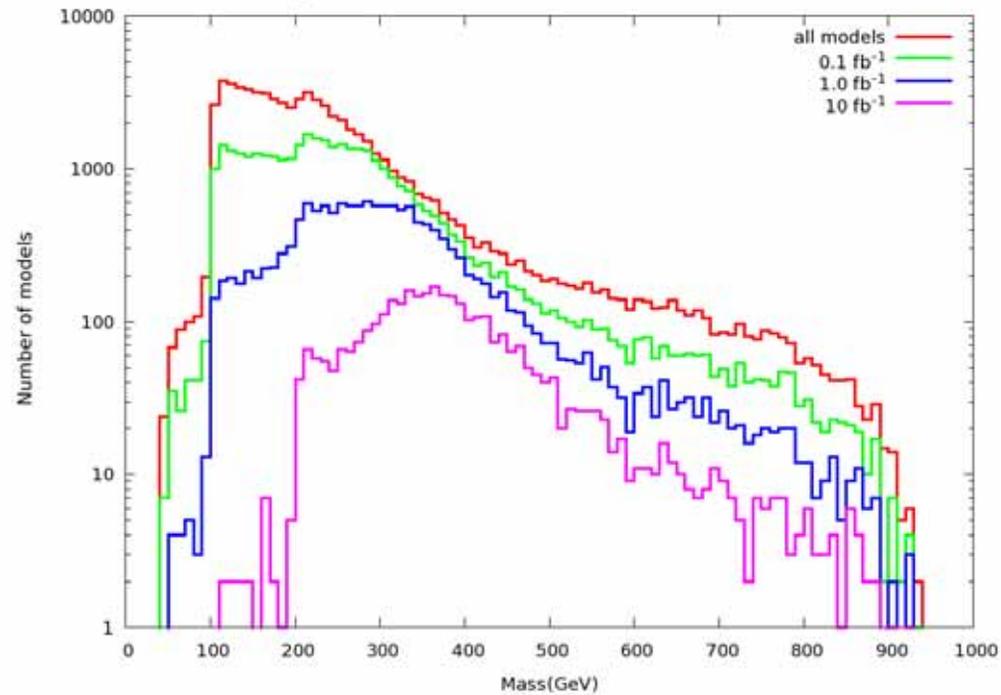


$e_R$  Mass Distribution for FLAT models failed for 50% error

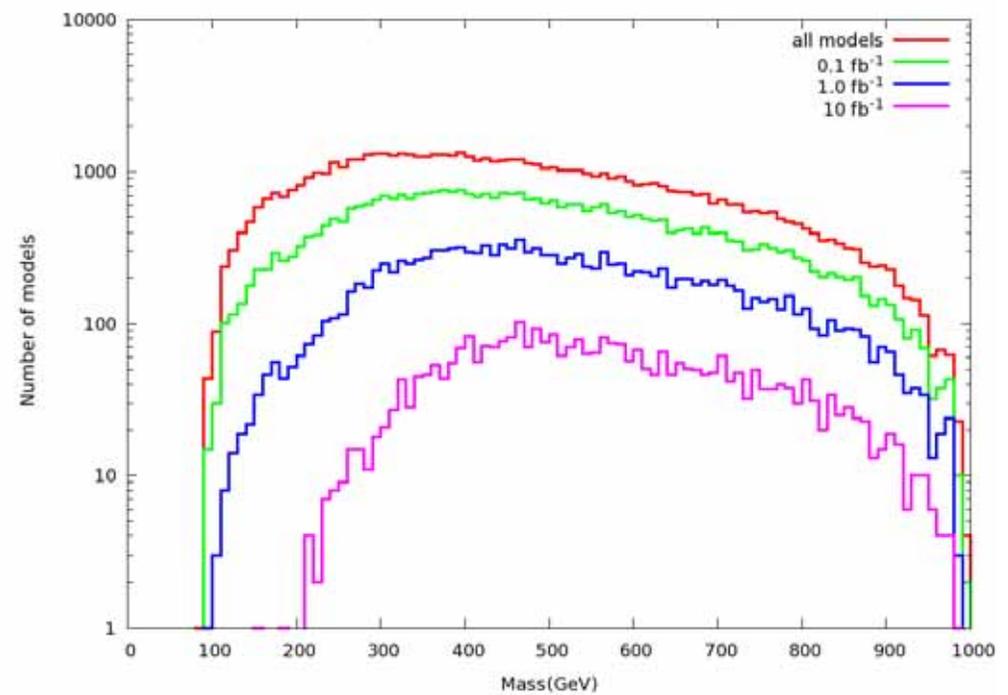


FLAT

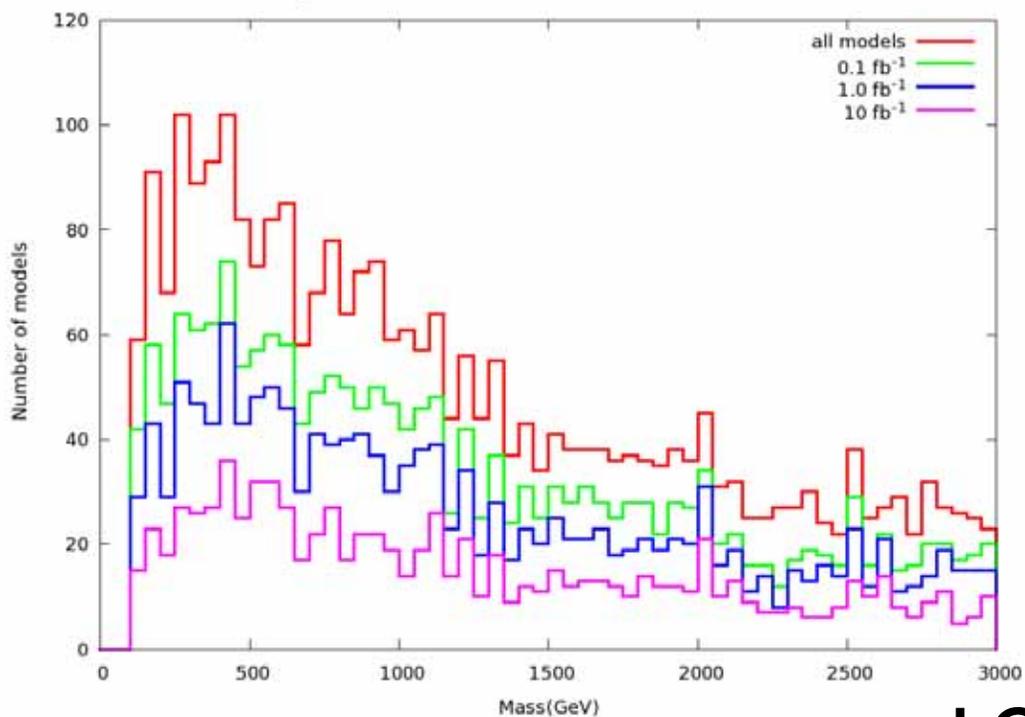
$\chi_1^+$  Mass Distribution for FLAT models failed for 50% error



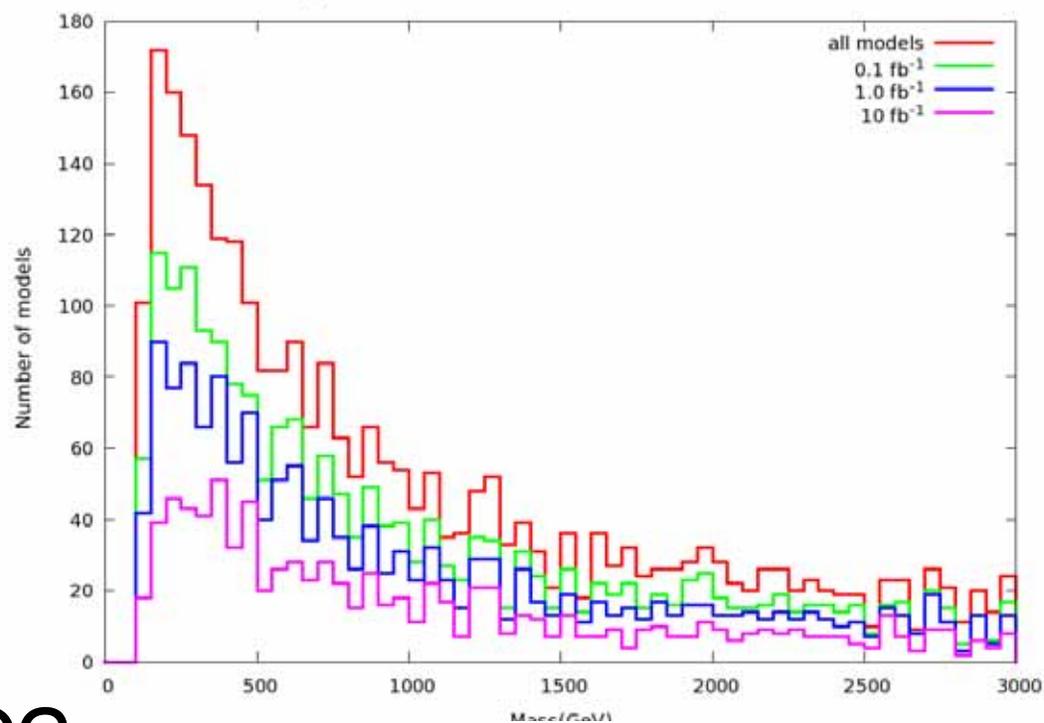
$\tau_1$  Mass Distribution for FLAT models failed for 50% error



$e_L$  Mass Distribution for LOG models failed for 50% error

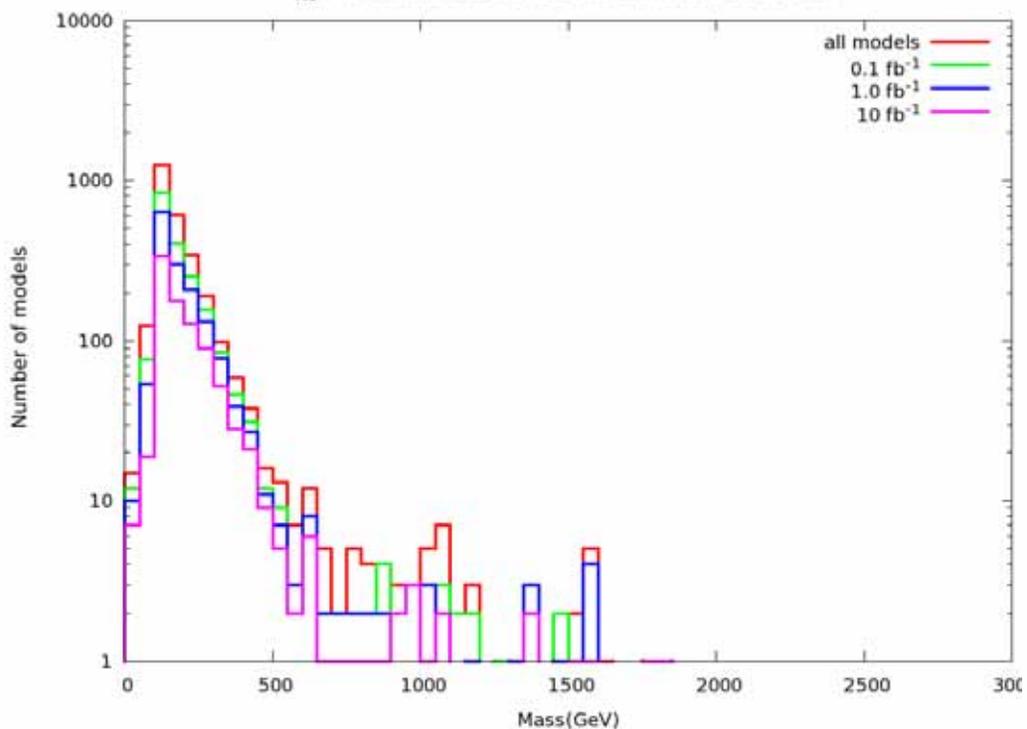


$e_R$  Mass Distribution for LOG models failed for 50% error

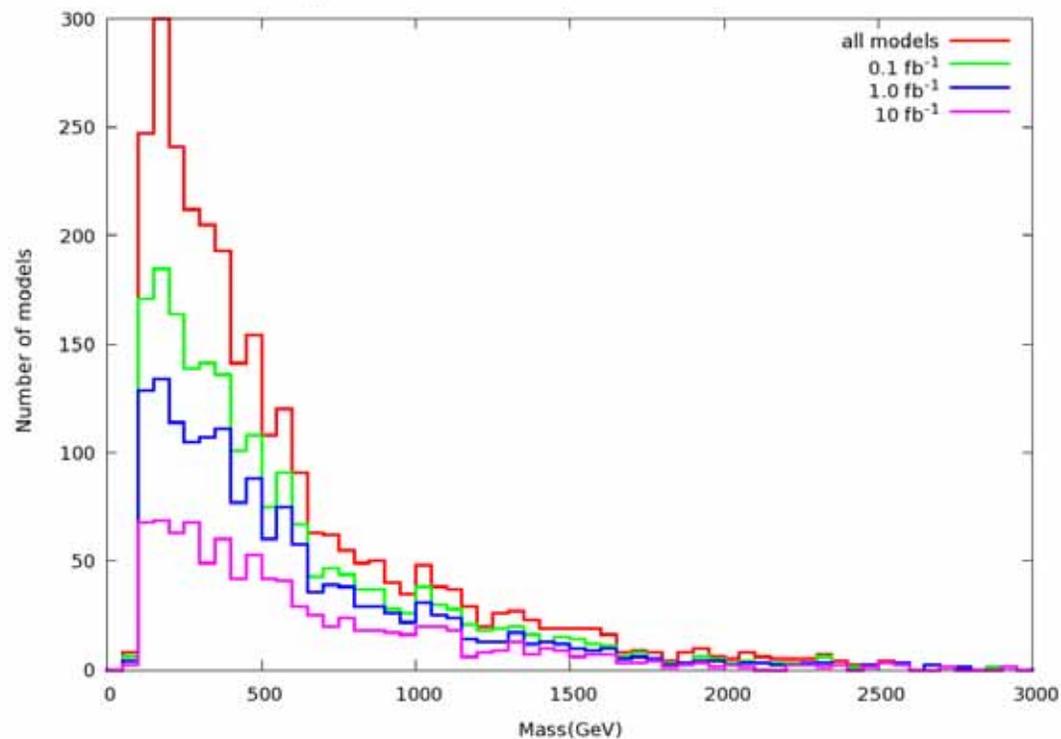


LOG

$\chi_1^+$  Mass Distribution for LOG models failed for 50% error



$\tau_1$  Mass Distribution for LOG models failed for 50% error



| Sparticle            | $\sqrt{s} = 500 \text{ GeV}$ |      | $\sqrt{s} = 1 \text{ TeV}$ |      |
|----------------------|------------------------------|------|----------------------------|------|
|                      | Flat                         | Log  | Flat                       | Log  |
| $\tilde{e}_L$        | 107                          | 101  | 3052                       | 347  |
| $\tilde{e}_R$        | 260                          | 209  | 3938                       | 565  |
| $\tilde{\tau}_1$     | 730                          | 381  | 7431                       | 869  |
| $\tilde{\tau}_2$     | 30                           | 36   | 1288                       | 207  |
| $\tilde{\nu}_e$      | 151                          | 117  | 3168                       | 356  |
| $\tilde{\nu}_\tau$   | 386                          | 236  | 4366                       | 553  |
| $\tilde{\chi}_1^0$   | 5487                         | 1312 | 14,510                     | 1539 |
| $\tilde{\chi}_2^0$   | 2738                         | 1035 | 10,714                     | 1395 |
| $\tilde{\chi}_3^0$   | 429                          | 352  | 5667                       | 903  |
| $\tilde{\chi}_4^0$   | 10                           | 18   | 1267                       | 202  |
| $\tilde{\chi}_1^\pm$ | 4856                         | 1208 | 13,561                     | 1495 |
| $\tilde{\chi}_2^\pm$ | 94                           | 54   | 3412                       | 456  |
| $\tilde{g}$          | 0                            | 0    | 1088                       | 65   |
| $\tilde{d}_L$        | 35                           | 11   | 2459                       | 117  |
| $\tilde{d}_R$        | 220                          | 96   | 3630                       | 526  |
| $\tilde{u}_L$        | 52                           | 16   | 2545                       | 123  |
| $\tilde{u}_R$        | 124                          | 64   | 3581                       | 273  |
| $\tilde{b}_1$        | 289                          | 75   | 5553                       | 590  |
| $\tilde{b}_2$        | 1                            | 0    | 409                        | 21   |
| $\tilde{t}_1$        | 93                           | 9    | 3727                       | 217  |
| $\tilde{t}_2$        | 0                            | 0    | 2                          | 0    |

## LC Implications

← In the set of 14623 (1546) **FLAT(LOG)** models **NOT** found at 7 TeV w/ 1 fb<sup>-1</sup> &  $\delta B=50\%$  we find there are....

This doesn't look too bad for 500 GeV or 1 TeV !

**But →**

| Sparticle            | $\sqrt{s} = 500 \text{ GeV}$ |     | $\sqrt{s} = 1 \text{ TeV}$ |     |
|----------------------|------------------------------|-----|----------------------------|-----|
|                      | Flat                         | Log | Flat                       | Log |
| $\tilde{e}_L$        | 0                            | 37  | 63                         | 142 |
| $\tilde{e}_R$        | 0                            | 72  | 53                         | 223 |
| $\tilde{\tau}_1$     | 2                            | 142 | 165                        | 338 |
| $\tilde{\tau}_2$     | 0                            | 11  | 9                          | 69  |
| $\tilde{\nu}_e$      | 0                            | 42  | 64                         | 146 |
| $\tilde{\nu}_\tau$   | 0                            | 85  | 81                         | 236 |
| $\tilde{\chi}_1^0$   | 26                           | 507 | 587                        | 626 |
| $\tilde{\chi}_2^0$   | 4                            | 397 | 352                        | 557 |
| $\tilde{\chi}_3^0$   | 0                            | 136 | 57                         | 357 |
| $\tilde{\chi}_4^0$   | 0                            | 5   | 5                          | 66  |
| $\tilde{\chi}_1^\pm$ | 25                           | 467 | 505                        | 608 |
| $\tilde{\chi}_2^\pm$ | 0                            | 17  | 16                         | 170 |
| $\tilde{g}$          | 0                            | 0   | 27                         | 5   |
| $\tilde{d}_L$        | 0                            | 3   | 73                         | 24  |
| $\tilde{d}_R$        | 1                            | 18  | 63                         | 157 |
| $\tilde{u}_L$        | 0                            | 5   | 81                         | 24  |
| $\tilde{u}_R$        | 0                            | 14  | 86                         | 79  |
| $\tilde{b}_1$        | 0                            | 20  | 103                        | 189 |
| $\tilde{b}_2$        | 0                            | 0   | 3                          | 4   |
| $\tilde{t}_1$        | 1                            | 2   | 94                         | 58  |
| $\tilde{t}_2$        | 0                            | 0   | 0                          | 0   |

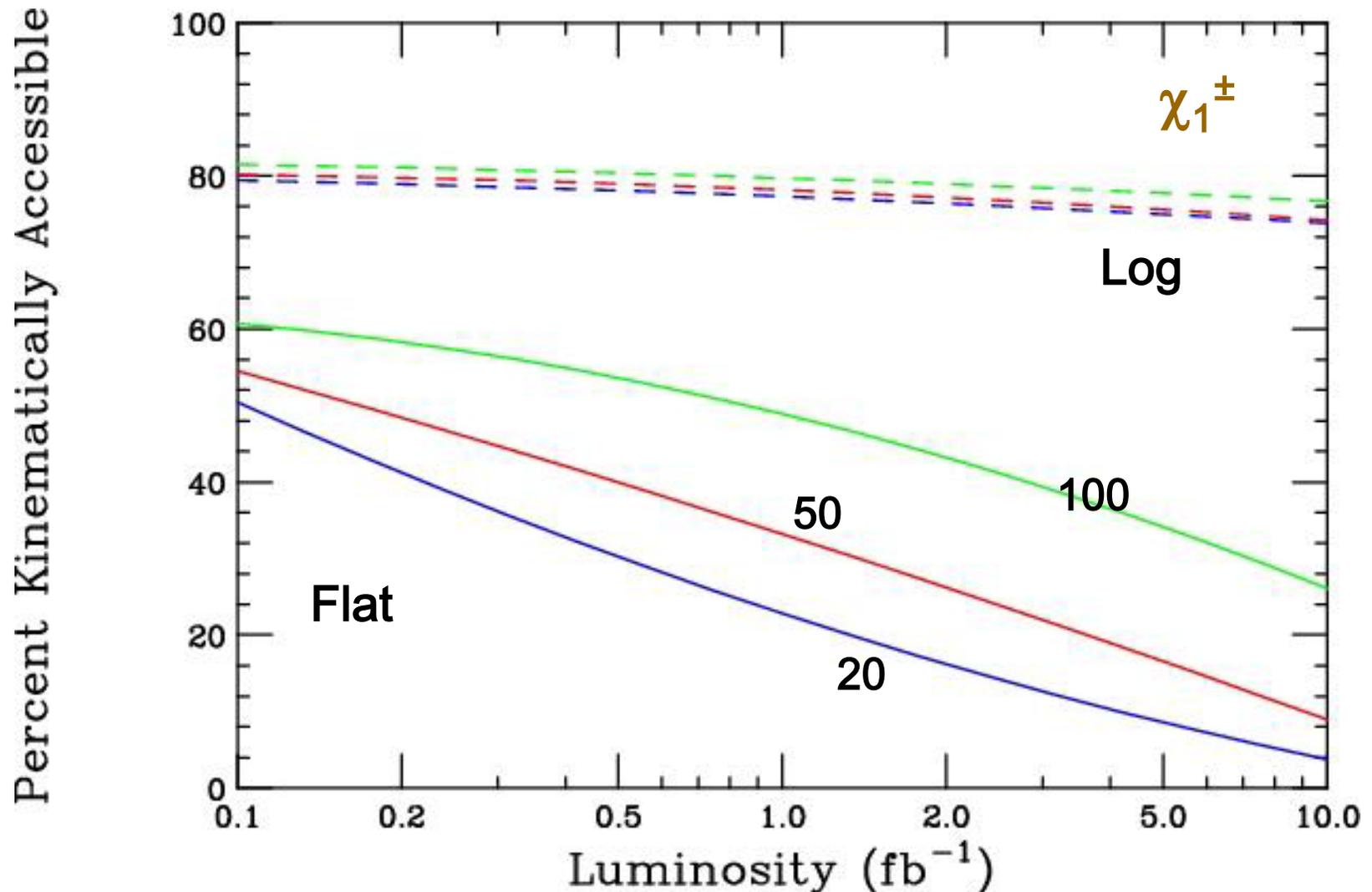
In the set of 672(663)  
**FLAT(LOG)** models  
**NOT** found @ 7 TeV  
w/ 10 fb<sup>-1</sup> and  $\delta B=20\%$   
there are **a lot fewer** !

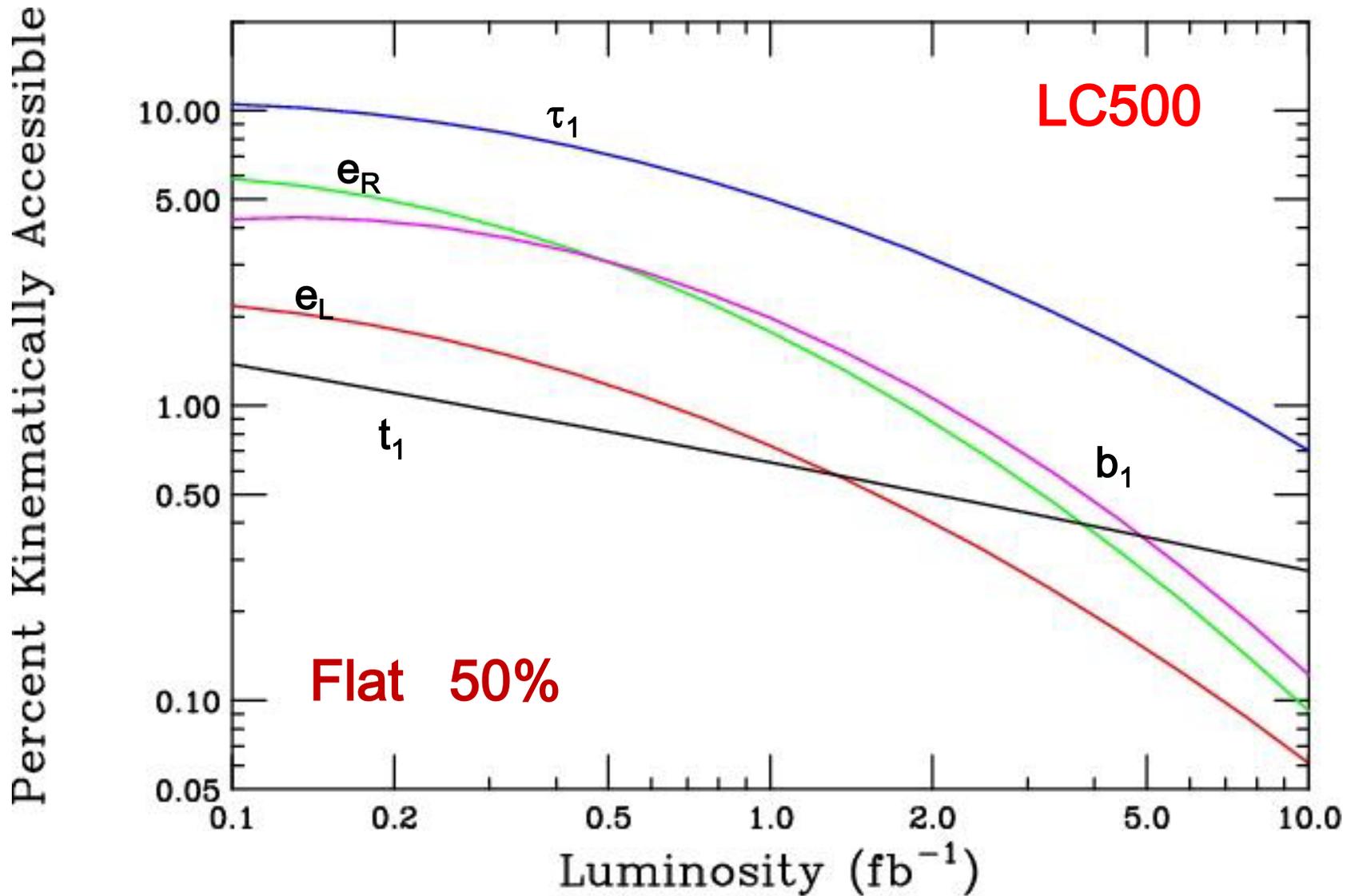
← Not a whole lot in the  
**FLAT** case at 500 GeV !  
...so the **SUSY @ the LC**  
**window depends** where **we**  
**end up & the spectrum**  
**shape details**

However, a 1TeV LC is  
vastly BETTER !

- Of course what's **REALLY** important is the fraction of the 'undiscovered' models which have kinematically accessible sparticles at the LC...

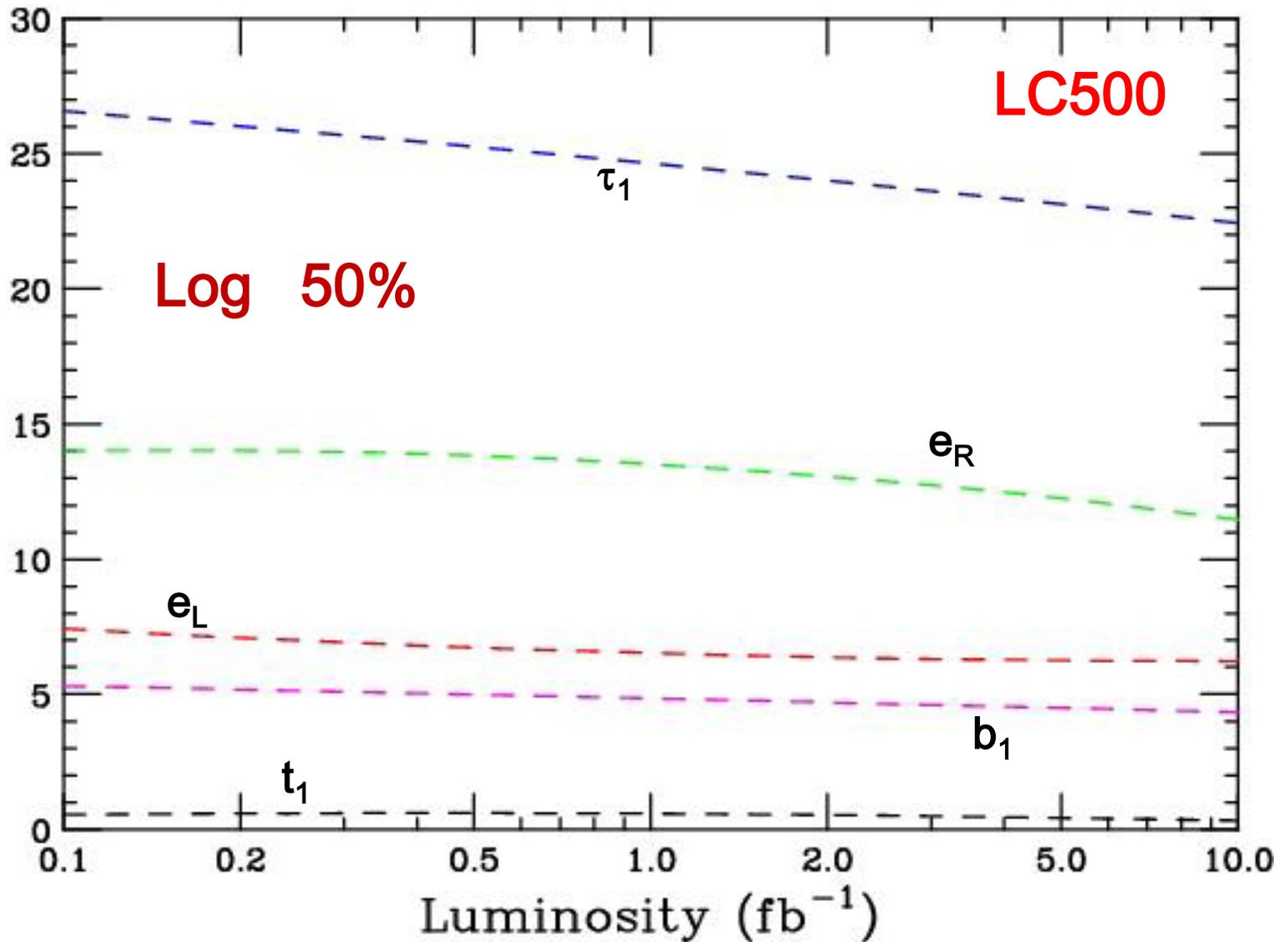
LC500





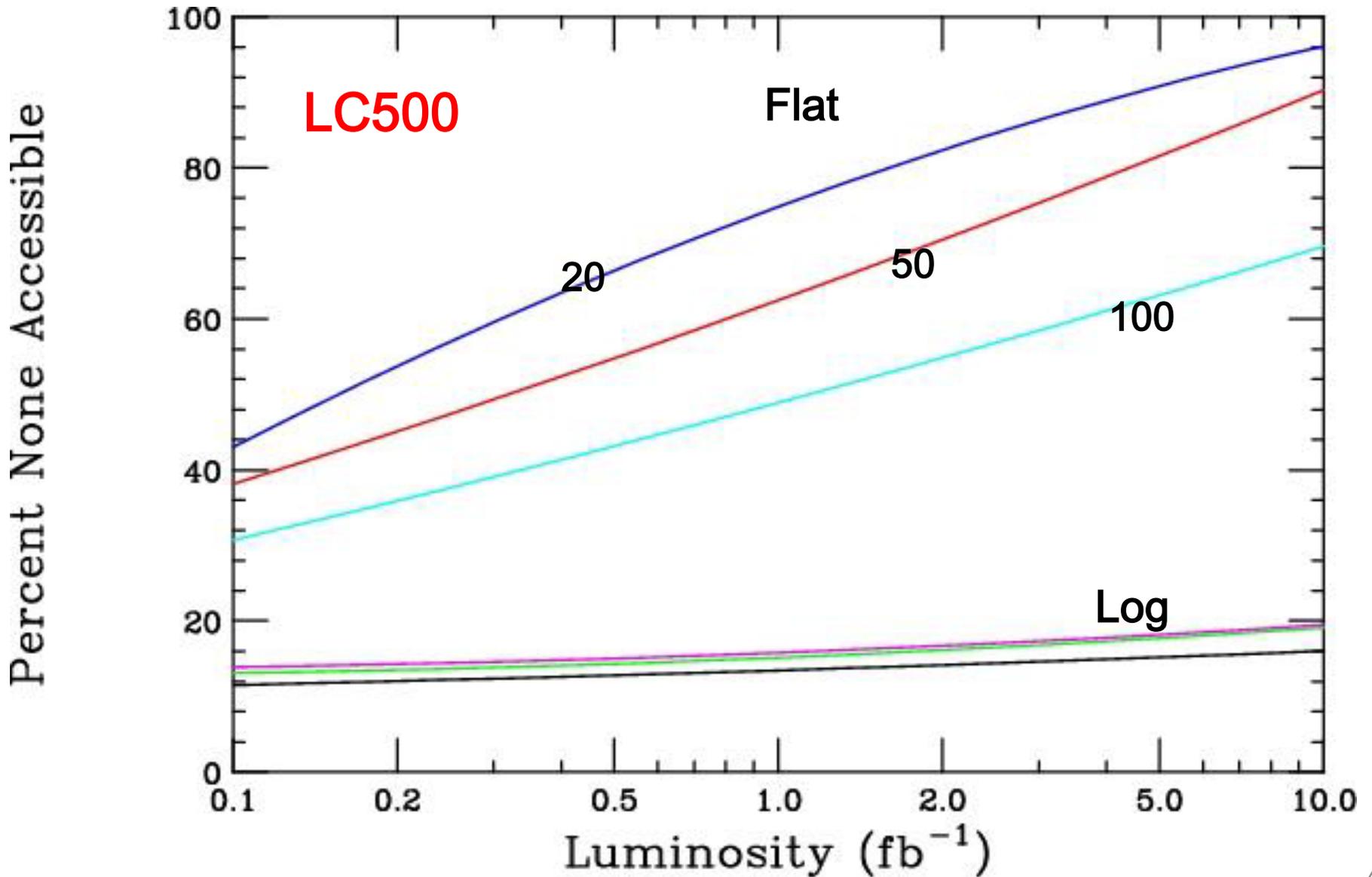
As the lumi increases the fraction of flat prior models with **LC500** observable sparticles becomes very small  $<1\%$ . However...

Percent Kinematically Accessible

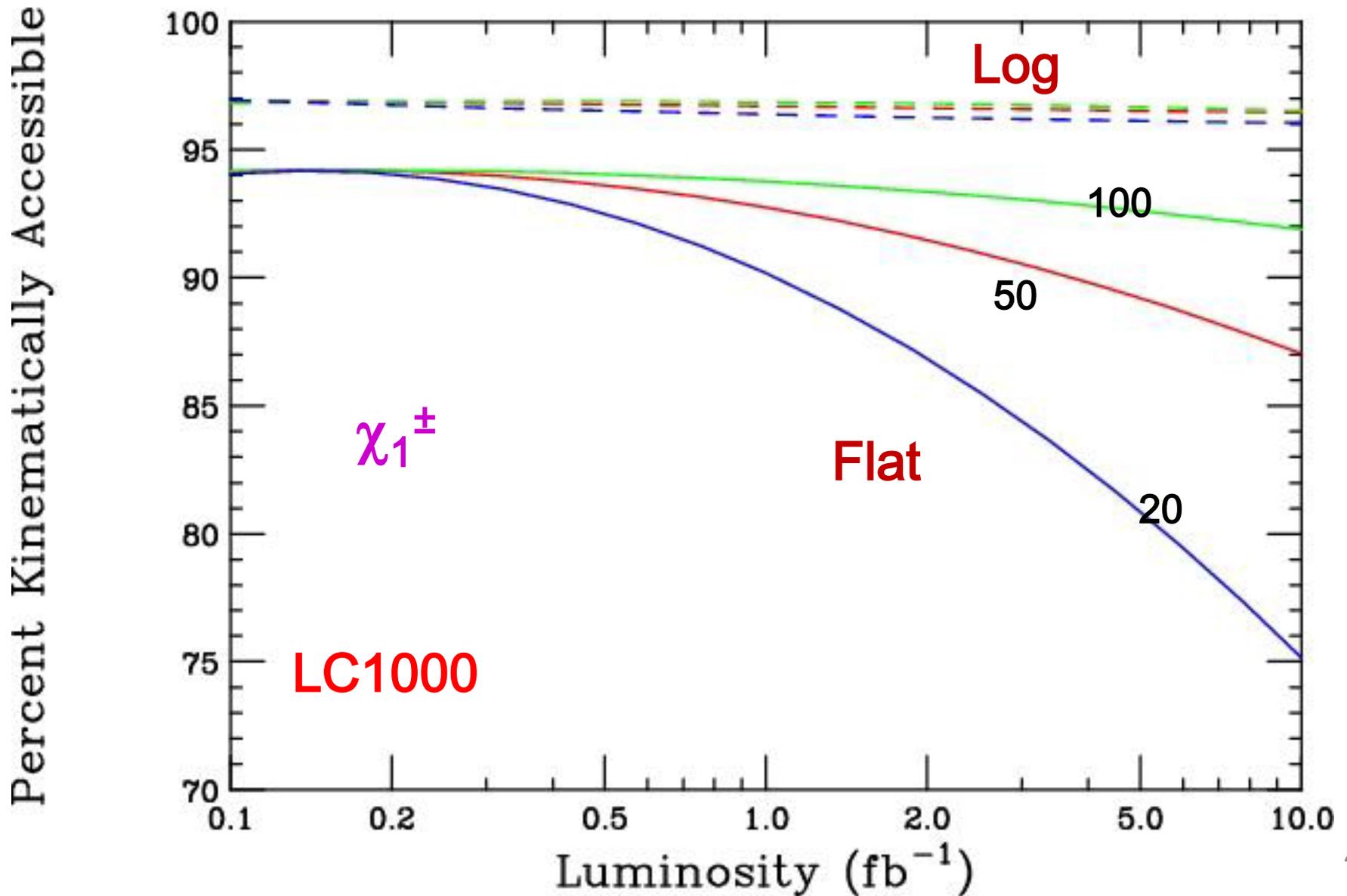


As the lumi increases the fraction of log prior models with LC500 observable particles declines quite slowly !

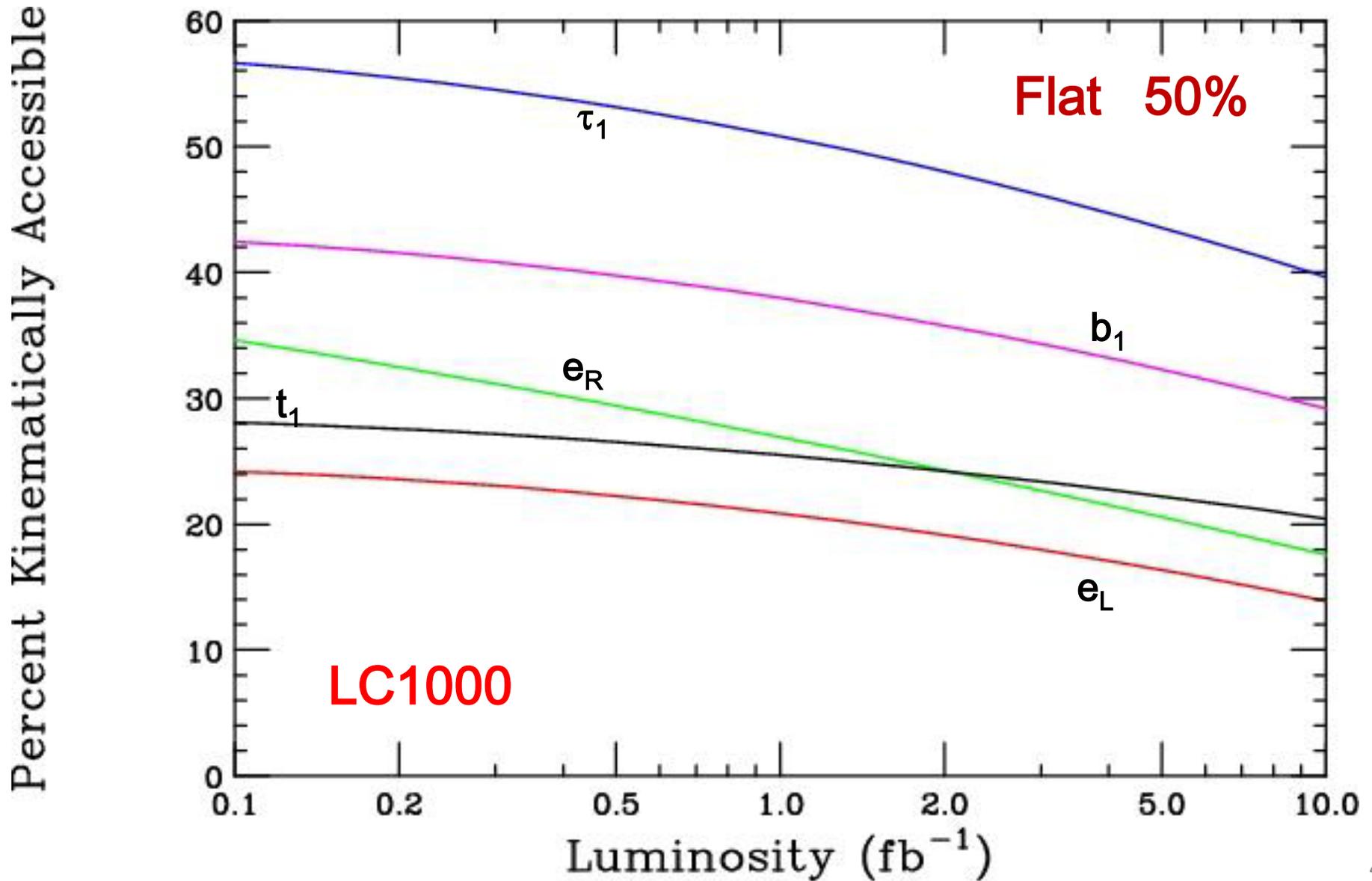
The fraction of models with **NO** sparticles accessible at **LC500** grows rapidly for the **flats** but almost not at all for the **logs**

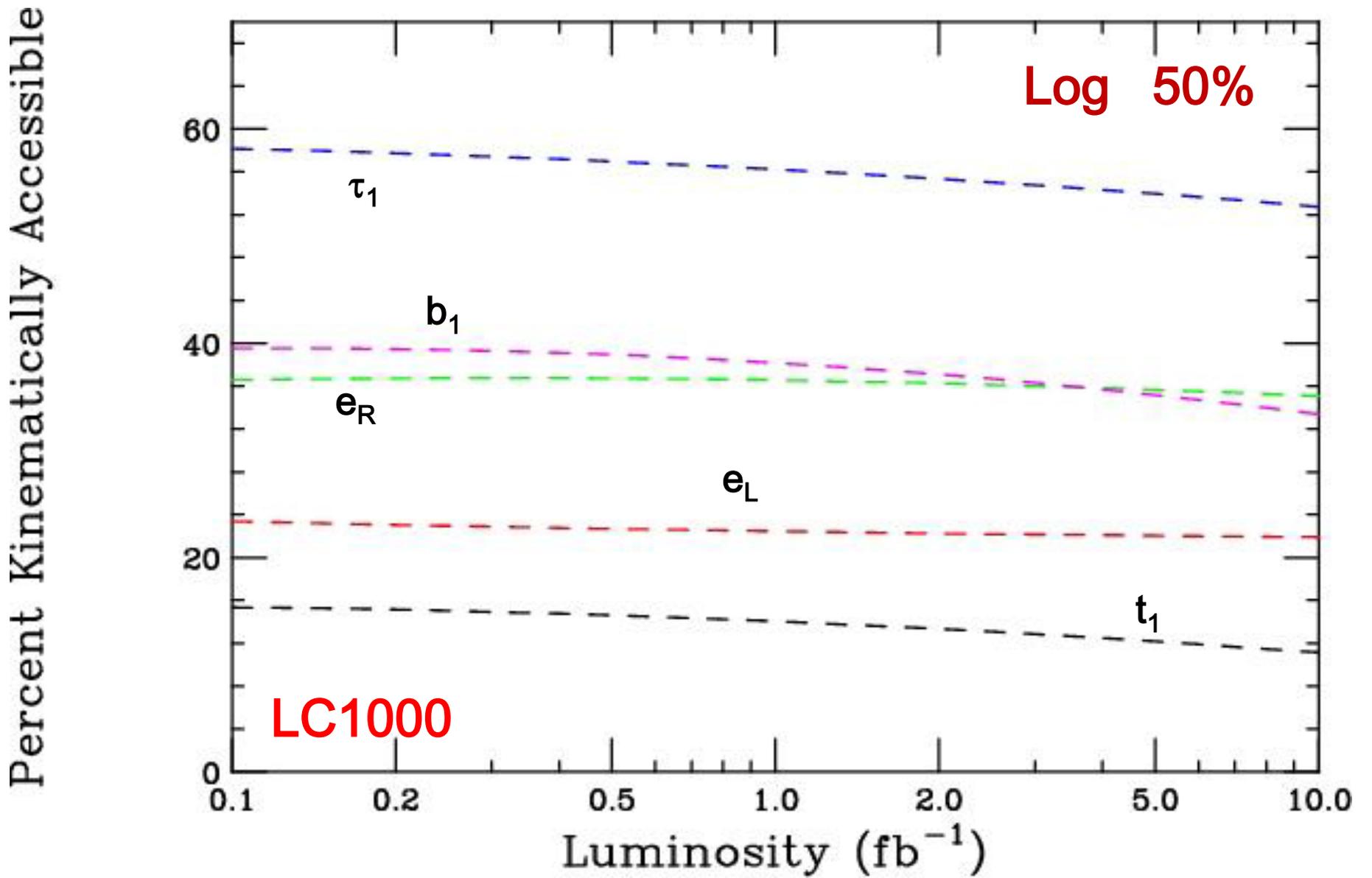


Things do change a lot at **LC1000**... a very large fraction of models have accessible charginos.



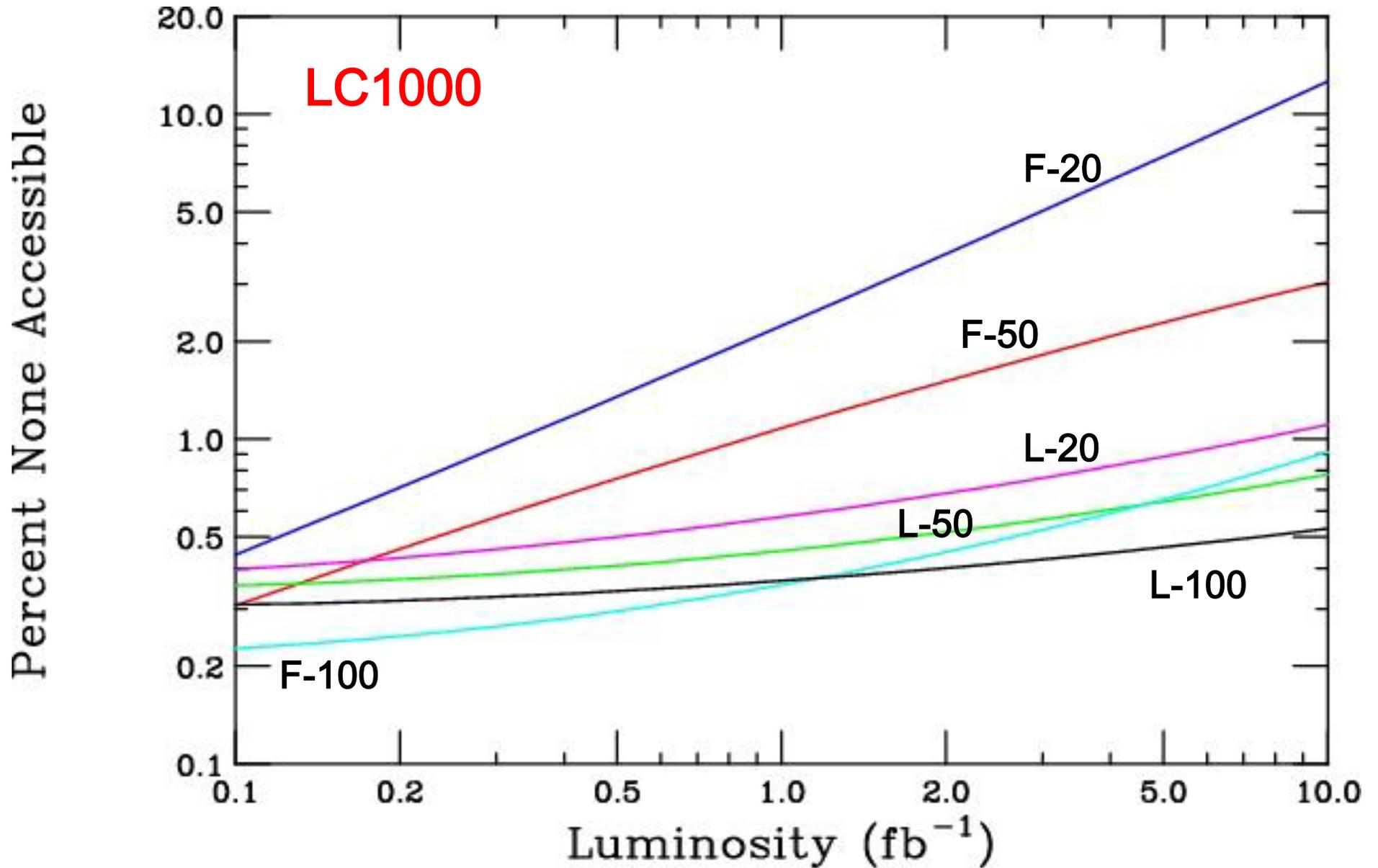
# Other sparticles are also reasonably accessible at LC1000





For the Log case many sparticles are accessible at LC1000

The fraction of models w/o any SUSY accessible is small



- If nothing is seen, just how visible SUSY may or may not be at LC500 will depend on 2 factors that we **presently don't know**:
    - (i) What luminosity/background systematics will be achieved during the 7 TeV LHC 2011-12 run?  
(We just need to wait to find out...)
    - (ii) How are the many sparticles in the SUSY spectrum distributed ? (i.e., how is SUSY broken..who can say?)
- If things are **remotely like** in the Flat prior case **LC500** is **NOT** a good place for SUSY studies

# Summary & Conclusions : I

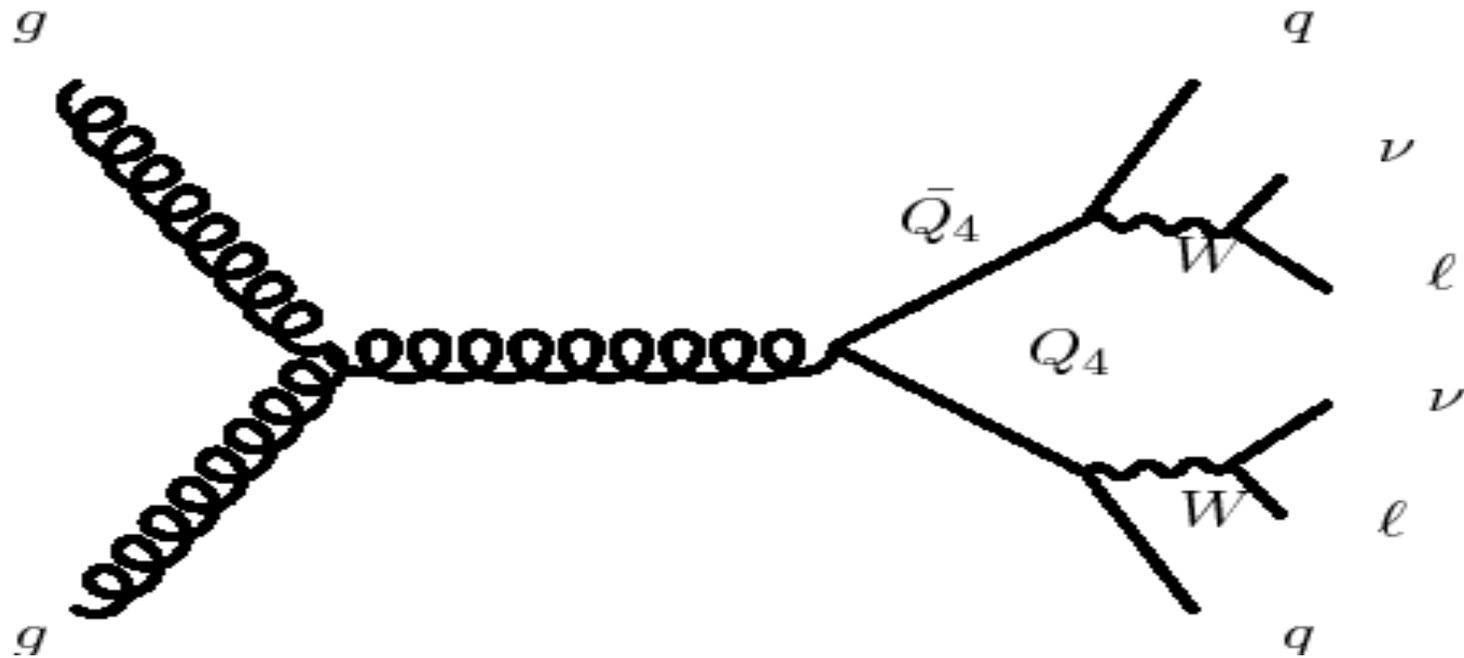
- The LHC is **already** placing limits on BSM physics beyond the reach of the Tevatron since no new signals have **yet** been seen
- **If a Z' is not seen by the time  $\sim 5 \text{ fb}^{-1}$  is accumulated then it is likely to be beyond the range of LC500 to study either because it is too heavy or too weakly coupled**
- **ADD graviton exchange visibility is not likely at LC500**
- **'Weak' constraints on some parts of a model can sometimes yield stronger constraints in another sector as in the LRM case**

# Summary & Conclusions : II

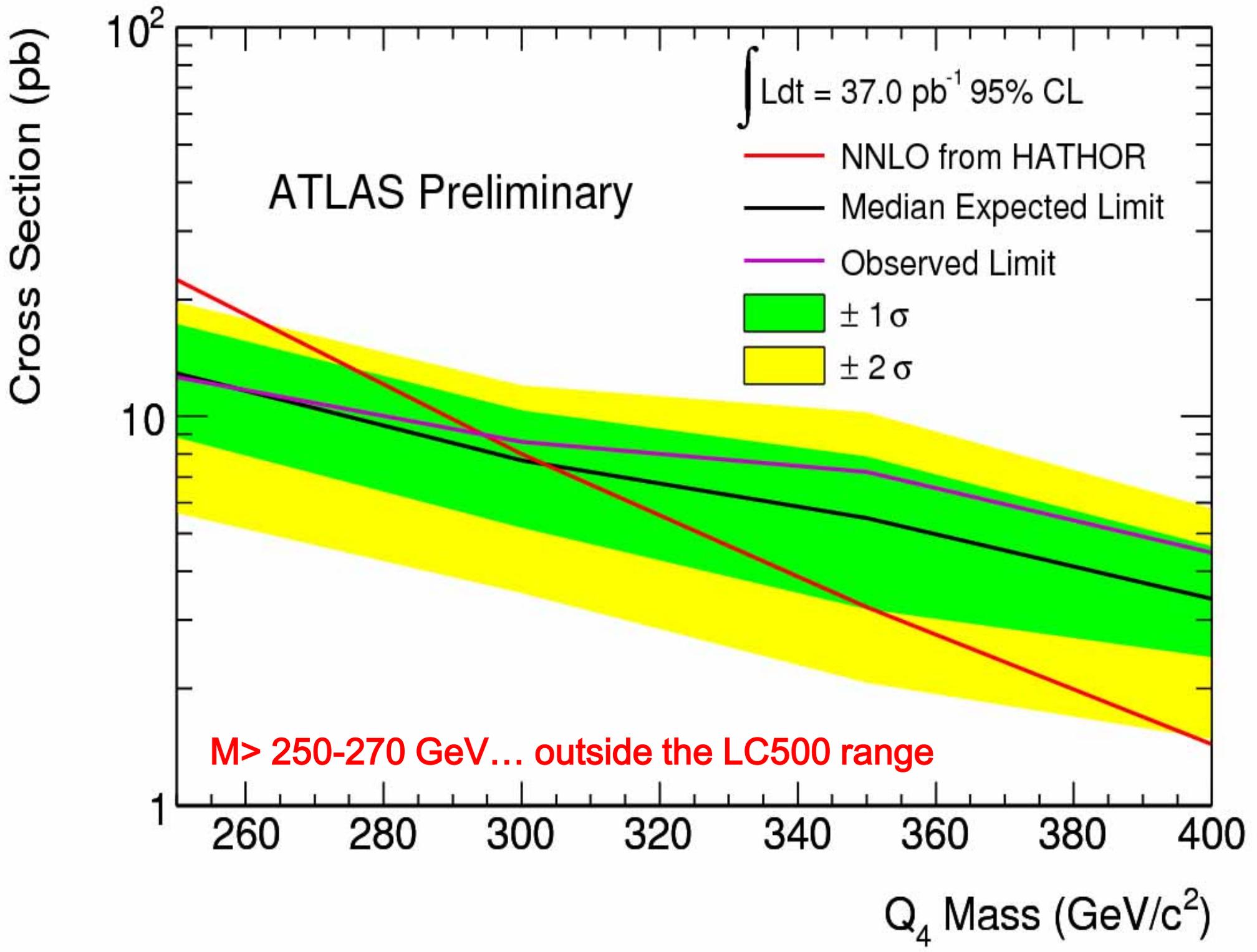
- LHC explorations of SUSY are underway & a fair bit of the **mSUGRA parameter space** is already excluded
- **With  $\sim 35 \text{ pb}^{-1}$  , a reasonable fraction of our pMSSM model space has also been 'covered' !**
- **Combined chargino + slepton studies at LC500 in mSUGRA are dead**
- **At best, LC500 looks 'iffy' for pMSSM SUSY studies BUT we don't know either the final performance of the LHC or what the sparticle spectrum looks like**
- **We're anxious to see more data & remember the caveats!**

# BACKUP SLIDES

Another possibility: 4<sup>th</sup> generation quark or, even more interestingly, an exotic vector-like quark decaying by mixing w/ ANY SM quark:

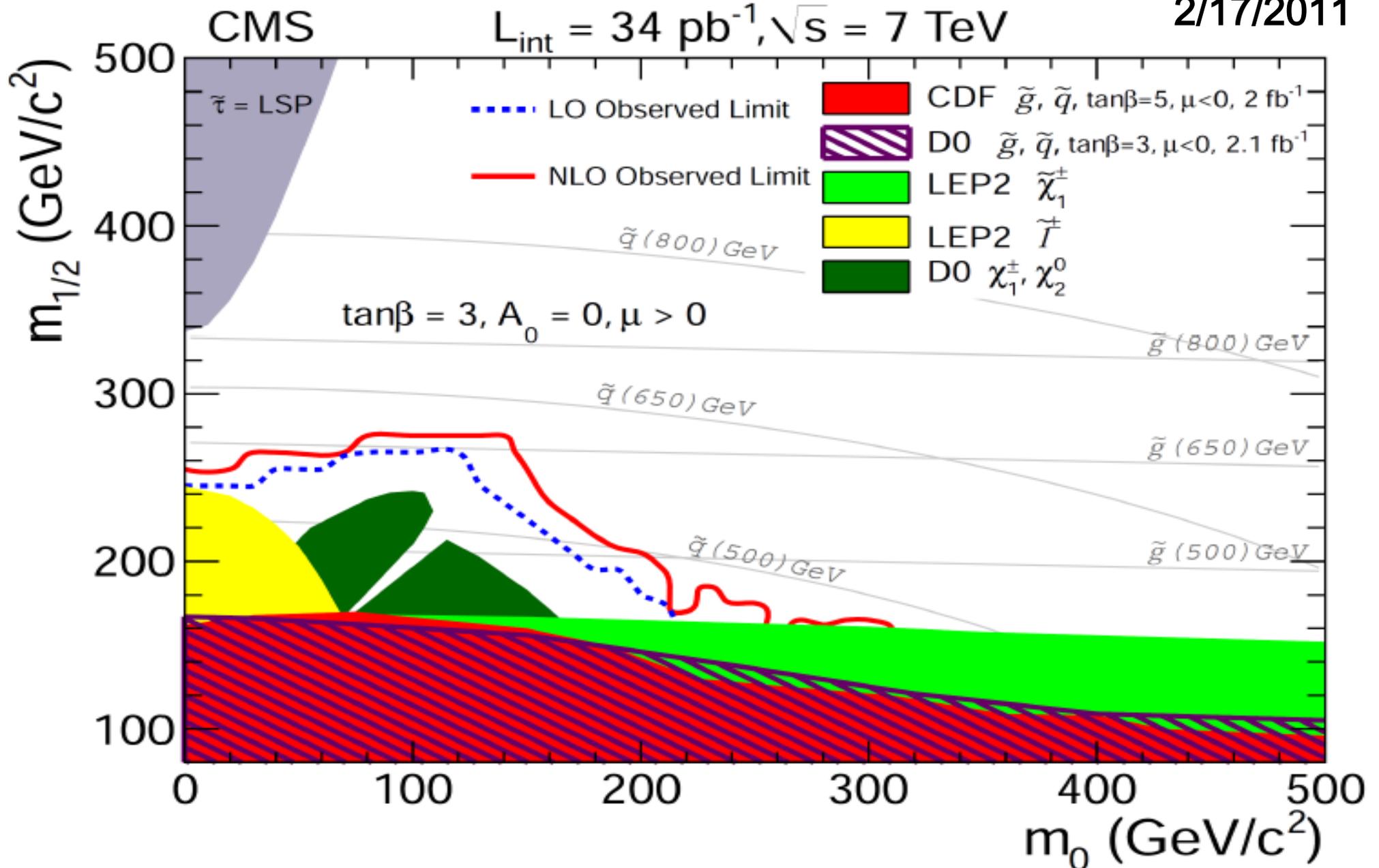


For exotics we need to use  $\sigma \cdot B_W^2$



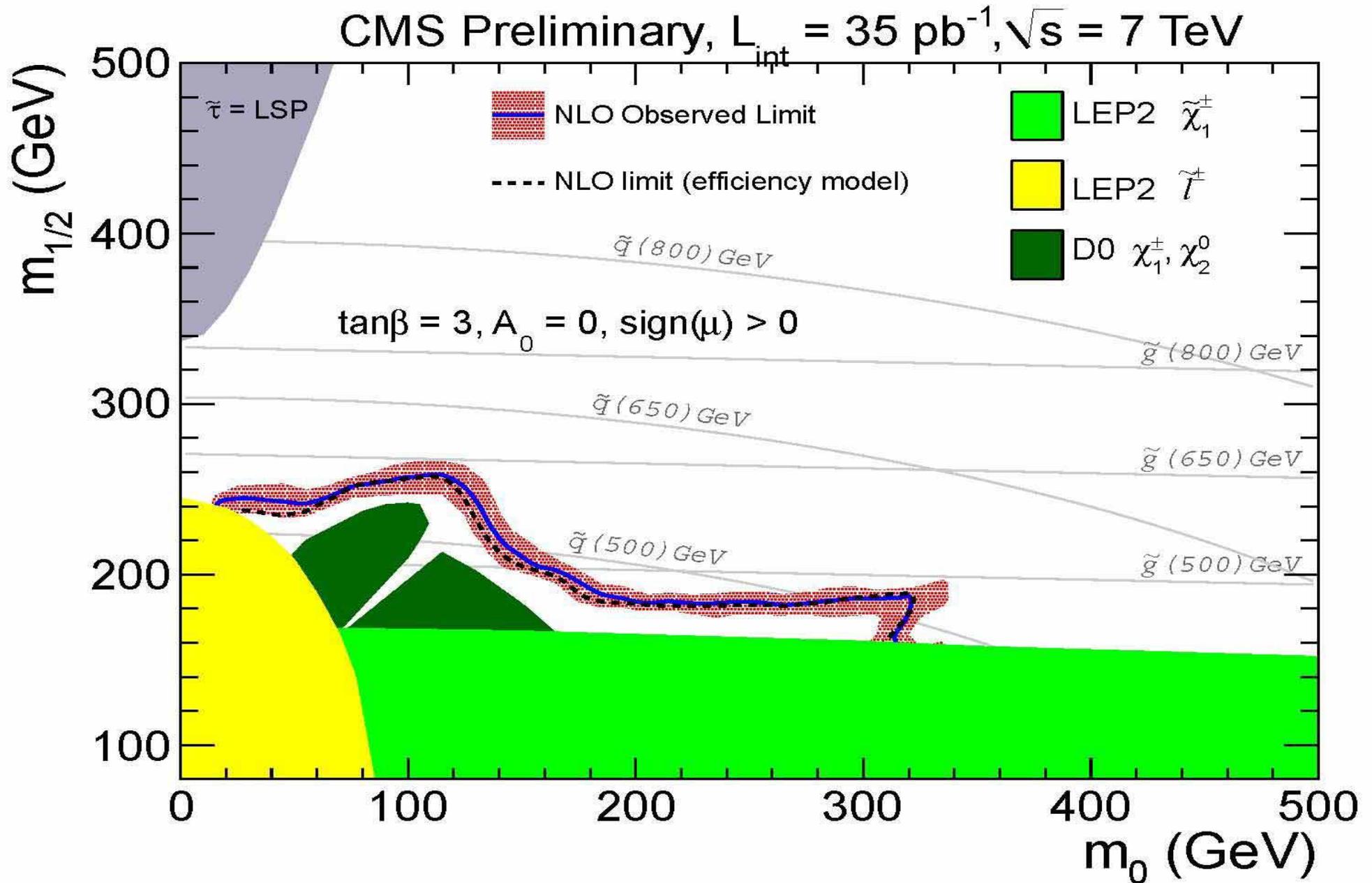
# 2jOSDL..more icing on the cake...

2/17/2011



# SSDL..even more icing on the cake...

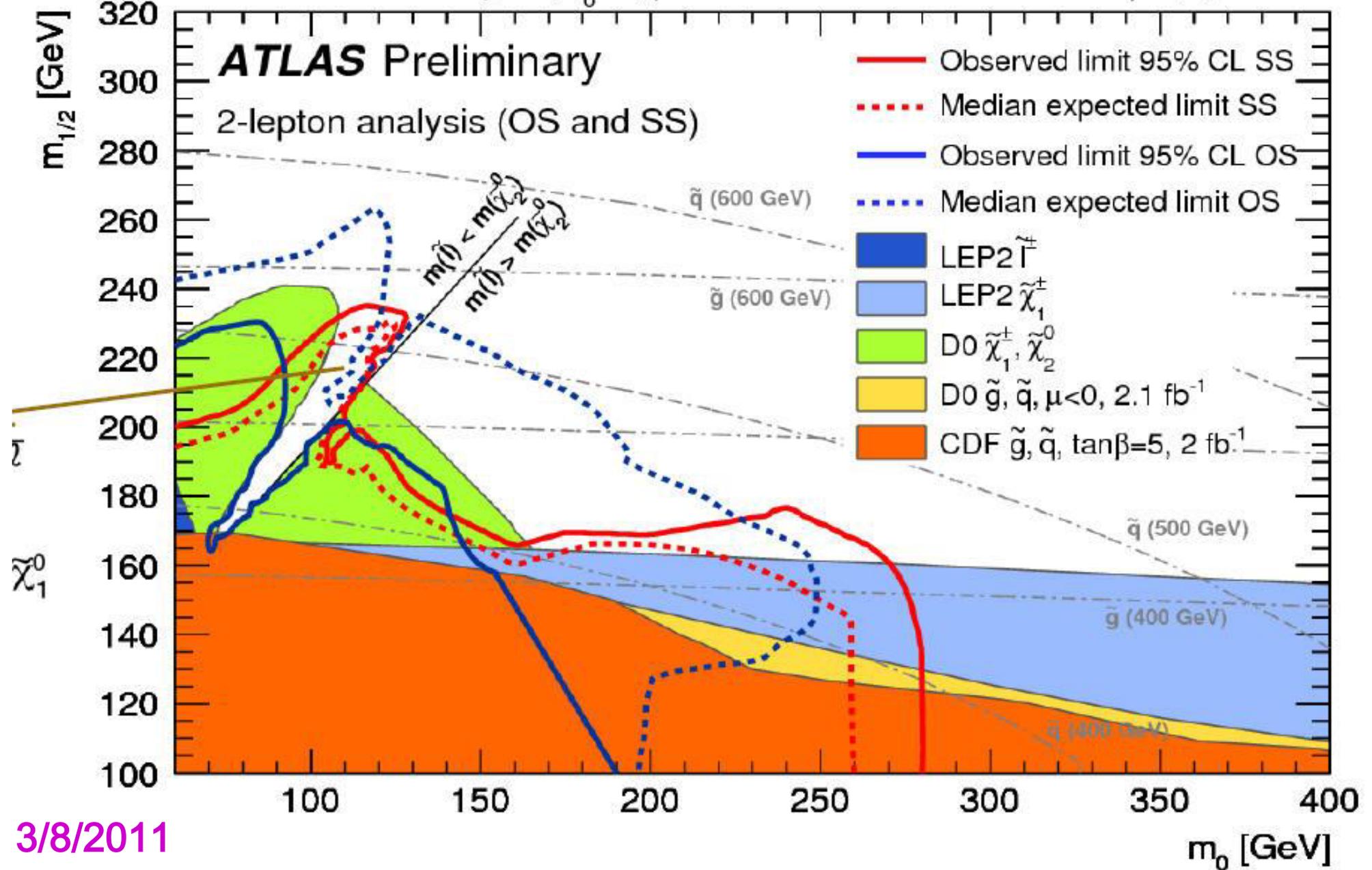
3/6/2011



..& the results keep right on coming !

MSUGRA/CMSSM:  $\tan\beta = 3, A_0 = 0, \mu > 0$

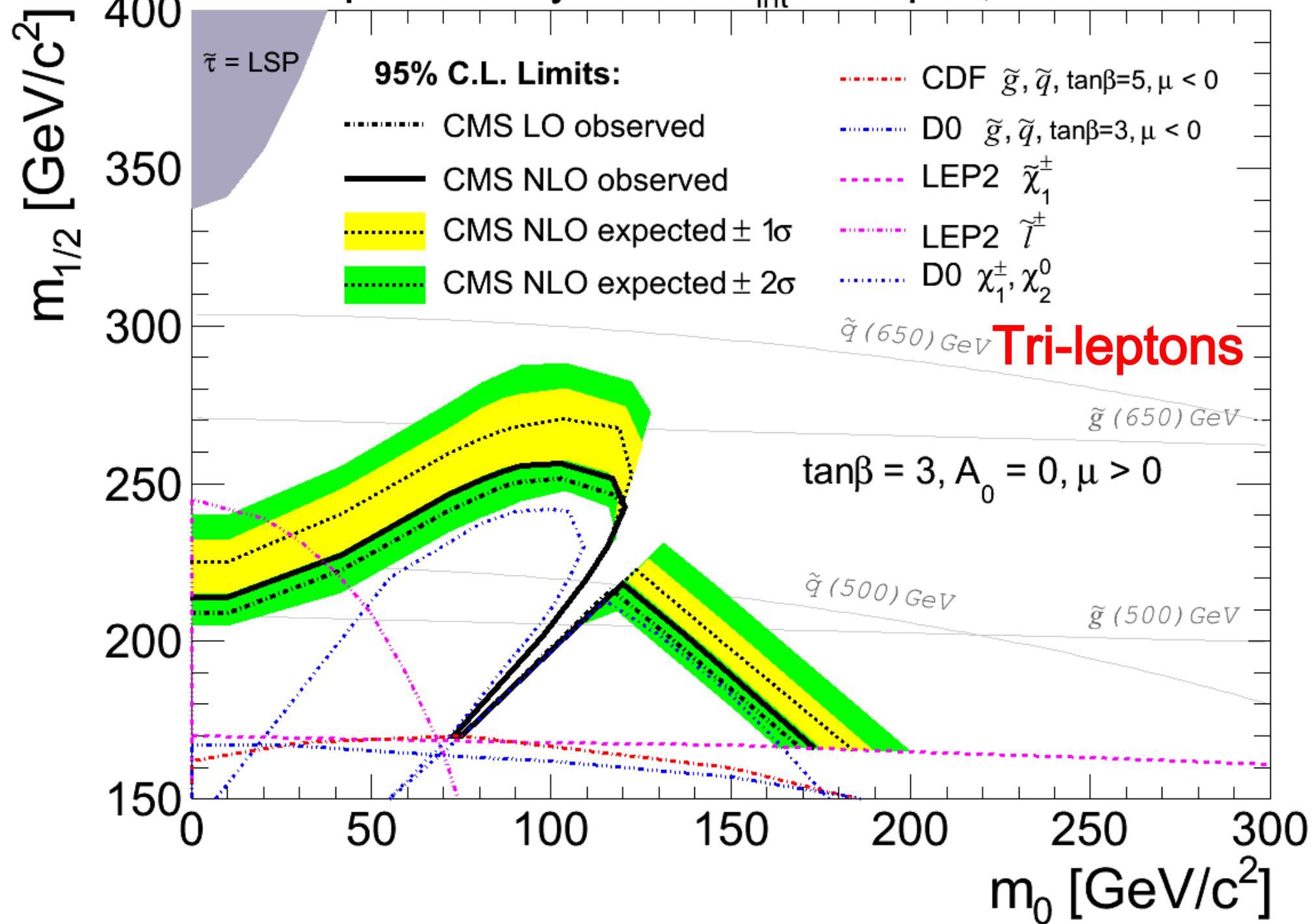
$L^{\text{int}} = 35 \text{ pb}^{-1}, \sqrt{s} = 7 \text{ TeV}$



3/8/2011

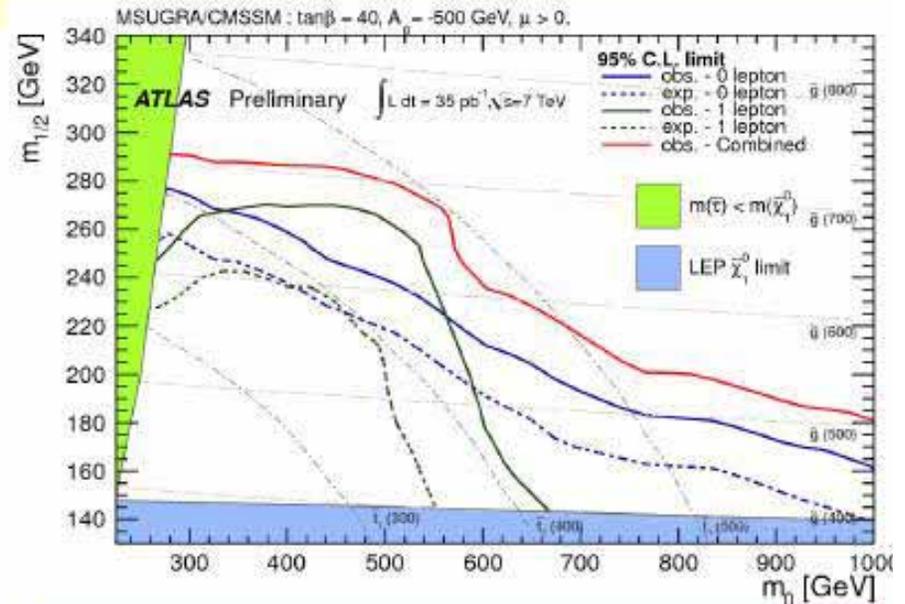
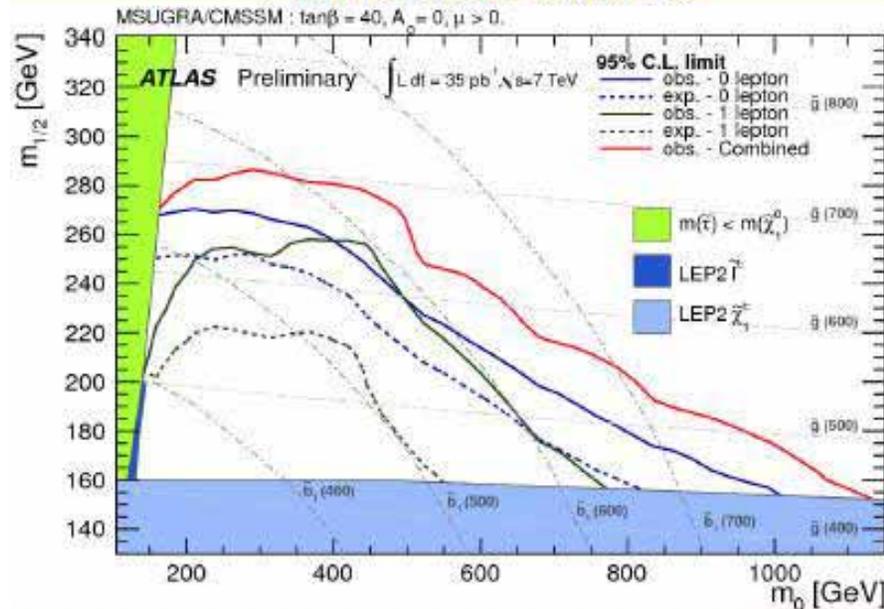
CMS preliminary

$L_{\text{int}} = 35 \text{ pb}^{-1}, \sqrt{s} = 7 \text{ TeV}$

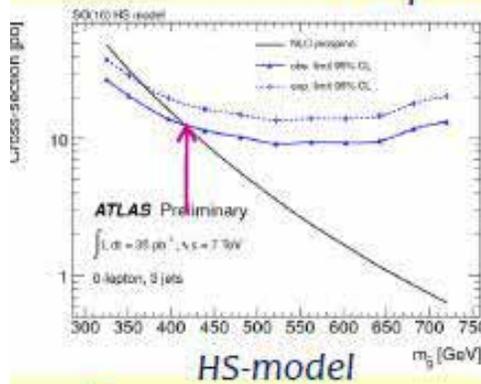


# Specific SUSY models

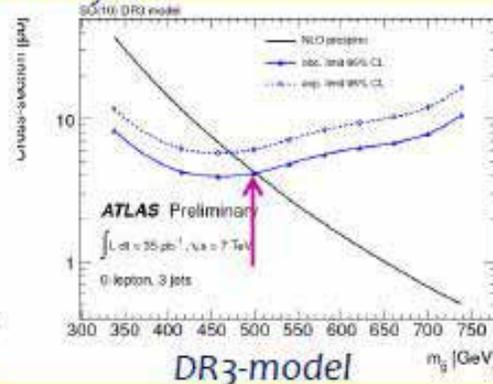
0- and 1-lepton analyses



0-lepton analysis

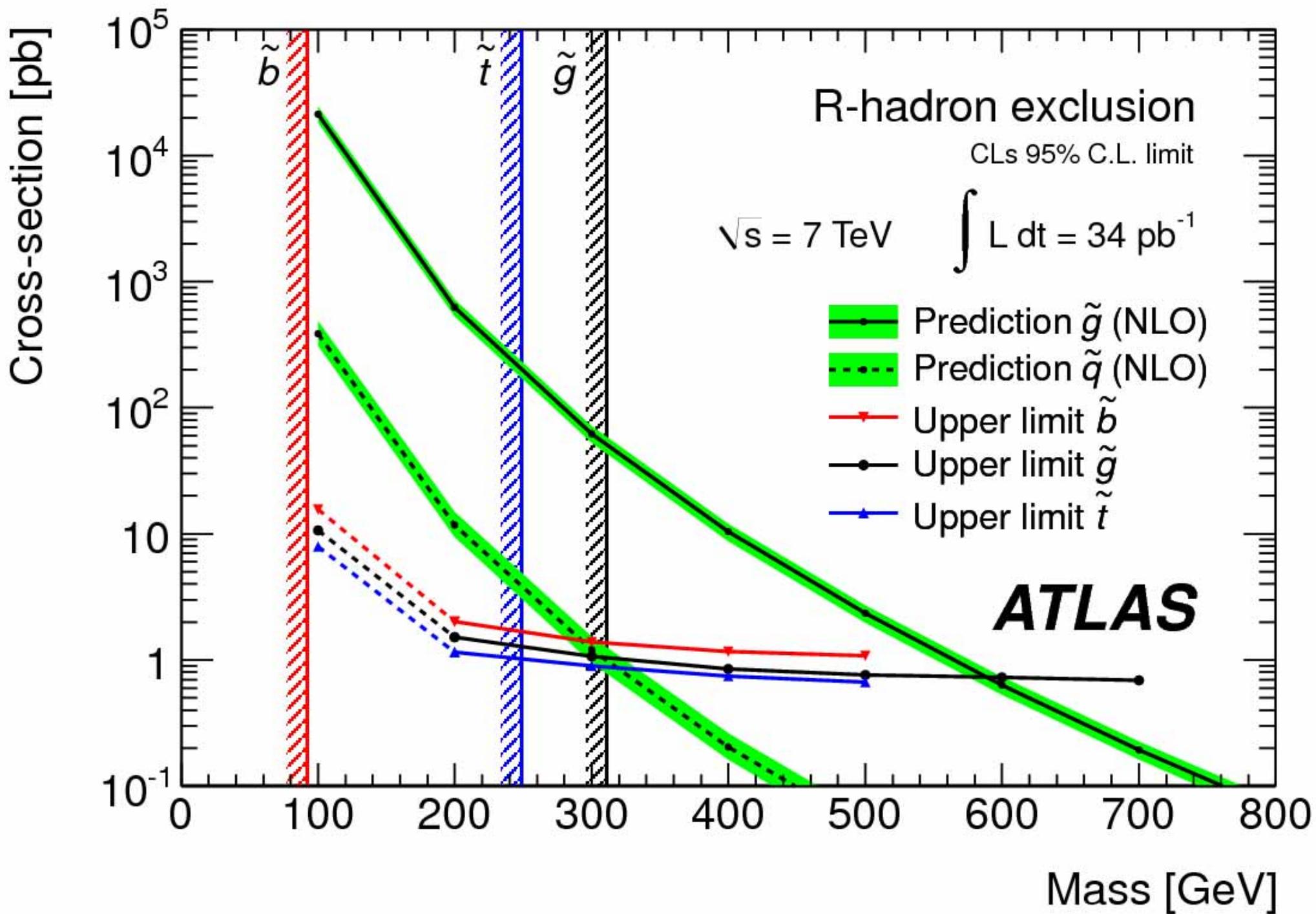


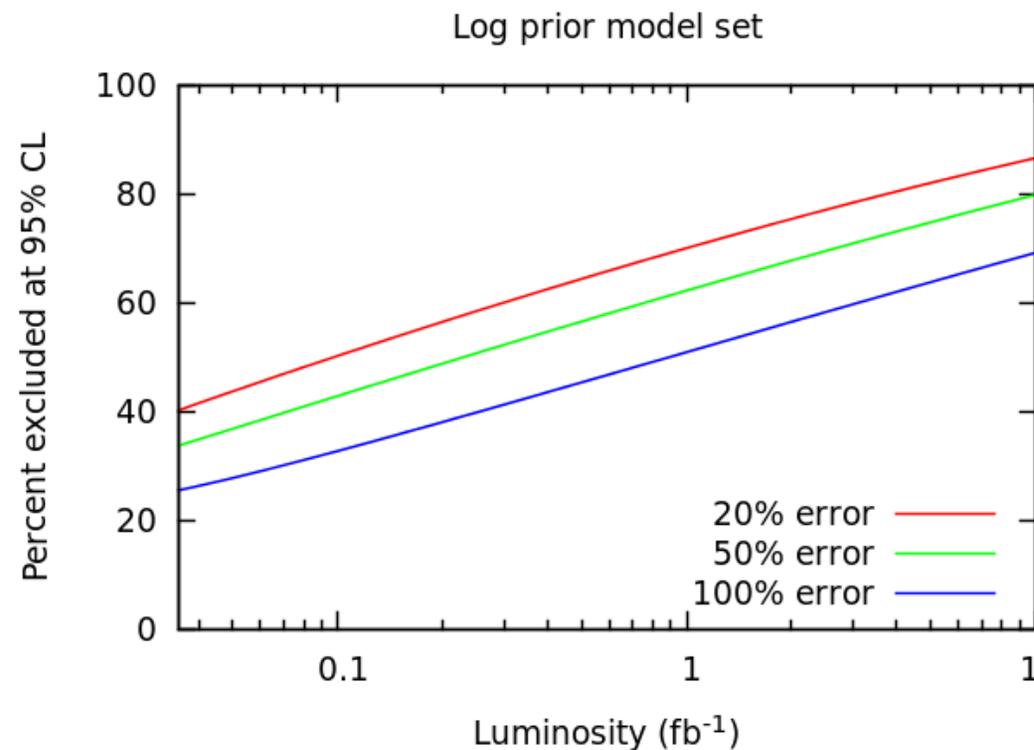
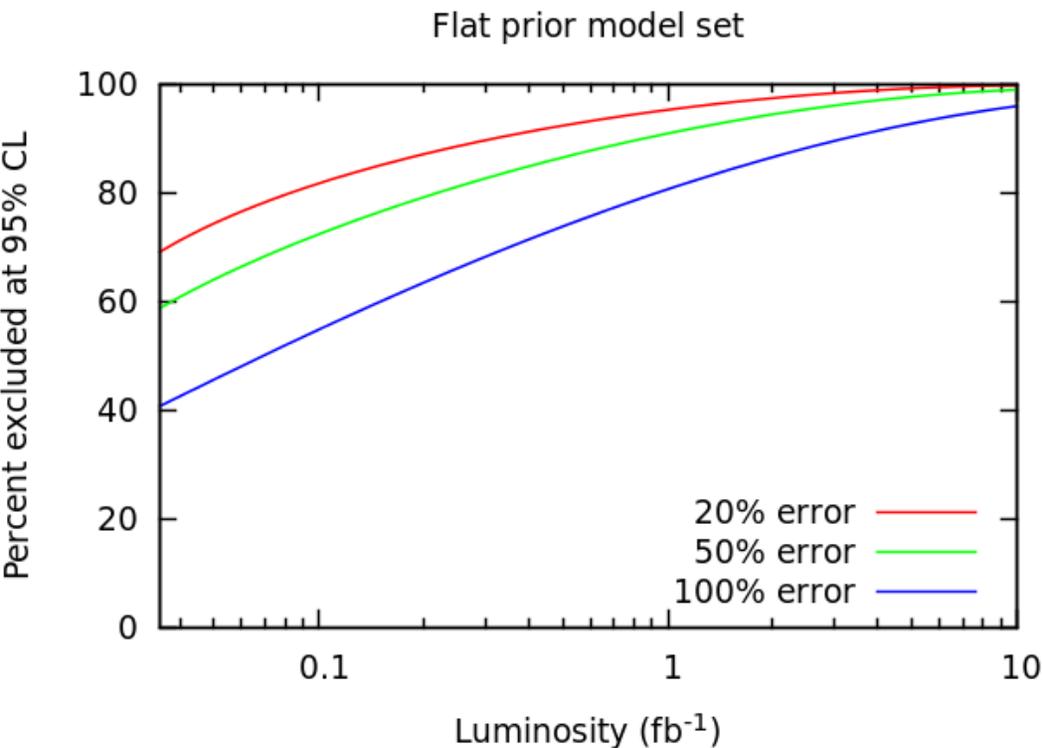
HS-model



DR3-model

- **mSUGRA:** large  $\tan \beta$  or low  $A_0$  values:
  - For each  $(m_0, m_{1/2})$  sbottom/stop masses lower than in low  $\tan \beta$  scenarios
  - **Exclude gluino masses up to 500 GeV for  $m_0 < 1$  TeV**
- **SO(10) models:** gluino pair production one of the dominant processes:
  - Gluino  $\rightarrow b\bar{b}\chi_1^0$  (DR3) or  $b\bar{b}\chi_2^0$  (HS)
  - **Exclude masses up to 500(420) GeV**

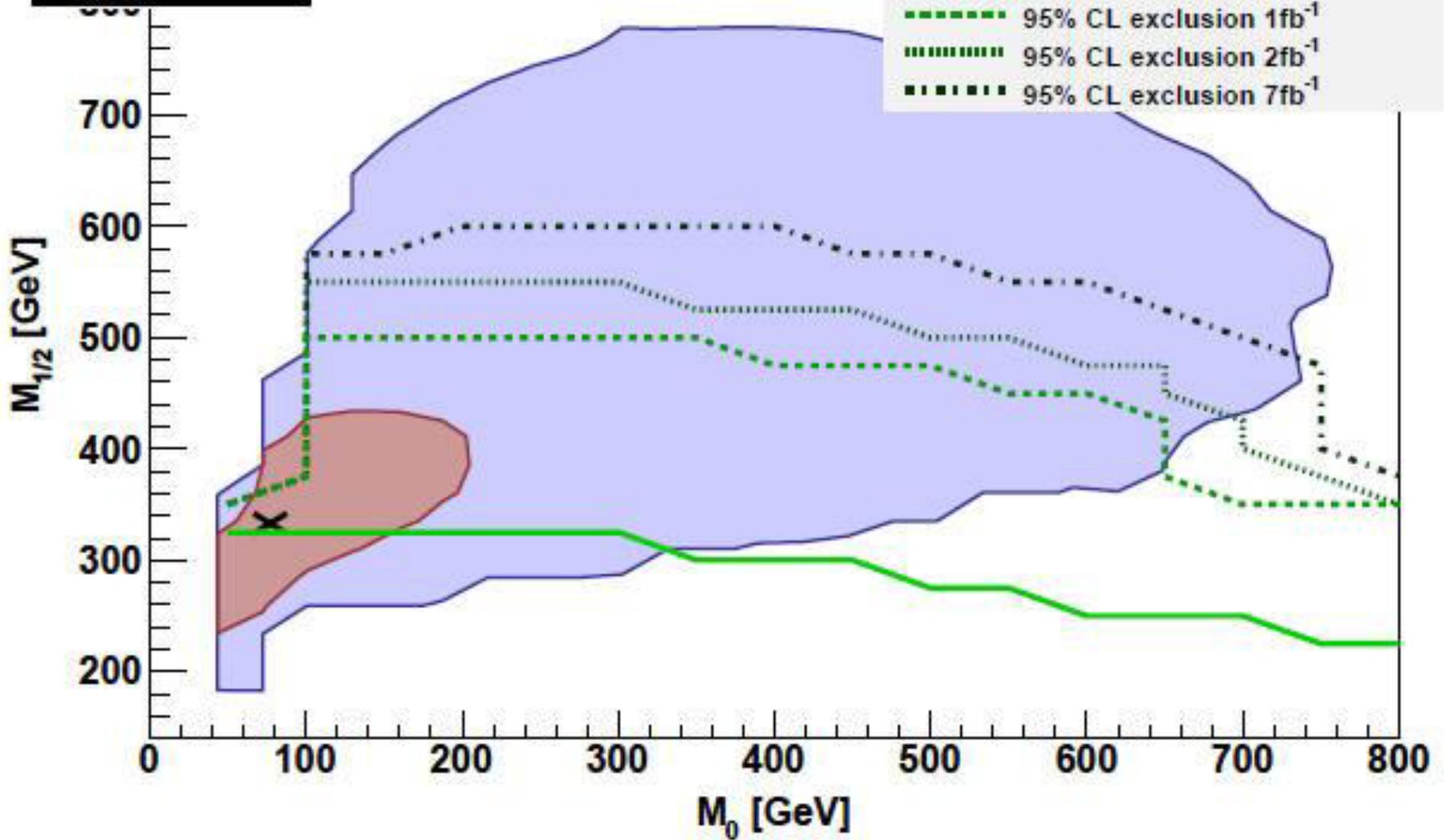




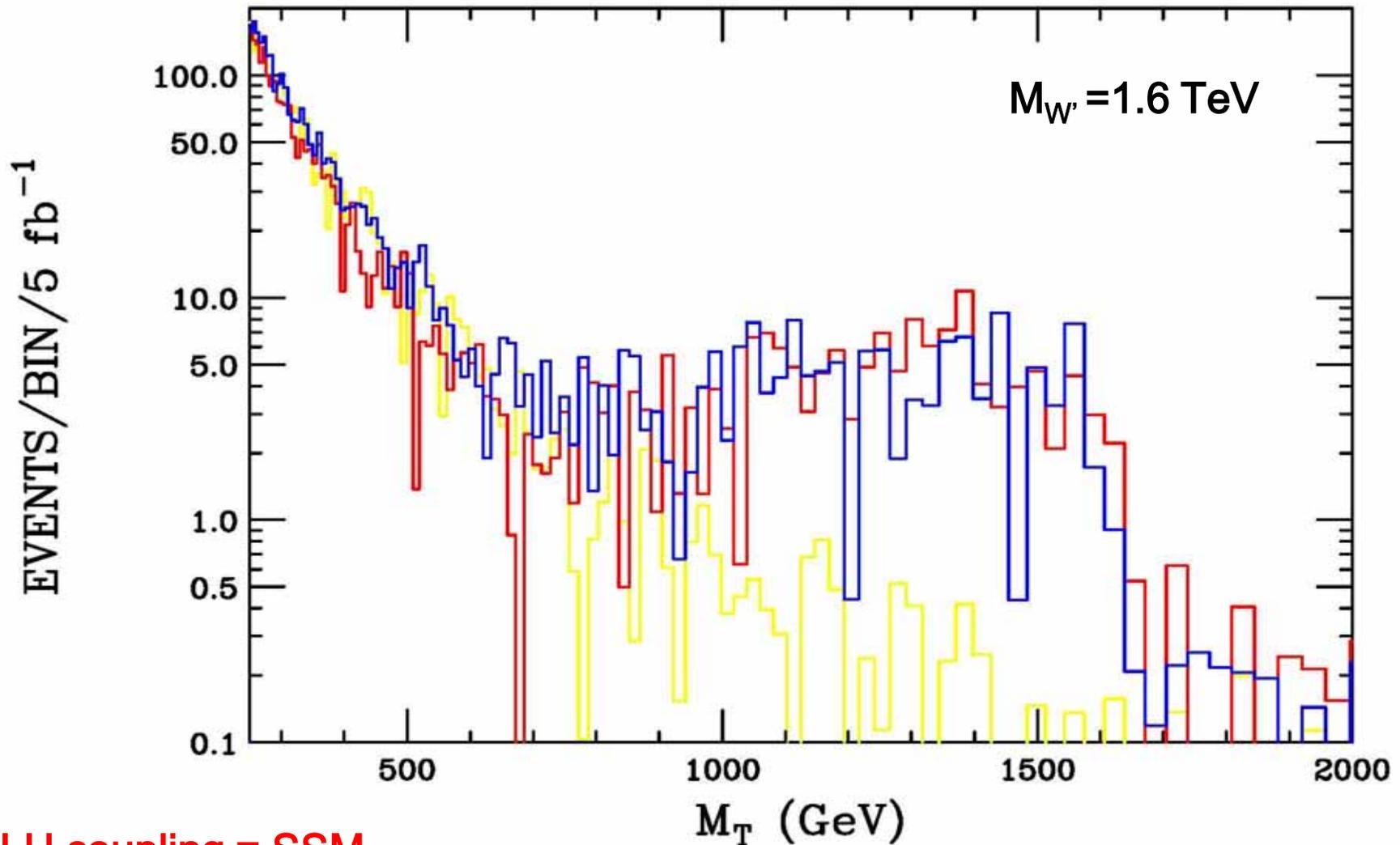
If we ask for the **current 95% CL** would-be **EXCLUSION** limits (assuming all these analyses were performed by **ATLAS as stated..which they were not**) then these model sets are even more constrained



Before LHC



Note:  $W'$  production has some dependency on coupling helicity (& **many** other **non-vanilla** assumptions !)



LH coupling = SSM

RH coupling = LRM

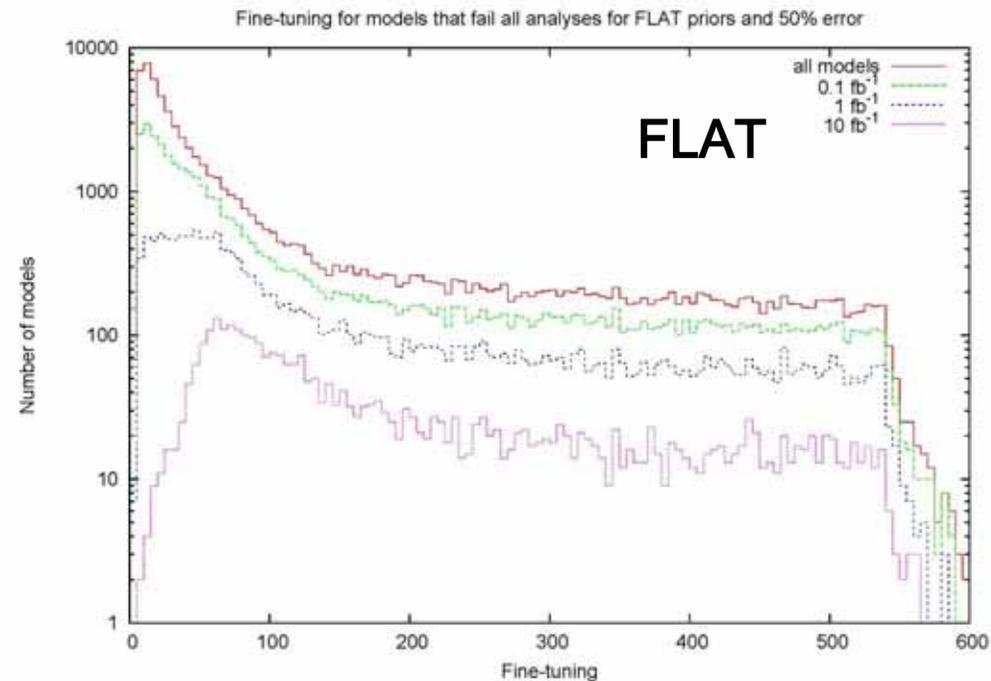
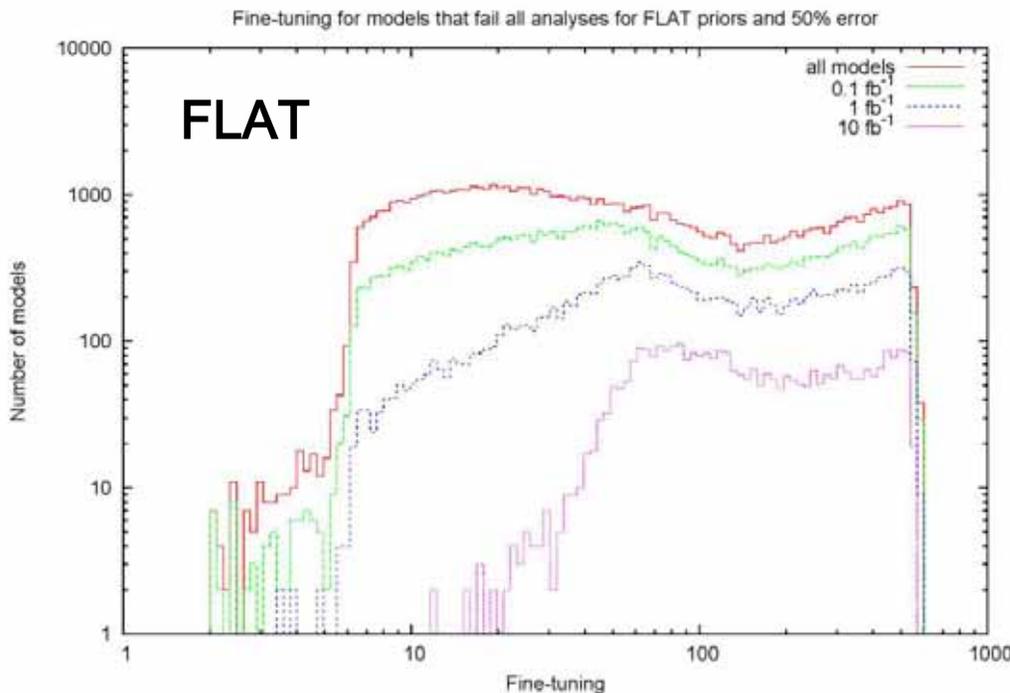
SM only

RH couplings do not interfere w/ the SM  $W$  !

LH couplings destructively interfere w/ SM

# Fine-Tuning SUSY ?

- It is often claimed that if the LHC (@7 TeV) does **not** find anything then SUSY must be **VERY** fine-tuned & so 'less likely'.  
**Is this true for the pMSSM??**



→ It is certainly true that models w/ low tuning do appear to 'suffer' more than those w/ larger values from null SUSY searches

- The amount of fine tuning in the LOG prior set is somewhat less influenced by null ATLAS searches due to spectrum differences

